^{<10}Cosmic Homogeneity: Inflation

WHAT: Sudden growth of the cosmic horizon WHEN: 10⁻³⁵ s ABB WHERE: everywhere

Space inflated (multiplied) in scale by a factor of 10^{50} between I = 10^{-35} and 10^{-33} s ABB.

♦ the horizon (radius of spheres of influence) ballooned from $10^{-28.5}$ km to $10^{+23.5}$ km ≈100 x the radius or today's visible universe l.y.

Size of proton grew from 10⁻¹³ cm to 10³² km = 10¹⁹ l.y. !
Regions once in close touch suddenly lost touch and no longer influence each other's evolution (and never will)
However, these regions shard (still share) their "inherited genes" - physical properties, forces, & processes

Inflation

Horizou

Exponential Stretch

Rapid Expansion

Inflation resolves problems

WHAT: Expansion by 10⁵⁰ in 10⁻³³ sec WHEN: 10⁻³⁵ s to 10⁻³⁵ s ABB WHERE: everywhere

• Horizon Problem: Inflation insured that every part of the visible universe, even 180 degrees apart, were once in thermal contact, and would have the same temperature.

Horizou

<10-35s

Exponential Stretch

Rapid Expansion

Smoothness Problem: The number of particles in the initial unit that became our visible universe after inflation was much smaller. So Inflation prevents clumping from being larger than observed (rather than being too smooth).

Flatness Problem: The inflation decreased the curvature of space-time by a factor of 10⁵⁰. That made it flat to all practical purposes.

Inflation

Cosmic Homogeneity: Inflation



Who Needs Inflation?

Horizon Problem

Why does each part of the sky look much the same?

the number and size of density fluctuations on both sides of the sky are similar, yet they are separated by a distance that is greater than the speed of light times the age of the Universe, i.e. they should have no knowledge of each other by special relativity



Horizon Solution

at some time in the early Universe, all parts of spacetime were causally connected, this must have happened after the spactime foam era, and before the time where thermalization of matter occurred.

Who Needs Inflation?

Why is the Universe Flat? Inflation helps!

Flatness Problem

A в C

Lots of inflation decreases the curvature of the universe

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Who Needs Inflation?

Inflation set just the right ripple contrast in the CMB



Who Needs Inflation? Us! No galaxies then no stars, planets, or life

Anthropic Principle

theories of the universe are constrained by the necessity to allow human existence.



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Looking Ahead: Cosmic Outcomes







Long-term Cosmic Outcomes



Long-term Cosmic Outcomes What do you expect?



Surely you just made DE up.



Do you really believe this stuff?

Alternates to Dark Energy: Quintessence and Modified Gravity Chasing Prescience

Quintessence: pervasive cosmic repulsive force field with gentle form and very gradual change in structure (text page 256) imagine an almost stationary energy fluid roughly similar (and possibly related) to dark matter "quintessence" means the fifth essence of existence –after earth, sky, fire, and water

Modified gravity: GR's description of gravity needs modification so that the force of gravity loses its influence over cosmic distances



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