

Mutual inductance is the principle behind transformer: I co!

An Ac transformer may increase (stepsip)

of decroses voltage from primary to

Secondary. Ignition in car taker

input voltage after between and riv

and stepsit up to high out put voltage so that

E at spark plug > break down

field for the air fuel in (stane in cylude,

spark is produced, my twe igniter

Transformers - Mistary of AC, electric power Librat trans mysioning) ase than houseld

Developmentof transfor mers was an important step in the historial the development Long dist transmission into we much higher voltage than household appliances of electric lights. Transformers can be socnon to lephone poles

The cakulation of E, is an example
ofseltenductance: If there is a currenting
loop There will a B and there forea
flux. The flux is propto current  For earbitererauple  so we write $\hat{I} = L I$ . $L = \frac{10N^2S}{2}$
L is the industance. If I change with t
E= -d\$ =-LdZ
Field , B carries energy be cause
we must do work to create the carrents and keep them going  Dower that must be supplied to
Mantain a e wront I in a circur, because
the induced EHF=ELdI
P=dW = charge point time > Valtos det
2-IE
So = dV = -IE=+ IddI

We want an expression for which has more details - B(r) varies with space want every density

For Electric fields the stored every density,

Cherry per unit volume 15 & E2

need expression for magnetic contribution

This is important for understanding

energy & momentum carried by a

beam of light,

Last Time energy in magnetic field W= 1 LI2 energy required to make magnetie field We want to get one epression in terms of Birth. Why? PECall energy density of me electric field is I GEZ, EOB are two sides of Same coin. Want theory electromagnetum When computing energy need both every flows between electren a maquetic

The LI = SB-da W= \frac{1}{2} \land \frac{1}{ W= 1 SAF de restanos aunas

Epiter m fiell

Magnetosistation

Magnetosistation W= 1 A. (7×B) dC B-Zutby
parts  $\vec{\nabla} \cdot (\vec{A} \times \vec{B}) = \vec{B} \cdot (\vec{\nabla} \times \vec{A}) - \vec{A} \cdot (\vec{\nabla} \times \vec{B})$ = L Satorio - Sala. (Aris) 2M Sobomda alv

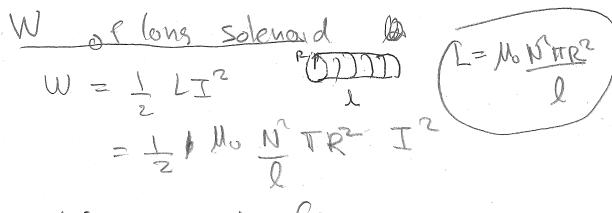
Integrationeral space 5-00 rade h Magnetuslatus
partofstory 2 der

This is not the whole story

ar important element is missing 
with this arguments of Far

lightnesses would not propagate

## Mackart Formula



get from W= 1 2 No St de

M=TBJUST

12=12 l

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before

	Desc		
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•	7. B 20		
Foodsy	74B=-2B		
Ampere	マ×B=Mot	うこうして	<b>?</b>
EXT	ra Conservationo	& charge T	0= 76+D17
LPL	P = 9 LÉ +3	( Ž 5	0
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	But there are anim	fuite & of	possible surfaces a that bound 8

Capacitor charging Mconsistency Further more month can take F. (TOB) = No TiJ 0 2 7.5 entitrae for steady carrat put  $\nabla . \vec{J} = -\partial P$  thousistancy

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Maxwell add aterm (not historically acrossly that wantained charge conservation サイトラールのサイズ ガーク・ガーカーア・ブナガ・ズ 7. X = MO 29 Look at Coulomb X = Mo E, DE 7. X = MOE 27. E = MO EODS So Marquell's eq changes Ampere to TO B = MOJA MO ED & Jo = displacement Current Eo DE 70 B = No (J+Jo) gB dlz Mo Jenchord + No Edt Now considerprevious example area O (J.da= I Joda = Eussala Tenelosedzo but Jo exists DE EA = O A de 2 de A 2 de J

50. da = eo dolle le 2delle =I

These are complete

## Magnetic monopoles

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whatis charge in Dne ver frame is

current in another. So \$ 15 transformed int B

T five look at Hanvell in case \$ =0, \$\bar{7} =0\$

have \$\bar{7} \cdot \bar{E} =0 \$\bar{7} \cdot \bar{E} = \bar{DB} \\
\$\bar{7} \cdot \bar{B} = \bar{D} \\
\$\bar{1} \cdot \bar{B} \\
\$\bar{1} \cdot \bar{D} = \bar{D} \\
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\$\bar{D} = \bar{D} \\
\$\bar{D} = \bar{D} \\

Equations invariant under transform

E 78

B 74060E

F Symmetry between of B of P,J zo

but with sources F.E = P/E6 T.B = 0

Major disterence between of B

We can make a more symmetric

Nersion of Marwel's Bans

7. = 8/6° 7 x E= Juo Jm - OB J.B=Pn Mo TXB = Mo Je +Mo GO JE but in mag. chargedens ity - it mustober continuity Pm +7. Im=0, 7. AxE) mustvanish This possibility has been ver attractive to physicists. There have beent least 1400 papers with title magnetic Monopoles Physicists have searched here thore o everywhere - North pole south pole Duterspace cosmice vars moon vocks Important in grand unified theory product too many Monopoles Inflationary unjuste Will talka' one monopole was found but it went away PRL 48,1378

Pollock, Phys. Rev. A 24, 1544 (1981).

<sup>3</sup>F. J. Rogers, H. E. De Witt, and D. B. Boercker, Phys. Lett. <u>82A</u>, 331 (1981); D. B. Boercker, Phys. Rev. A <u>23</u>, 1969 (1981); D. B. Boercker, F. J. Rogers, and H. E. De Witt, Phys. Rev. A <u>25</u>, 1623 (1982).

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<sup>9</sup>R. J. Bearman and J. G. Kirwood, J. Chem. Phys. 28, 136 (1958).

ring magniflip create E s accept

## First Results from a Superconductive Detector for Moving Magnetic Monopoles

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A velocity- and mass-independent search for moving magnetic monopoles is being performed by continuously monitoring the current in a  $20\text{-cm}^2$ -area superconducting loop. A single candidate event, consistent with one Dirac unit of magnetic charge, has been detected during five runs totaling 151 days. These data set an upper limit of  $6.1\times10^{-10}~\mathrm{cm}^{-2}$  sec<sup>-1</sup> sr<sup>-1</sup> for magnetically charged particles moving through the earth's surface.

PACS numbers: 14.80.Hv

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The detection of a moving magnetic charge with a superconducting ring is based solely on the long-range electromagnetic interactions between the magnetic charge and the macroscopic quantum state of the ring.¹ Such a detector measures a moving particle's magnetic charge regardless of its velocity, mass, electric charge, or magnetic dipole moment. In this paper, the first experimental results from use of this scheme are presented.

Superconductors make natural magnetic charge detectors, as suggested by comparing the flux quantum of superconductivity  $\varphi_0=hc/2e$  with the flux emanating from a single Dirac charge  $4\pi g=hc/e$ . Dirac² was led to his value for the elementary magnetic charge by postulating that the wave function of a single electron in the field of a pole should be single valued. In superconductivity, the postulate of a single-valued macroscopic wave function leads to flux quantization. The factor of 2 arises from the electric charge, 2e, of the Cooper pairs.

Consider a magnetic charge g moving at velocity v along the axis of a superconducting wire ring of radius b. Integrating Maxwell's generalized equation for the monopole current,  $\operatorname{curl}(\vec{\mathbf{E}}) + (1/c)d\vec{\mathbf{B}}/dt = -(4\pi/c)\vec{\mathbf{j}}_m$ , over the area  $S_\Gamma$  in the plane of the ring, we obtain

$$\oint_{\Gamma} \vec{E} \cdot d\vec{l} + c^{-1} \left( d\varphi / dt \right) = - \left( 4\pi g / c \right) \delta(t), \qquad (1)$$

where  $S_{\Gamma}$  is bounded by a path  $\Gamma$  that is everywhere inside the wire. If we neglect the finite response time of the superelectrons,  $\vec{E}$  will vanish along  $\Gamma$  and  $\varphi(t) = -4\pi g \theta(t)$ , where we set  $\varphi = 0$  for  $t = -\infty$ . The total flux through  $S_{\Gamma}$  is  $\varphi = \varphi_g + \varphi_g$ ,  $\varphi_g$  from the monopole and  $\varphi_s$  from the induced supercurrent. We find

 $\varphi_{g}(t) = 2\pi g \left[1 - 2\theta(t) + \frac{\gamma vt}{\left[(\gamma vt)^{2} + b^{2}\right]^{1/2}}\right],$ 

and  $\varphi_s = -I(t)L$ , where I(t) is the induced supercurrent and L the self-inductance of the ring. Thus, substituting  $4\pi g = 2\varphi_0$ , we obtain

$$I(t) = \frac{\varphi_0}{L} \left[ 1 + \frac{\gamma vt}{\left[ (\gamma vt)^2 + b^2 \right]^{1/2}} \right].$$

This result is independent of the choice of surface  $S_{\Gamma}$  bounded by the path  $\Gamma$  and corresponds to a change of  $2\varphi_0$  through the ring (Fig. 1). The change in current will occur with a characteristic time given by  $b/\gamma v$ .

In the general case, any trajectory of a magnetic charge g which passes through the ring will result in a flux-quanta change of 2, while one which misses the ring will produce no flux change. In the less likely event that a magnetic charge passes through the ring wire itself, it will leave a trapped doubly quantized vortex, and some intermediate total current will persist. Any electric charge or magnetic dipole moment of the particle

1378 Jm = g hv S (2-wh) S(A) S(4)
= 9 h S (K) S(A) S(4)

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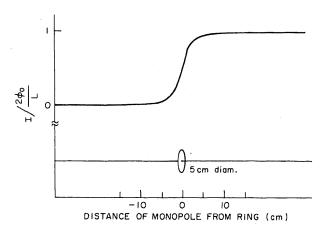


FIG. 1. Induced current in a superconducting ring for an axial monopole trajectory.

will cause only small transient fluctuations and no dc shifts. Thus, moving magnetically charged particles can be detected by monitoring the current in the superconducting ring.

In this experiment a four-turn, 5-cm-diam loop, positioned with its axis vertical, is connected to the superconducting input coil of a SQUID (superconducting quantum interference device) magnetometer.3 The passage of a single Dirac charge through the loop would result in an  $8\varphi_0$  change in the flux through the superconducting circuit, comprised of the detection loop and the SQUID input coil (a factor of 2 from  $4\pi g = 2\varphi_0$  and of 4 from the turns in the pickup loop). The SQUID and loop are inside a 20-cm-diam, 1-m-long cylindrical superconducting shield closed at the bottom, and these are mounted inside a single Mumetal cylinder. The combined shielding provides 180dB isolation from external magnetic field changes and an ambient field of 5×10-8 G.4

The voltage output of the SQUID electronics, which is directly proportional to the supercurrent in the detection loop, is continuously recorded through a 0.1-Hz low-pass filter onto a stripchart recorder. In addition, several times per day digital voltmeter readings are taken to guard against recorder failures.

The detector sensitivity has been calibrated in three independent ways: (a) by measuring the SQUID response to a known current in calibration Helmholtz coils and calculating their mutual inductance to the superconducting loop ( $\pm 4\%$ ); (b) by estimating the self-inductance of the superconducting circuit ( $\pm 30\%$ ); and (c) by directly observing flux quantization within the superconducting circuit ( $\pm 10\%$ ). All three methods agree with-

in their independent uncertainties.

Two additional effects influence the exact detector response. First, a magnetic monopole whose trajectory intersects either the transformer loop in the SQUID, the twisted leads from the SQUID to the loop, or the loop wire itself would produce a shift of nonintegral magnitude. Computation of the average area ratio of the loop to the remainder of the transformer circuit indicates that such events will be suppressed by a factor of 25 compared to loop events. Second, a particle traversing the superconducting shield will leave doubly quantized vortices wherever the trajectory intersects a wall. The effect is a magnetic field change inside the shield and an applied flux change across the loop. The total induced current change in the loop is  $\Delta I = (8\varphi_0/L)[\eta - \xi(A_1/A_S)]$ , where  $A_1/A_s = 0.06$  is the ratio of the loop to shield cross-sectional areas. For a trajectory that intersects the loop,  $\eta = 1$ , and for one that misses.  $\eta = 0$ . The geometric factor  $\xi$  depends on the trajectory impact parameter and inclination angles. and has maximum value of 1 for axial trajectories through the shield and a minimum of 0 for transverse ones. Current changes of  $(0.06)8\varphi_0/L$  or less will be observed for trajectories that pass through the shield but not the loop. The probability for such events with  $\Delta I > (0.02)8\varphi_0/L$  is about 10 times larger than for the loop.

As of 11 March 1982 data have been recorded for a total of 151 days. Several intervals throughout a continous one-month time period are shown in Fig. 2(a), where no adjustment of the dc level has been made. Typical disturbances caused by daily liquid-nitrogen and weekly liquid-helium transfers are evident. A single large event was recorded [Fig. 2(b)]. It is consistent with the passage of a single Dirac charge within a combined uncertainty of  $\pm 5\%$  (resulting from the calibration uncertainty and the distribution of geometric factor  $\xi$ ). It is the largest event of any kind in the record. In Fig. 3 are plotted the 27 events exceeding a threshold of  $0.2\varphi_0$ , which remain after exclusion of known disturbances such as transfers of liquid helium and nitrogen.<sup>5</sup> An event is defined as a sharp offset with well-defined stable levels for at least 1 h before and after. Only six events were recorded during the 70% of the running time when the laboratory was unoccupied.

The following statements about spurious detector response can be made:

(a) Line voltage fluctuations caused by two power outages and their accompanying transients

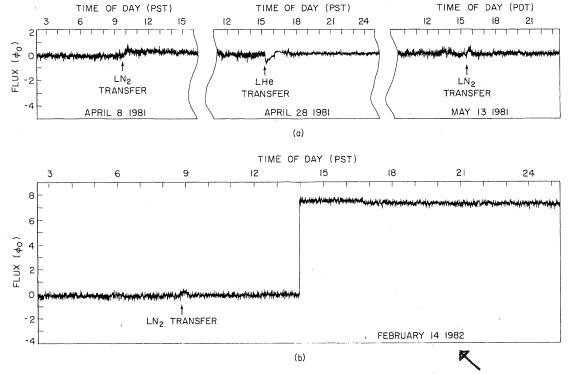


FIG. 2. Data records showing (a) typical stability and (b) the candidate monopole event.

failed to cause detectable offsets.

- (b) rf interference from the motor brushes of a heat gun failed to produce any offsets when operated in close proximity to the detector.
- (c) External magnetic field changes are attenuated by 180 dB, primarily from an exponential factor of  $e^{-1.63z/a}$ , where z=72 cm is the distance in from the open top of the superconducting shield and a=10 cm is the shield radius.
- (d) Ferromagnetic contamination is minimized using clean-bench assembly techniques and

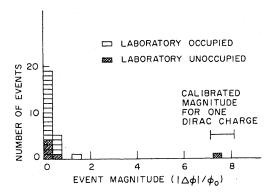


FIG. 3. Histogram of all event magnitudes.

- checked with magnetometer measurements within the shield.
- (e) The critical current of the loop is not reached for currents a thousand times greater than  $8\varphi_0/L$  and is typically  $10^8$  times greater.
- (f) Mechanically induced offsets have been intentionally generated and are probably caused by shifts of the four-turn loop-wire geometry which produce inductance changes. Sharp raps with a screwdriver handle against the detector assembly cause such offsets. On two occasions out of 25 attempts these have exceeded  $6\varphi_0$  (75% of the shift expected from one Dirac charge); however, drifts in the level were seen during the next hour.
- (g) No seismic disturbance occurred on 14 February 1982.
- (h) Energetic cosmic rays depositing  $\lesssim 1~\text{GeV/cm}$  in traversing the wire would raise the local wire temperature by only  $\lesssim 0.01~\text{K}$ , but a 5-K change is needed to reach the critical temperature.

A spontaneous and large external mechanical impulse is not seen as a possible cause for the event; however, the evidence presented by this single event does not preclude the possibility of a spontaneous internal stress release mechanism. Regardless, to date the experiment has set an

upper limit of  $6.1 \times 10^{-10}$  cm<sup>-2</sup> sec<sup>-1</sup> sr<sup>-1</sup> for the isotropic distribution of any moving particles with magnetic charge greater than 0.06 g.

An observational upper bound on the mass density of monopoles is given by limits on the local "missing mass." 6 Visible matter has a measured local density of 0.09  $M_{\odot}/\mathrm{pc^3}$  (solar masses per cubic parsec), whereas the mass density calculated from the velocity distribution out of the galactic plane is  $0.14 M_{\odot}/pc^{3}$ . This local "missing mass" density estimate of  $0.05 M_{\odot}/pc^3$  is in good agreement with the halo mass estimates extrapolated back to our local galactic radius, which give  $0.03 M_{\odot}/pc^3$ . If we assume this entire "hidden mass" to be made up of monopoles of mass  $10^{16}$  GeV/ $c^2$  with isotropic velocities of order 300 km/sec, as suggested from grand unification theories, the number passing through the earth's surface would be 4×10<sup>-10</sup> cm<sup>-2</sup> sec<sup>-1</sup> sr<sup>-1</sup>.8 This would result in 1.5 events per year through the detector loop.

The search with the present detector is being continued, and two new systems of larger sensing area are being built.

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<sup>1</sup>L. W. Alvarez, Lawrence Radiation Laboratory Physics Note No. 470, 1963 (unpublished); P. Eberhard, Lawrence Radiation Laboratory Physics Note No. 506, 1964 (unpublished); L. J. Tassie, Nuovo Cimento 38, 1935 (1965); L. Vant-Hull, Phys. Rev. 173, 1412 (1968); P. Eberhard, D. Ross, L. Alvarez, and R. Watt, Phys. Rev. D 4, 3260 (1971).

<sup>2</sup>P. A. M. Dirac, Proc. Roy. Soc. London, Ser. A <u>133</u>, 60 (1931), and Phys. Rev. 74, 817 (1948).

<sup>3</sup>Model 330x SQUID system, manufactured by S. H. E. Corp., San Diego, Cal. For theory see J. Clarke, Proc. IEEE 61, 8 (1973).

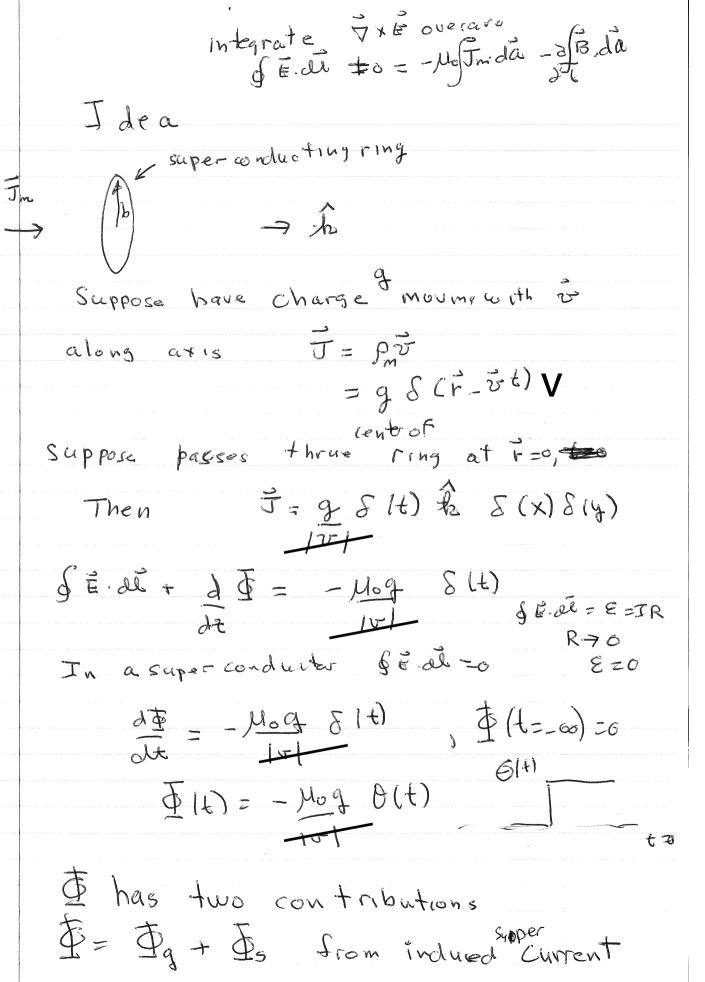
<sup>4</sup>B. Cabrera, Ph.D. thesis, Stanford University, 1974 (unpublished); B. Cabrera and F. van Kann, Acta Astronaut. <u>5</u>, 127 (1978).

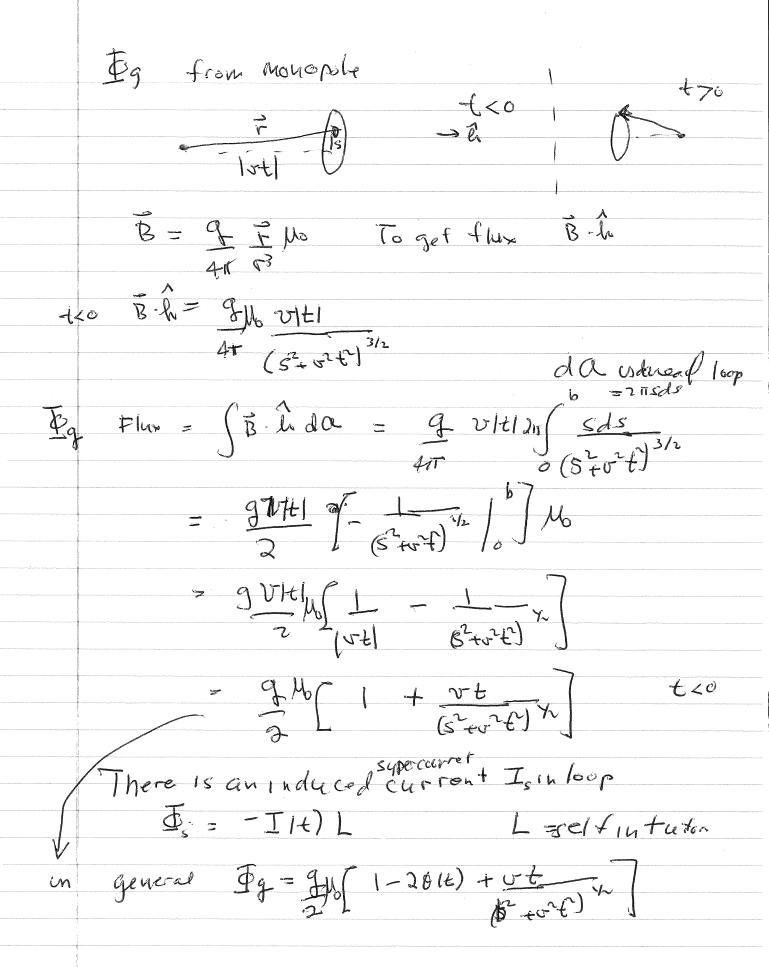
 $^5\mathrm{An}$  event during the first run, of approximate magnitude  $7\varphi_0$  and which left the recorder range, is directly correlated with probe assembly work for another experiment next to the detector. It occurred before the installation of a mechanical stabilizing frame and has been ignored.

<sup>6</sup>For review, see S. M. Faber and J. S. Gallagher, Ann. Rev. Astron. Astrophys. <u>17</u>, 135 (1979).

<sup>7</sup>J. P. Preskill, Phys. Rev. Lett. <u>43</u>, 1365 (1979); G. Lazarides, Q. Shafi, and T. F. Walsh, Phys. Lett. 100B, 21 (1981).

 $^8$ Recently J. D. Ullman, Phys. Rev. Lett. <u>47</u>, 289 (1981), has reported on the use of a proportional counter array to set an upper flux limit of  $3\times10^{-11}$  cm<sup>-2</sup> sec<sup>-1</sup> sr<sup>-1</sup>. However, differing estimates for slow monopole ionization rates [S. Geer and W. G. Scott, CERN  $\bar{p}p$  Note 69, April 1981 (unpublished); K. Hayashi, to be published; J. S. Trefil, to be published] indicate the complexities of such calculations. These calculations are crucial to predicting whether slow monopoles are observable in a given ionization detector.





$$I = \frac{9 \mu_0}{2 L} \left[ \frac{1 + v + v}{2 L} \right]$$

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$$I = \frac{9 \mu_0}{2 L} \left[ \frac{1 + v + v}{2 L} \right]$$

$$I = \frac{9 \mu_0}{2 L}$$

Marwell equs in Matter
In matter there are bound electrice
and magnetic dipoles (M)
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bound charge density = 86z-7. P
Have also written Maswell ean ingeneral
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free charges 95, free current Is
& want to use H & D. This done for
He statuc case. Now we have
to consider other effects This.
involves polarization - dipple moments/voluces consider tiny chunk of material that is polarized.  OPD There is a charge density ob=P
on right and $\sigma_b = -P$ (from $\sigma_b = \vec{I} \cdot \hat{n}$ )
Now suppose P increases a bit. Then there must be a cyrrent from teft to right
dI = 206 da, = 2P da,

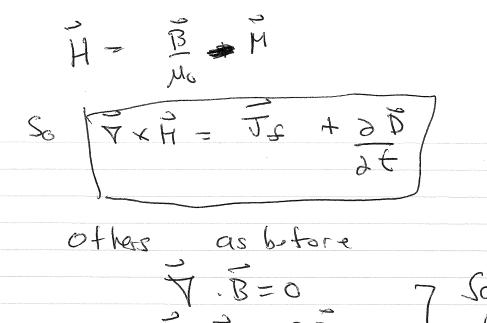
In general the current is carrent dens, to times area There is a current density

Je = 21

Polarization current This current density is consistent Coutinuity eq. Recall that Sb = -7. P  $\vec{\nabla} \cdot \vec{J}_{P} = \vec{\partial} \vec{\nabla} \cdot \vec{P} = - \vec{\partial} \vec{P}_{0} \quad \text{HES}^{!} \quad \text{Includins}$ Jp essential for rouservation of bound charge. Now let's write all the so areas charges and cyrrents 9=9s+Pb=9s-7.P J=Js+Jb+Jp=Js+7×M+2P Gauss Law 7. E = P = 1 (Ps - Y.B)

Fe call D = E + P 7.5 = Ps Ampere Maxwoll

T × B = MoT = Mo (Ix + T T × M + 2P) +MOGO DE



others as before

7.B=0

7xE=-2B

free

26

These are May well eque in media

To proceed heed to know relation between Dit and

B, n depends on properties of material

For Linear Media

P= EoXe F H = M2 m H D=6E+D=60(1+Xe)F=EE 18 = Mo (1+Xm) H 一儿前