## VIII. A Dynamical Theory of the Electromagnetic Field Maxwell 1864

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In order to bring these results within the power of symbolical calculation, I then express them in the form of the General Equations of the Electromagnetic Field. These equations express—

- (A) The relation between electric displacement, true conduction, and the total current, compounded of both.
- (B) The relation between the lines of magnetic force and the inductive coefficients of a circuit, as already deduced from the laws of induction.
- (C) The relation between the strength of a current and its magnetic effects, according to the electromagnetic system of measurement.
- (D) The value of the electromotive force in a body, as arising from the motion of the body in the field, the alteration of the field itself, and the variation of electric potential from one part of the field to another.
- (E) The relation between electric displacement, and the electromotive force which produces it.
- (F) The relation between an electric current, and the electromotive force which produces it.
- (G) The relation between the amount of free electricity at any point, and the electric displacements in the neighbourhood.
- (H) The relation between the increase or diminution of free electricity and the electric currents in the neighbourhood.

There are twenty of these equations in all, involving twenty variable quantities.

(70) In these equations of the electromagnetic field we have assumed twenty variable 3 v 2

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quantities,	namely,			*												
For	Electromagne	tic Mon	nentur	n										F	G	
	Magnetic Int	ensity												α	β	
22	Electromotive	e Force												Ρ	Q	
22	Current due t	to true o	onduc	etion	ι.									p	q	
55	Electric Displ	lacemen	t										. ,	f	q	
,,	Total Current	(includ	ing va	iriat	ion	of	dis	pla	cer	nen	t)			p'	$q^{\prime}$	
22	Quantity of f	ree Elec	tricity											e	-	
33	Electric Poter	ntial .												Ψ		
Between	these twenty of	quantitie	es we	have	e fo	une	d ty	ver	itv	eat	at	ions	. v	17.		
Т	hree equations	s of Mag	metic	For	ce					- 1			., .		(B)	
*		Elec	etric C	urre	ents								÷		(C)	
	1 33	Elec	etromo	otive	Fo	rce									(D)	
	32	Elec	etric E	last	icit	y									(E)	
	22	Elec	tric R	esis	tan	ce				•					(F)	
	33	Tota	l Cur	rent	s										(A)	
0	ne equation of	Free E	lectric	city											(G)	
	22	Contin	uity											•.	(H)	

These equations are therefore sufficient to determine all the quantities which occur in them, provided we know the conditions of the problem. In many questions, however, only a few of the equations are required.

#### **Maxwell's Equations**

#### The Original Equations

With the knowledge of fluid mechanics MAXWELL<sup>[15]</sup> has introduced the following eight equations to the electromagnetic fields (the right equations correspond with the original text, the left equations correspond with today's vector notation):

$$\frac{d\gamma}{dy} - \frac{d\beta}{dz} = 4\pi p' 
\frac{d\alpha}{dz} - \frac{d\gamma}{dx} = 4\pi q' 
\frac{d\beta}{dx} - \frac{d\alpha}{dy} = 4\pi r' 
\frac{d\beta}{dx} - \frac{d\beta}{dy} = 4\pi r'$$
(1.3)

$$P = \mu \left( \gamma \frac{dy}{dt} - \beta \frac{dz}{dt} \right) - \frac{dF}{dt} - \frac{d\Psi}{dx}$$

$$E_{1} = \mu \left( H_{3}v_{2} - H_{2}v_{3} \right) - \frac{dA_{1}}{dt} - \frac{d\varphi}{dx}$$

$$Q = \mu \left( \alpha \frac{dz}{dt} - \gamma \frac{dx}{dt} \right) - \frac{dG}{dt} - \frac{d\Psi}{dy}$$

$$\Rightarrow E_{2} = \mu \left( H_{1}v_{3} - H_{3}v_{1} \right) - \frac{dA_{2}}{dt} - \frac{d\varphi}{dy}$$

$$E_{3} = \mu \left( H_{2}v_{1} - H_{1}v_{2} \right) - \frac{dA_{3}}{dt} - \frac{d\varphi}{dz}$$

$$\Rightarrow E_{4} = \mu \left( H_{2}v_{1} - H_{1}v_{2} \right) - \frac{dA_{3}}{dt} - \frac{d\varphi}{dz}$$

$$(1.4)$$

$$\Rightarrow E_{4} = \mu \left( \mathbf{v} \times \mathbf{H} \right) - \frac{\partial \mathbf{A}}{\partial t} - \nabla \varphi$$

$$\begin{array}{c} P = -\zeta p \\ Q = -\zeta q \\ R = -\zeta r \end{array} \right\} \qquad \begin{array}{c} \sigma E_1 = j_1 \\ \rightarrow & \sigma E_2 = j_2 \\ \sigma E_3 = j_3 \end{array} \right\} \Rightarrow \sigma E = j \qquad (1.6)$$

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$$\mathbf{e} + \frac{\mathrm{d}\,f}{\mathrm{d}x} + \frac{\mathrm{d}\,g}{\mathrm{d}y} + \frac{\mathrm{d}\,h}{\mathrm{d}z} = 0 \quad \Rightarrow \quad \rho + \frac{\partial \mathbf{D}_1}{\partial x} + \frac{\partial \mathbf{D}_2}{\partial y} + \frac{\partial \mathbf{D}_3}{\partial z} = 0 \quad \Rightarrow \quad -\rho = \nabla \cdot \mathbf{D} \tag{1.7}$$

$$\frac{\mathrm{d}e}{\mathrm{d}t} + \frac{\mathrm{d}p}{\mathrm{d}x} + \frac{\mathrm{d}q}{\mathrm{d}y} + \frac{\mathrm{d}r}{\mathrm{d}z} = 0 \quad \Rightarrow \quad \frac{\partial\rho}{\partial t} + \frac{\partial j_1}{\partial x} + \frac{\partial j_2}{\partial y} + \frac{\partial j_3}{\partial z} = 0 \quad \Rightarrow \quad -\frac{\partial\rho}{\partial t} = \nabla \cdot \mathbf{j}$$
(1.8)

This original equations do not strictly correspond to today's vector equations. The original equations, for example, contains the vector potential **A**, which today usually is eliminated.

Three Maxwell equations can be found quickly in the original set, together with OHM's law (1.6), the FARADAY-force (1.4) and the continuity equation (1.8) for a region containing charges.

#### The Original Quaternion Form of Maxwell's Equations

In his Treatise<sup>[16]</sup> of 1873 MAXWELL has already modified his original equations of 1865. In addition Maxwell tried to introduce the quaternion notation by writing down his results also in a quaternion form. However, he has never really calculated with quaternions but only uses either the scalar or the vector part of a quaternion in his equations.

A general quaternion has a scalar (real) and a vector (imaginary) part. In the example below ,a' is the scalar part and 'ib + jc + kd' is the vector part.

$$\mathbb{Q} = \mathbf{a} + i\mathbf{b} + j\mathbf{c} + k\mathbf{d}$$

Here a, b, c and d are real numbers and i, j, k are the so-called HAMILTON 'ian<sup>[7]</sup> unit vectors with the magnitude of  $\sqrt{-1}$ . They fulfill the equations

$$i^2 = j^2 = k^2 = ijk = -1$$

and

$$ij = k$$
  $jk = i$   $ki = j$   
 $ij = -ji$   $jk = -kj$   $ki = -ik$ 

A nice presentation about the rotation capabilities of the HAMILTON'ian unit vectors in a three-dimensional ARGAND diagram was published by GOUGH<sup>[6]</sup>.

Now MAXWELL has defined the field vectors (for example  $\mathbf{B} = B_1 i + B_2 j + B_3 k$ ) as quaternions without scalar part and scalars as quaternions without vector part. In addition he defined a quaternion operator without scalar part

$$\nabla = \frac{\mathrm{d}}{\mathrm{d}x_1}i + \frac{\mathrm{d}}{\mathrm{d}x_2}j + \frac{\mathrm{d}}{\mathrm{d}x_3}k$$

which he used in his equations. Maxwell devided a single quaternion with two prefixes into a scalar and vector. This prefixes he defined according to

$$S.\mathbb{Q} = S.(a + ib + jc + kd) = a$$

$$V.\mathbb{Q} = V.(a + ib + jc + kd) = ib + jc + kd$$

The original Maxwell quaternion equations are now for isotrope media (no changes except fonts, normal letter = scalar, capital letter = quaternion without scalar):

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FALSE IN GENERAL (true only for statics)	TRUE ALWAYS					
$F = \frac{1}{4\pi\epsilon_0} \frac{q_1 q_2}{r^2} $ (Coulomb's law)	$F = q(E + v \times B) $ (Lorentz force) $\rightarrow \nabla \cdot E = \frac{\rho}{\epsilon_0} $ (Gauss' law)					
$\nabla \times E = 0$	$\rightarrow \nabla \times E = -\frac{\partial B}{\partial t} $ (Faraday's law)					
$E = -\nabla \phi$	$E = -\nabla\phi - \frac{\partial A}{\partial t}$					
$E(1) = \frac{1}{4\pi\epsilon_0} \frac{\rho(2)e_{12}}{r_{12}^2} dV_2$						
For conductors, $E = 0$ , $\phi = \text{constant}$ . $Q = CV$	In a conductor, <i>E</i> makes currents.					
	$  \overrightarrow{\nabla} \cdot B = 0 $ (No magnetic charges) $ B = \nabla \times A $					
$c^2 \nabla \times \boldsymbol{B} = \frac{\boldsymbol{j}}{\epsilon_0}$ (Ampere's law)	$\rightarrow c^2 \nabla \times B = \frac{j}{\epsilon_0} + \frac{\partial E}{\partial t}$					
$B(1) = \frac{1}{4\pi\epsilon_0 c^2} \int \frac{j(2) \times e_{12}}{r_{12}^2}  dV_2$						
$\nabla^2 \phi = -\frac{\rho}{\epsilon_0} $ (Poisson's equation)	$ \left( \begin{array}{c} \nabla^2 \phi \ - \frac{1}{c^2} \frac{\partial^2 \phi}{\partial t^2} = - \frac{\rho}{\epsilon_0} \end{array} \right) $					
$\nabla^2 A = -\frac{j}{\epsilon_0 c^2}$	with $\nabla^2 A - \frac{1}{c^2} \frac{\partial^2 A}{\partial t^2} = -\frac{j}{\epsilon_0 c^2}$					
$\nabla \cdot \boldsymbol{A} = \boldsymbol{0}$	$\begin{bmatrix} with \\ c^2 \nabla \cdot A + \frac{\partial \phi}{\partial t} = 0 \end{bmatrix}$					
$\phi(1) = \frac{1}{4\pi\epsilon_0} \int \frac{\rho(2)}{r_{12}} dV_2$	$\phi(1,t) = \frac{1}{4\pi\epsilon_0} \int \frac{\rho(2,t')}{r_{12}} dV_2$					
$A(1) = \frac{1}{4\pi\epsilon_0 c^2} \int \frac{j(2)}{r_{12}}  dV_2$	and $A(1, t) = \frac{1}{4\pi\epsilon_0 c^2} \int \frac{j(2, t')}{r_{12}} dV_2$ with $t' = t - \frac{r_{12}}{c}$					
$U = \frac{1}{2} \int \rho \phi  dV + \frac{1}{2} \int \mathbf{j} \cdot \mathbf{A}  dV$	$U = \int \left(\frac{\epsilon_0}{2} \mathbf{E} \cdot \mathbf{E} + \frac{\epsilon_0 c^2}{2} \mathbf{B} \cdot \mathbf{B}\right) dV$					

The equations marked by an arrow  $(\rightarrow)$  are Maxwell's equations.

Table 15–1

#### 15–15

#### **Maxwell's Equations**

In general :

$$\begin{cases} \mathbf{\nabla} \cdot \mathbf{E} = \frac{1}{\epsilon_0} \rho \\ \mathbf{\nabla} \times \mathbf{E} = -\frac{\partial \mathbf{B}}{\partial t} \\ \mathbf{\nabla} \cdot \mathbf{B} = 0 \\ \mathbf{\nabla} \times \mathbf{B} = \mu_0 \mathbf{J} + \mu_0 \epsilon_0 \frac{\partial \mathbf{E}}{\partial t} \end{cases}$$

In matter :

$$\nabla \cdot \mathbf{D} = \rho_f$$
$$\nabla \times \mathbf{E} = -\frac{\partial \mathbf{B}}{\partial t}$$
$$\nabla \cdot \mathbf{B} = 0$$
$$\nabla \times \mathbf{H} = \mathbf{J}_f + \frac{\partial \mathbf{D}}{\partial t}$$

#### **Auxiliary Fields**

**Definitions** :

$$\mathbf{D} = \epsilon_0 \mathbf{E} + \mathbf{P}$$
$$\mathbf{H} = \frac{1}{\mu_0} \mathbf{B} - \mathbf{M}$$

Linear media :

$$\mathbf{P} = \epsilon_0 \chi_e \mathbf{E}, \quad \mathbf{D} = \epsilon \mathbf{E}$$
  
 $\mathbf{M} = \chi_m \mathbf{H}, \quad \mathbf{H} = \frac{1}{\mu} \mathbf{E}$ 

Feynman: These Eight Equations Contain All of Classical Physics

(1-4) Maxwell's Equations

$$\nabla \cdot \mathbf{E} = \rho/\epsilon_0$$
$$\nabla \cdot \mathbf{B} = 0$$
$$\nabla \times \mathbf{E} = -\partial \mathbf{B}/\partial t$$
$$\nabla \times \mathbf{B} = \mu_0 \mathbf{J} + \mu_0 \epsilon_0 \ \partial \mathbf{E}/\partial t$$

(5) The Conservation of Charge

$$\nabla \cdot \mathbf{J} = -\partial \rho / \partial t$$

(6) The Lorentz Force Law

$$\mathbf{F} = q\mathbf{E} + q\mathbf{v} \times \mathbf{B}$$

- (7) The Law of Motion
  - $d\mathbf{p}/dt = \mathbf{F}$  where  $\mathbf{p} = \gamma m \mathbf{v}$
- (8) Newton's Law of Gravitation

$$\mathbf{F} = -Gm_1m_2 \; \hat{\mathbf{r}}/r^2$$

#### Maxwell's Equations

Two vector fields,  ${\bf E}$  and  ${\bf B}$ 

#### **Two Divergence Equations**

Gauss' Law for Electricity

 $\nabla \cdot \mathbf{E} = \rho / \epsilon_0$ 

Gauss' Law for Magnetism

$$\nabla \cdot \mathbf{B} = 0$$

**Two Curl Equations** 

Faraday's Law of Induction

 $\nabla \times \mathbf{E} = -\partial \mathbf{B} / \partial t$ 

#### Ampere's Law with Maxwell's Extension

 $\nabla \times \mathbf{B} = \mu_0 \mathbf{J} + \mu_0 \epsilon_0 \partial \mathbf{E} / \partial t$ 

$$\nabla \times \mathbf{B} = \mu_0 \mathbf{J} + \mu_0 \mathbf{J}_{\mathbf{d}}$$

Maxwell called  $\mathbf{J}_{\mathbf{d}}$  the displacement current

For static fields, the four Maxwell Equations are the best formulation

$$\nabla \cdot \mathbf{E} = \rho/\epsilon_0$$
$$\nabla \cdot \mathbf{B} = 0$$
$$\nabla \times \mathbf{E} = -\partial \mathbf{B}/\partial t$$
$$\nabla \times \mathbf{B} = \mu_0 \mathbf{J} + \mu_0 \epsilon_0 \partial \mathbf{E}/\partial t$$

Maxwell's equations are four coupled first-order differential equations The sources are  $\rho$ , J, and each other

For dynamic fields, the wave equations are the best formulation Decoupling the four Maxwell Equations produces two second-order differential equations

$$\nabla^{2}\mathbf{E} = \mu_{0}\epsilon_{0} \ \partial^{2}\mathbf{E}/\partial^{2}t = (1/c^{2}) \ \partial^{2}\mathbf{E}/\partial^{2}t$$
$$\nabla^{2}\mathbf{B} = \mu_{0}\epsilon_{0} \ \partial^{2}\mathbf{B}/\partial^{2}t = (1/c^{2}) \ \partial^{2}\mathbf{B}/\partial^{2}t$$

These two wave equations have six degrees-of-freedom, three for E and three for B. They apply in regions where there are no sources.

The vector potential and the scalar potential obey the same wave equation

$$\begin{aligned} \nabla^2 \mathbf{A} &= \partial^2 \mathbf{A} / \partial^2 t \\ \nabla^2 \phi &= \partial^2 \phi / \partial^2 t \end{aligned}$$

These two wave equations have four degrees-of-freedom, one for  $\phi$  and three for A

As we will see later, the vector potential and the scalar potential are the components of the four-potential  $A_{\mu}$ 

$$A_{\mu} = (\phi, A_x, A_y, A_z)$$

In the radiation gauge,

 $\partial \phi / \partial t + \nabla \cdot \mathbf{A} = 0$ 

the four-potential  $A_{\mu}$  obeys the same wave equation

$$\begin{bmatrix} \nabla^2 - \partial^2 / \partial^2 t \end{bmatrix} \phi = 0$$
$$\begin{bmatrix} \nabla^2 - \partial^2 / \partial^2 t \end{bmatrix} Ax = 0$$
$$\begin{bmatrix} \nabla^2 - \partial^2 / \partial^2 t \end{bmatrix} Ay = 0$$
$$\begin{bmatrix} \nabla^2 - \partial^2 / \partial^2 t \end{bmatrix} Ay = 0$$

Using Feynman's four-vector notation, Maxwell's equations, in the radiation gauge, are given by

$$abla_{\mu} 
abla_{\mu} A_{\mu} = j_{\mu} / \epsilon_0$$
 $abla_{\mu} j_{\mu} = 0$ 

Note that  $\nabla_{\mu}\nabla_{\mu}$  becomes  $[\partial^2/\partial^2 t, -\nabla^2]$  but  $\nabla_{\mu}$  becomes  $[\partial/\partial t, \nabla]$ 

Using modern notation, Maxwell's equations, in the radiation gauge, are given by

$$\partial_{\mu}\partial^{\mu} A^{\nu} = 4\pi j^{\nu}$$
  
 $\partial_{\nu}j^{\nu} = 0$ 

## TABLE 3.1 ANALOGY BETWEEN THE ELECTROMAGNETIC AND GRAVITATIONAL FIELD THEORIES

Electromagnetism

Gravitation
. (Linear Approximation)

 $d\tau^2 = (\eta_{\alpha\beta} + \kappa h_{\alpha\beta}) dx^{\alpha} dx^{\beta}$ 

Source of field Conservation law Field Field equation

Gauge transformation Preferred gauge condition Field equation in preferred gauge Energy-momentum exchange between field and particle Equation of motion of particle Energy-momentum of particle Proper time interval

j <sup>v</sup>	$T^{\mu u}$
$\partial_{\nu}j^{\nu}=0$	$\partial_{\nu}T^{\mu\nu}=0$
$A^{\nu}$	h <sup>µv</sup>
$\partial_{\mu}\partial^{\mu}A^{\nu}-\partial^{\nu}\partial_{\mu}A^{\mu}=4\pi j^{\nu}$	$\partial_{\lambda}\partial^{\lambda}h^{\mu\nu} - 2\partial_{\lambda}\partial^{(\nu}h^{\mu)\lambda} + \partial^{\mu}\partial^{\nu}h \\ -\eta^{\mu\nu}\partial_{\lambda}\partial^{\lambda}h + \eta^{\mu\nu}\partial_{\lambda}\partial_{\sigma}h^{\lambda\sigma} = -\kappa T^{\mu\nu}$
$A^{\mu} \rightarrow A^{\mu} + \partial^{\mu} \Lambda$	$h^{\mu u}  ightarrow h^{\mu u} + \partial^{( u} \Lambda^{\mu)}$
$\partial_{\mu}A^{\mu}=0$	$\partial_{\mu}(h^{\mu\nu}-rac{1}{2}\eta^{\mu\nu}h)=0$
$\partial_{\mu}\partial^{\mu}A^{\nu}=4\pi j^{\nu}$	$\partial_{\lambda}\partial^{\lambda}(h^{\mu\nu}-\tfrac{1}{2}\eta^{\mu\nu}h)=-\kappa T^{\mu\nu}$
$\partial_{\nu} T_{(m)\mu}{}^{\nu} = F_{\mu\nu} j^{\nu}$	$\partial_{\nu} T_{(m)\mu}^{\nu} = \frac{\kappa}{2} m h_{\alpha\beta,\mu} T_{(m)}^{\alpha\beta}$
$\frac{d}{d\tau}p_{\mu}=qF_{\mu\nu}u^{\nu}$	$\frac{d}{d\tau} P_{\mu} = \frac{\kappa}{2} m h_{\alpha\beta,\mu} u^{\alpha} u^{\beta}$
$p_{\mu} = m \mu_{\mu}$	$P_{\mu} = m u_{\mu} + m \kappa h_{\mu \alpha} u^{\alpha}$

 $p_{\mu} = m u_{\mu}$  $d\tau^2 = \eta_{\alpha\beta} dx^{\alpha} dx^{\beta}$ 

The Principle of Relativity

All laws of physics must be invariant under Lorentz transformations.

Invariant means:

(1) The law has the same mathematical form

(2) All numerical constants have the same values

A potpouri of four-vectors and four-operators

spacetime 4-vector s = (t, x, y, z)

energy-momentum 4-vector  $(E, p_x, p_y, p_z)$ 

4-potential  $A_{\mu} = (\phi, A_x, A_y, A_z)$ 

4-current  $j_{\mu} = (\rho, j_x, j_y, j_z)$ 

Feynman's  $\nabla$  in 4-dimensions  $\nabla_{\mu} = (\partial/\partial t, \nabla)$ 

Feynman's D'Alembertian =  $\nabla_{\mu}\nabla_{\mu} = (\partial^2/\partial t^2, -\nabla^2)$  4-Laplacian

In modern notation we use  $\partial_{\mu}$  and  $\partial_{\mu}\partial^{\mu}$ 

All 4-vectors transform in precisely the same way.

The Structure of Spacetime Virtual Text 2 starts by saying "Lorentz transformations are just hyperbolic rotations." I am going to go one step further by asserting

Special Relativity is hyperbolic geometry.

Newtonian Space And Time Space and time are different separate entities. Distances are calculated using the Pythagorean Theorem  $s^2 = x^2 + y^2 + z^2$ 

Euclidean Spacetime, Cartesian Spacetime, Galilean Spacetime For 90 proofs of the Pythagorean Theorem see: http://www.cut-the-knot.org/pythagoras/index.shtml

Spacetime in SR (aka Minkowski Spacetime) Space and time are one entity. Distances are calculated differently  $s^2 = t^2 - x^2 - y^2 - z^2$ In two dimensions

$$s^2 = t^2 - x^2$$

This is how distance is measured in two-dimensional hyperbolic geometry.

#### **Calculating Distances**

In Euclidean Spacetime, distances are calculated using the inner product

$$\mathbf{r} = \begin{bmatrix} x \\ y \end{bmatrix}$$
$$\mathbf{r} \cdot \mathbf{r} = |\mathbf{r}|^2 = x^2 + y^2$$

In Minkowski Spacetime, distances are calculated using a different inner product

$$\mathbf{s} = \begin{bmatrix} t \\ x \end{bmatrix}$$
$$\mathbf{s} \cdot \mathbf{s} = |\mathbf{s}|^2 = t^2 - x^2$$

Formally, we can define the inner product using a matrix

For Euclidean Spacetime

$$\mathbf{M} = \begin{bmatrix} 1 & 0\\ 0 & 1 \end{bmatrix}$$

For Minkowski Spacetime

$$\mathbf{M} = \begin{bmatrix} 1 & 0\\ 0 & -1 \end{bmatrix}$$

Distance is given by the inner product  $\mathbf{s} \cdot \mathbf{s} = \mathbf{s}^T \mathbf{M} \mathbf{s}$ The matrix M that defines distance is called the metric.

On the west coast

$$\mathbf{M} = \begin{bmatrix} 1 & 0 & 0 & 0 \\ 0 & -1 & 0 & 0 \\ 0 & 0 & -1 & 0 \\ 0 & 0 & 0 & -1 \end{bmatrix} \text{ and } s^2 = t^2 - x^2 - y^2 - z^2$$

On the east coast

$$\mathbf{M} = \begin{bmatrix} -1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{bmatrix} \text{ and } s^2 = -t^2 + x^2 + y^2 + z^2$$

The important thing about distances—also called intervals—is that they have the same value independent of the coordinate system used to measure them.

The Lorentz Transformations

Coordinate system changes in Minkowski Spacetime are given by the Lorentz transformations

The Lorentz transformations are the only linear coordinate transformations that produce invariant distances. The transformations must be linear so that the worldlines of free particles are straight lines in all coordinate systems.

$$t' = At + Bx$$
$$x' = Ct + Dx$$

The Lorentz transformations are given by

$$t' = \frac{t - vx/c}{\sqrt{1 - (v^2/c^2)}}$$
$$x' = \frac{x - vt}{\sqrt{1 - (v^2/c^2)}}$$

Written more compactly

$$t' = \gamma t - \beta \gamma x$$
$$x' = \gamma x - \beta \gamma t$$

Where

$$\beta = v/c$$
 
$$\gamma = \frac{1}{\sqrt{1-(v^2/c^2)}}$$

As the author of Virtual Text 2 asserted, we will see later that the Lorentz transformations are hyperbolic rotations.

 $\label{eq:http://www.univie.ac.at/future.media/moe/galerie/struct/struct.html \\ http://webphysics.davidson.edu/applets/Minkowski/Minkowski_FEL.html \\$ 

### Matrix form

This Lorentz transformation is called a "boost" in the x-direction and is often expressed in matrix form as

.

$$\begin{bmatrix} ct' \\ x' \\ y' \\ z' \end{bmatrix} = \begin{bmatrix} \gamma & -\beta\gamma & 0 & 0 \\ -\beta\gamma & \gamma & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{bmatrix} \begin{bmatrix} ct \\ x \\ y \\ z \end{bmatrix}$$

This transformation matrix is universal for all four-vectors.

More generally for a boost in any arbitrary direction  $(\beta_x, \beta_y, \beta_z)$ ,

$$\begin{bmatrix} ct'\\x'\\y'\\z' \end{bmatrix} = \begin{bmatrix} \gamma & -\beta_x \gamma & -\beta_y \gamma & -\beta_z \gamma \\ -\beta_x \gamma & 1 + (\gamma - 1)\frac{\beta_x^2}{\beta^2} & (\gamma - 1)\frac{\beta_x\beta_y}{\beta^2} & (\gamma - 1)\frac{\beta_x\beta_z}{\beta^2} \\ -\beta_y \gamma & (\gamma - 1)\frac{\beta_y\beta_x}{\beta^2} & 1 + (\gamma - 1)\frac{\beta_y^2}{\beta^2} & (\gamma - 1)\frac{\beta_y\beta_z}{\beta^2} \\ -\beta_z \gamma & (\gamma - 1)\frac{\beta_z\beta_x}{\beta^2} & (\gamma - 1)\frac{\beta_z\beta_y}{\beta^2} & 1 + (\gamma - 1)\frac{\beta_z^2}{\beta^2} \end{bmatrix} \begin{bmatrix} ct\\x\\y\\z \end{bmatrix}$$
where  $\beta = \frac{v}{c} = \frac{|\vec{v}|}{c}$  and  $\gamma = \frac{1}{\sqrt{1 - \beta^2}}$ .

# **Hyperbolic Geometry**

Three forms of geometry Euclid's Fifth Postulate

- 1 parallel line => Flat (zero curvature) Euclidean Geometry
- no parallel lines => positive curvature Spherical Geometry
- infinite number => negative curvature Hyperbolic Geometry

# The three geometries

	Elliptical Spherical Riemannian	Parabolic Planar Euclidean	Hyperbolic Saddle Lobachevskian
curvature	positive	0	negative
triangles	>180 <sup>0</sup>	=180 <sup>0</sup>	<180 <sup>0</sup>
parallel lines	zero	1	infinite
projections	Mercator		<b>Poincare disk</b>
	<b>Gall-Peters</b>		Poincare half plane
	Dymaxion		Klein disk
	Cuboctahedron		Minkowski

## For this spherical triangle the sum of the internal angles = 270 degrees







F1G. 234a

F1G. 234b

H

F1G. 234c







# How can we project the hyperbolic plane onto the Euclidian plane?

