- 1. Consider an arbitrary change of coordinates (or diffeomorphism) from a set $\{x^{\mu}\}$ to a new set $\{\tilde{x}^{\alpha}\}$ such that $x^{\mu} = f^{\mu}(\tilde{x})$ for some set of (differentiable) functions $\{f^{\mu}(\tilde{x})\}$. Under such a change of coordinates, a scalar function $\phi \to \tilde{\phi}$ where $\phi(\tilde{x}) \equiv \phi(x)$ evaluated at x = f(y). Using coordinate bases, the new and old basis vectors are related via $\tilde{\mathbf{e}}_{\alpha}(\tilde{x}) =$ $(\partial x^{\mu}/\partial \tilde{x}^{\alpha}) \mathbf{e}_{\mu}(x)$ (again evaluated at $x = f(\tilde{x})$). (i) How do the components of a vector field transform? (ii) How do the components of a one-form transform? (iii) How do the components of the metric transform? (iv) How does the determinant of the metric transform? (v) Explain why $(-g)^{1/2}$ is the spacetime volume element, where $g \equiv \det ||g_{\mu\nu}||$.
- 2. With $g \equiv \det ||g_{\alpha\beta}||$, prove that:
 - (i) $(\ln g)_{,\alpha} = g^{\beta\gamma}g_{\beta\gamma,\alpha}$ (ii) $\Gamma^{\alpha}{}_{\alpha\beta} = (\ln |g|^{1/2})_{,\beta}$ in a coordinate frame. (iii) $A^{\alpha}{}_{;\alpha} = |g|^{-1/2}(|g|^{1/2}A^{\alpha})_{,\alpha}$ in a coordinate frame. (iv) $\Box \phi \equiv \phi_{;\alpha}{}^{;\alpha} = |g|^{-1/2}(|g|^{1/2}g^{\alpha\beta}\phi_{,\alpha})_{,\beta}$ in a coordinate frame.
- 3. The metric of S^D , a *D*-dimensional sphere, may be written in the form

$$ds^2 = d\alpha_1^2 + \sin^2 \alpha_1 \, d\alpha_2^2 + \sin^2 \alpha_1 \, \sin^2 \alpha_2 \, d\alpha_3^2 + \dots + \left(\prod_{j=1}^{D-1} \sin^2 \alpha_j\right) \, d\alpha_D^2 \,,$$

where $\{\alpha_1, \dots, \alpha_D\}$ are the angular coordinates one introduces when defining spherical coordinates in \mathbb{R}^{D+1} . (i) Justify this assertion. What are the appropriate ranges for each angle? (ii) Calculate the connection coefficients and show that great circles are geodesics. Demonstrating this for great circles passing through the "north pole" is sufficient. (iii) Calculate the Riemann curvature tensor, the Ricci tensor, and the Ricci scalar. Show that $R_{ijkl} = K(g_{ik} g_{jl} - g_{il} g_{jk}), R_{jl} = (D-1) K g_{jl}$, and R = D(D-1) K for some constant K.

Hints: Calculating with pencil and paper is, of course, fine. So is using computer assistance. If you choose the later, I recommend becoming acquainted with, and using, the Riemannian Geometry & Tensor Calculus package with Mathematica (MMA). Links to these are on the class web page. When using this package with MMA, the spacetime dimension D must be a number and cannot be left free. Write a MMA function which will let you perform the above calculations for any given value of D, and verify the assertions above for all $D \leq 6$.