$$\begin{array}{c} \mbox{Tuesday } Oct 1, 2019 \\ \hline \\ \mbox{Stokes' Theorem:} \\ \end{tabular} \\ \mbox{Stokes' Theorem:} \\ \mbox{Stoke$$



Example:
$$\vec{\nabla}(x,y_1,z) = 4y\hat{i} + x\hat{j} + 2z\hat{k}$$

Compute $\int (\vec{\nabla} \cdot \vec{\nabla}) \cdot d\vec{a}$ over the hemisphere
 $x^2 + y^2 + z^2 = R^2$, $z \ge 0$. We'll do this 3 ways.
1) Direct evaluation:
 $\vec{\nabla} \cdot \vec{\nabla} = det \begin{vmatrix} \partial_x & \partial_y & \partial_z \\ V_x & V_y & V_z \\ \hat{i} & \hat{j} & \hat{k} \end{vmatrix}$
 $= (\frac{\partial V_z}{\partial Y} - \frac{\partial V_z}{\partial z})\hat{i} + (\frac{\partial V_x}{\partial z} - \frac{\partial V_z}{\partial y})\hat{j} + (\frac{\partial V_y}{\partial x} - \frac{\partial V_x}{\partial y})\hat{k}$
 $= 0\hat{i} + 0\hat{j} + (1-4)\hat{k} = -3\hat{k}$

$$d\vec{a} = R^{2} \sin\theta d\theta d\phi \hat{r} \Rightarrow \hat{k} \cdot \hat{r} = \cos\theta$$
$$\int (\vec{\nabla} \times \vec{v}) d\vec{a} = \int_{0}^{21} d\phi \int_{0}^{\pi_{2}} (-3) R^{2} \sin\theta \cos\theta d\theta = -6\pi R^{2} \int_{0}^{1} u du = -3\pi R^{2}.$$
$$u = \sin\theta$$

2) Stokes' theorem
$$\int_{S} (\bar{\nabla} x \bar{\nabla}) d\bar{a} = \oint_{P} \bar{\nabla} \cdot d\bar{\ell}$$

The boundary is a circle $\chi^{2} + \gamma^{2} = R^{2}$, $Z = 0$.
Which way around? Right hand rule \rightarrow counter clockwise.
 $d\bar{\ell} = Ra\phi \hat{\phi} = Rd\phi (\cos\phi \hat{j} - \sin\phi \hat{\iota})$
 $\bar{\nabla} \cdot d\bar{\ell} = Rd\phi (\chi \cos\phi - 4\gamma \sin\phi) \qquad \chi = R\cos\phi , \gamma = R\sin\phi$
 $= Rd\phi (R\cos^{2}\phi - 4R\sin^{2}\phi)$
 $\oint_{P} \bar{\nabla} \cdot d\bar{\ell} = \int_{0}^{2\pi} R^{2} (\cos^{2}\phi - 4S\sin^{2}\phi) d\phi$
 $= R^{2} \cdot (\pi - 4\pi) = -3\pi R^{2}$
3) Use Stokes' theorem to relate integral over the

3) Use Stokes' theorem to relate integral over the hemisphere to an integral in the plane with the same boundary. Now $d\vec{a} = dxdy \hat{k}$ (sign from RHR) $\vec{k} = \vec{k} + \vec{k} = -3\hat{k} \Rightarrow (\vec{\nabla} x \vec{v}) \cdot d\vec{a} = -3 dxdy$ $\int_{x}^{R} (\vec{\nabla} x \vec{v}) \cdot d\vec{a} = -3 \int_{x}^{S} dxdy = -3 \pi R^{2}$

Dirac S function
Suppose
$$\vec{V}(\vec{r}) = \frac{1}{r^2} \hat{r} = \frac{\vec{r}}{r^3}$$

What is the divergence of \vec{V} ?
Using the formula for $\vec{r} \cdot \vec{v}$ in spherical coordinates:
 $\vec{\nabla} \cdot \vec{v} = \frac{1}{r^2} \frac{\partial}{\partial r} (r^2 V_r) + \frac{1}{r \sin \theta} \frac{\partial}{\partial \theta} (V_\theta \sin \theta) + \frac{1}{r \sin \theta} \frac{\partial}{\partial \phi} V_{\phi}$
 $= \frac{1}{r^2} \frac{\partial}{\partial r} (1) + 0 + 0 = 0$
This is strange, because $\vec{V}(\vec{r})$ definitely looks "divergencey".
What's going on?
Let's try with the divergence theorem over some
sphere of vadius R centered on $r=0$.
 $\int [\vec{\nabla} \cdot \vec{v}] d^2r = \int \vec{\nabla} \cdot d\vec{a}$ $d\vec{a} = R \sin \theta d\theta d\phi \hat{r}$
 $= \iint \frac{1}{R^2} R^2 \sin \theta d\theta d\phi$
 $= 4\pi$ independent of R (!?!)
The source of the trouble is that all of the
divergence in $\vec{V}(\vec{r})$ is located at $r=0$. The
formula for $\vec{\nabla} \cdot \vec{v}$ in spherical coordinates involves
 $\vec{\tau}^2$ and $\vec{\tau} \sin \phi$, so it cannet be valid at $r=0$.

We also notice that $|\vec{v}(\vec{r})| \rightarrow \infty$ as $r \rightarrow 0$

To summarize, $\vec{\nabla} \cdot \vec{\nabla}$ is zero everywhere except for at r=0, where it is 00. But it is 00 in such a way that integrating over it gives a finite value, namely 4π .

This is a strange situation, but it is very important in E#M when dealing with point charges, so we should figure out how to deal with it. These properties are described by the Dirac delta function.

Dirac S in 1D:

$$S(x) = \begin{cases} 0 & \text{if } x \neq 0 \\ \infty & \text{if } x = 0 \end{cases}, \quad \int_{a}^{b} \delta(x) \, dx = \begin{cases} 1 & \text{if } a < 0 < b \\ 0 & \text{otherwise} \end{cases}$$





"Sifting" property: Consider some "normal" (i.e. finite) function
$$f(x)$$

$$\int_{-\infty}^{\infty} \delta(x) f(x) dx = f(0)$$
This is the most important feature of a S function.
Note that the integral need not run from -20 to 20.
It just needs to include O.

$$f(x)$$

To understand this property, let's use the limit of the a rectangle step function, which allows us to limit the integral $\int_{a}^{\infty} \int_{-\infty}^{\infty} \delta(x) f(x) dx = \lim_{a \to 0} \int_{-a}^{a} \frac{1}{2a} f(x) dx$

and expand
$$f(x)$$
 as a Taylor series about $x=0$:

$$f(x) = f(0) + f'(0)x + \frac{1}{2}f'(0)x^{2} + \frac{1}{3!}f''(0)x^{3} + \dots$$

So
$$\int_{-\infty}^{\infty} \delta(x) f(x) dx = \lim_{a \to 0} \frac{1}{2a} \left[\int_{-a}^{a} f(o) dx + \int_{a}^{a} f'(o) x dx + \int_{-a}^{a} \frac{1}{2} f''(o) x^{2} dx + \dots \right]$$

Now $\int_{-a}^{a} dx = 2a$, $\int_{-a}^{a} x dx = \frac{1}{2} (a^{2} - a^{2}) = 0$, $\int_{-a}^{a} x^{2} dx = \frac{1}{3} (a^{3} - (-a)^{3}) = \frac{2}{3} a^{3}$

$$So \int_{-\infty}^{\infty} \delta(x) f(x) dx = \lim_{\alpha \to 0} \frac{1}{2\alpha} \left[f(0) \cdot 2\alpha + f'(0) \cdot 0 + f''(0) \cdot \frac{2}{3}\alpha^3 + \dots \right]$$
$$= \lim_{\alpha \to 0} \left[f(0) + f''(0)\alpha^2 \cdot \frac{1}{3} + \dots \right]$$

In general, each term in the series has the form
$$f_{(o)}^{(m)}a^n \cdot [f_{inite number}]$$
, and all terms aside from $f(o)$ disappear in the limit $a \rightarrow 0$, and $\int_{-\infty}^{\infty} \delta(x) f(x) dx = f(o)$.

(Note that this only holds if the derivatives of fare finite at X=0.)

In practice, we don't need a to literally be Zero. We just need it so small that find an is negligible compared to, say, the precision of a measurement we can make. This means that the function f(x) should not change appreciably over a length scale a.

Coming back to electrodynamics, if we have some charge distribution p(x) which is localized within some length scale a, and we place it in an electric field $\vec{E}(x)$ which varies on a length scale $L \gg a$, then we will not make much error by assuming that all of the charge is located at a single point: $p(x) \rightarrow q S(x), q = \int_{a}^{a} p(x) dx.$ Now, back to the properties of the S function. We can shift the origin: $\delta(x-x_0)$ then $\int_{-\infty}^{\infty} \delta(x-x_0) f(x) dx = f(x_0)$ Units: Since $\int_{-\infty}^{\infty} \delta(x) dx = 1$, dx has units of meters, so $\delta(x)$ has units of density, i.e. $\frac{1}{m}$. Scaling: $\delta(kx) = \frac{1}{|k|} \delta(x)$ see Griffiths Ex. 1.15.

3D S function:

$$\delta^{(s)}(\vec{r}) = \delta(x)\delta(y)\delta(z) \quad \Leftrightarrow \text{ Units are } \frac{1}{m^3}$$

$$\int d^3r \ \delta^{(s)}(\vec{r}) = 1$$

$$\int d^3r \ \delta^{(s)}(\vec{r}-\vec{r}_s) f(\vec{r}) = f(\vec{r}_s)$$

Notation: Griffiths uses $S^{(3)}_{(2)}$ for a 3D 8. Other authors use S(2). Either way is fine, and ambiguity should not arise because if the argument is a vector, each element of the vector must be zero. Finally, back to $\vec{\nabla} \cdot \vec{\nabla}$. We found that $\vec{\nabla} \cdot \vec{\nabla} = \begin{cases} \infty & \vec{\tau} = 0 & \text{and} & \int d^3 r \ \vec{\nabla} \cdot \vec{\nabla} = 4\pi \\ 0 & \text{otherwise} \end{cases}$ So this can be explained by setting $\vec{\nabla} \cdot \left(\frac{\hat{r}}{r^2}\right) = 4\pi \delta^{(3)}(\vec{r})$. Also, since $\vec{\nabla} \left(\frac{1}{r}\right) = -\frac{\hat{r}}{r^2}$, we have $\vec{\nabla} \cdot \vec{\nabla}(\vec{+}) = \nabla^2 \left(\frac{1}{r}\right) = -4\pi \delta^{(5)}(\vec{r})$

[Bonus material, time permitting] When setting $p(\vec{r}) \propto S(\vec{r})$ works: gravity at Earths surface Force on a spherical cow: $\vec{F} = \int \vec{g}(\vec{r}) p(\vec{r}) d^{3}r$ Take $p_{mos}(\vec{r}) \rightarrow M_{cov} S(\vec{r}_{cov})$ $\vec{F} = M_{cov} \vec{g}(\vec{r}_{cov})$ $\vec{g}(\vec{r}) = \frac{-GMo}{r^{2}} \hat{r}$ What error have we made? Cow size is $a \sim 1m$. Leading correction $\sim \frac{\partial q}{\partial r} a \sim \frac{GMo}{r^{3}} a \approx g(R_{o}) \frac{a}{R_{o}}$. Correction is $\sim \frac{1m}{6 \times 10^{6}m}$ (small). Since G is only known to 5 sig figs, this is fine.]

When
$$p(\vec{r}) \propto S(\vec{r})$$
 doesn't work:
Self-energy of a point charge. (This will be covered
later, but hopefully you can believe for now) the
energy to construct a charge distribution $p(\vec{r})$
is
 $W = \frac{1}{2} \int p(\vec{r}) V(\vec{r}) d^3r$
here $V(\vec{r})$ is the electrostatic potential. For a point
charge, $p(\vec{r}) \rightarrow q S^{(3)}(\vec{r})$, $V(\vec{r}) = \frac{q}{4\pi\epsilon_0 r}$
so
 $W \rightarrow \frac{1}{2} q \cdot \frac{q}{4\pi\epsilon_0 0} = \infty$. Uh oh.
The source of the trouble is evident
if we go back and try to Taylor
expand $V(r)$ about $r=0$:
 $V(\circ) = \lim_{r\to 0} \frac{q}{4\pi\epsilon_0} (\frac{1}{r} - \frac{1}{r^2} + \frac{1}{r^3} - \frac{1}{r^4} + ...)$

Every term blows up at $r \rightarrow 0$, so the requirement that $f_{ini}^{(m)}a^n$ go to zero doesn't hold. V(r) is not approximately constant over any length scale around r=0and so the replacement $p(r) \rightarrow q_i S_{ir}^{(r)}$ doesn't work.

Another way to look at it is that $V(\neq)$ depends on the details of $p(\neq)$ in a way that doesn't vanish when the size of $p(\neq)$ shrinks to zero.

Thursday Oct 3, 2019

Electrostatics

Coulomb's Law

Force on "test" charge Q at \vec{r} due to a 'so charge Q, at \vec{r}' 元 Source" **≯**∘ Q $\vec{F} = \frac{qQ}{4\pi\epsilon} \frac{\vec{r} \cdot \vec{r}'}{|\vec{r} \cdot \vec{r}'|} = \frac{qQ}{4\pi\epsilon} \hat{f}$ where $\vec{\lambda} \equiv \vec{r} \cdot \vec{r}'$ points from source to observer. $E_0 = 8.85 \times 10^{-12} C^2$ "permittivity of free space" electron charge: $e = -1.6 \times 10^{-19} C$ \rightarrow Force between 2 electrons @ 1m: $\approx 2 \times 10^{-28}$ N. Superposition : Force on test charge Q due to all source charges 9, 22 23 ... is just the sum of the Q individual forces $\vec{F} = \vec{F} + \vec{F} + \vec{F} + ...$ $= \frac{Q}{4\pi c_{o}} \left[\frac{q_{1}}{|\vec{r} - \vec{r}_{1}|^{s}} (\vec{r} - \vec{r}_{1}) + \frac{q_{2}}{|\vec{r} - \vec{r}_{2}|^{s}} (\vec{r} - \vec{r}_{2}) + \dots \right]$ $\equiv Q \overline{E}(a)$

Electric field: $\vec{E}(\vec{r}) = \frac{1}{4\pi\epsilon_o} \sum_i \frac{q_i}{|\vec{r} - \vec{r}_i|^3} (\vec{r} - \vec{r}_i)$ Force per unit charge

What is a field?

Without bothering too much about philosophy, a field is a quantity (like a number, or a vector-i.e. 3 numbers) assigned to every point in space.



simplify, choose origin so that
$$\vec{r}=0$$
, $\vec{k}=-\vec{r}'$.

$$\vec{E}(\vec{r}=0) = \frac{1}{4\pi\epsilon_0} \int p(\vec{r}') \frac{(-\hat{r}')}{r'^2} r'^2 \sin\theta' dr' d\theta' d\phi'$$
powers of r cancel, no disaster



Example: Uniformly charged square



Do the integral for
$$E_z$$
, Since $E_x = E_y = 0$.
 $E_z(z) = \frac{1}{4\pi\epsilon_o} \int dq \frac{\hat{h}\cdot\hat{k}}{\hbar^2} = \frac{1}{4\pi\epsilon_o} \int_{-\frac{1}{2}}^{\frac{1}{2}} \int_{-\frac{1}{2}}^{\frac{1}{2}} \sigma dx dy \frac{h}{[x^2+y^2+k_0^2]^{3/2}}$

Not a particularly pleasant integral. Before pushing on,
let's look at some limiting cases.
Limit Z>>L
$$\Rightarrow \chi^2 + y^2 + z^2 \approx z^2 \Rightarrow E_z \approx \frac{1}{4\pi c_o} \frac{1}{z^2} \cdot \sigma L^2 = \frac{Q}{4\pi c_o} z^2$$

Limit Z<\Rightarrow \chi^2 + y^2 + z^2 \approx \chi^2 + y^2 \Rightarrow E_z \approx \frac{\sigma z}{4\pi c_o} \iint \frac{dxdy}{[\chi^2 + y^2]^{3/2}}
blows up@ x=y=0. uh oh.

What went wrong? In the neighborhood of x=y=0, we cavit make the approximation $x^2+y^2+z^2 \approx x^2+y^2$.



Êline = Eline k

It height
$$z$$
 above the
f charge of length $2w$.
 $dq = \lambda dx$ $\vec{\mathcal{A}} = -x\hat{i} + z\hat{k}$
By symmetry, we only need
to worry about the \hat{k}
component of \vec{E} .
 $\hat{\mathcal{A}}\cdot\hat{k} = \cos\theta$, $\mathcal{L} = \frac{z}{\cos\theta}$

$$\chi = Z \tan \theta$$
, $dx = Z \frac{1}{\cos^2 \theta} d\theta$

$$\begin{split} E_{\text{line}}(z) &= \frac{1}{4\pi\epsilon_{o}} \int \frac{dq}{\hbar^{2}} \hat{h} \cdot \hat{k} = \frac{1}{4\pi\epsilon_{o}} \int \frac{\lambda z d\theta}{\cos^{2} \theta} \frac{\cos^{2} \theta}{z^{2}} \cos \theta \\ &= \frac{1}{4\pi\epsilon_{o}} \hat{\lambda} = \int \frac{\theta_{w}}{\cos \theta} d\theta \\ &= \frac{1}{4\pi\epsilon_{o}} \hat{\lambda} = \int \frac{\theta_{w}}{\cos \theta} d\theta \\ &= \frac{1}{4\pi\epsilon_{o}} \frac{\lambda}{z} \left(\sin \theta_{w} - \sin \theta_{-w} \right) \\ &= \frac{1}{4\pi\epsilon_{o}} \frac{\lambda}{z} \left(\sin \theta_{w} - \sin \theta_{-w} \right) \\ &= \frac{2\lambda}{4\pi\epsilon_{o}} \frac{\lambda}{z} \left(\sin \theta_{w} - \sin \theta_{-w} \right) \\ &= \frac{2\lambda}{4\pi\epsilon_{o}} \frac{\lambda}{z} \left(\sin \theta_{w} - \sin \theta_{-w} \right) \\ &= \frac{2\lambda}{4\pi\epsilon_{o}} \frac{\lambda}{z} \left(\sin \theta_{w} - \sin \theta_{-w} \right) \\ &= \frac{2\lambda}{4\pi\epsilon_{o}} \frac{\lambda}{z} \left(\sin \theta_{w} - \sin \theta_{-w} \right) \\ &= \frac{2\lambda}{4\pi\epsilon_{o}} \frac{\lambda}{z} \left(\sin \theta_{w} - \sin \theta_{-w} \right) \\ &= \frac{2\lambda}{4\pi\epsilon_{o}} \frac{\lambda}{z} \left(\sin \theta_{w} - \sin \theta_{-w} \right) \\ &= \frac{2\lambda}{4\pi\epsilon_{o}} \frac{\lambda}{z} \left(\sin \theta_{w} - \sin \theta_{-w} \right) \\ &= \frac{2\lambda}{4\pi\epsilon_{o}} \frac{\lambda}{z} \left(\sin \theta_{w} - \sin \theta_{-w} \right) \\ &= \frac{2\lambda}{4\pi\epsilon_{o}} \frac{\lambda}{z} \left(\sin \theta_{w} - \sin \theta_{-w} \right) \\ &= \frac{2\lambda}{4\pi\epsilon_{o}} \frac{\lambda}{z} \left(\sin \theta_{w} - \sin \theta_{-w} \right) \\ &= \frac{2\lambda}{4\pi\epsilon_{o}} \frac{\lambda}{z} \left(\sin \theta_{w} - \sin \theta_{-w} \right) \\ &= \frac{2\lambda}{4\pi\epsilon_{o}} \frac{\lambda}{z} \left(\sin \theta_{w} - \sin \theta_{-w} \right) \\ &= \frac{\lambda}{4\pi\epsilon_{o}} \frac{\lambda}{z} \left(\sin \theta_{w} - \sin \theta_{-w} \right) \\ &= \frac{\lambda}{4\pi\epsilon_{o}} \frac{\lambda}{z} \left(\sin \theta_{w} - \sin \theta_{-w} \right) \\ &= \frac{\lambda}{4\pi\epsilon_{o}} \frac{\lambda}{z} \left(\sin \theta_{w} - \sin \theta_{-w} \right) \\ &= \frac{\lambda}{4\pi\epsilon_{o}} \frac{\lambda}{z} \left(\sin \theta_{w} - \sin \theta_{-w} \right) \\ &= \frac{\lambda}{4\pi\epsilon_{o}} \frac{\lambda}{z} \left(\sin \theta_{w} - \sin \theta_{-w} \right) \\ &= \frac{\lambda}{4\pi\epsilon_{o}} \frac{\lambda}{z} \left(\sin \theta_{w} - \sin \theta_{-w} \right) \\ &= \frac{\lambda}{4\pi\epsilon_{o}} \frac{\lambda}{z} \left(\sin \theta_{w} - \sin \theta_{-w} \right) \\ &= \frac{\lambda}{4\pi\epsilon_{o}} \frac{\lambda}{z} \left(\sin \theta_{w} - \sin \theta_{-w} \right) \\ &= \frac{\lambda}{4\pi\epsilon_{o}} \frac{\lambda}{z} \left(\sin \theta_{w} - \sin \theta_{-w} \right) \\ &= \frac{\lambda}{4\pi\epsilon_{o}} \frac{\lambda}{z} \left(\sin \theta_{-w} - \sin \theta_{-w} \right) \\ &= \frac{\lambda}{4\pi\epsilon_{o}} \frac{\lambda}{z} \left(\sin \theta_{-w} - \sin \theta_{-w} \right) \\ &= \frac{\lambda}{4\pi\epsilon_{o}} \frac{\lambda}{z} \left(\sin \theta_{-w} - \sin \theta_{-w} \right) \\ &= \frac{\lambda}{4\pi\epsilon_{o}} \frac{\lambda}{z} \left(\sin \theta_{-w} - \sin \theta_{-w} \right) \\ &= \frac{\lambda}{2} \left(\sin \theta_{-w} - \sin \theta_{-w} \right) \\ &= \frac{\lambda}{4\pi\epsilon_{o}} \frac{\lambda}{z} \left(\sin \theta_{-w} - \sin \theta_{-w} \right) \\ &= \frac{\lambda}{4\pi\epsilon_{o}} \frac{\lambda}{z} \left(\sin \theta_{-w} - \sin \theta_{-w} \right) \\ &= \frac{\lambda}{4\pi\epsilon_{o}} \frac{\lambda}{z} \left(\sin \theta_{-w} - \sin \theta_{-w} \right) \\ &= \frac{\lambda}{4\pi\epsilon_{o}} \frac{\lambda}{z} \left(\sin \theta_{-w} - \sin \theta_{-w} \right) \\ &= \frac{\lambda}{4\pi\epsilon_{o}} \frac{\lambda}{z} \left(\sin \theta_{-w} - \sin \theta_{-w} \right) \\ &= \frac{\lambda}{4\pi\epsilon_{o}} \frac{\lambda}{z} \left(\sin \theta_{-w} - \sin \theta_{-w} \right) \\ &= \frac{\lambda}{4\pi\epsilon_{o}} \frac{\lambda}{z} \left(\sin \theta_{-w} - \sin \theta_{-w} \right) \\ &=$$

Step 2.



This is an integral you can just ask Mathematica to do for you. But we can also make a clever substitution by drawing another triangle (Im not sure of its physical significance, but it helps make a good change of variables).

h

$$\begin{aligned}
\frac{1}{12(h^{2}+w^{2})} & tan\phi = \frac{h}{\sqrt{h^{2}+2w^{2}}} \\
\frac{1}{1005^{2}\phi} d\phi = \frac{-h}{2} \frac{4wdw}{[h^{2}+2w^{2}]^{3}h} \\
\frac{1}{1005^{2}\phi} d\phi = \frac{h^{2}+2w^{2}}{2(h^{2}+w^{2})} \\
\frac{1}{1005^{2}\phi} d\phi = \frac{h^{2}+2w^{2}}{2(h^{2}+w^{2})}$$

So
$$d\phi = \frac{h^2 + 2w^2}{2(h^2 + w^2)} \cdot \frac{-2hwdw}{[h^2 + 2w^2]^{3/2}} = \frac{-hwdw}{(h^2 + w^2)\sqrt{h^2 + 2w^2}}$$

$$E_{\text{square}} = \int_{0}^{\frac{42}{25 \cdot dW \cdot Wh}} \int_{0}^{\frac{25 \cdot dW \cdot Wh}{\pi \varepsilon_{0} (h^{2} + W^{2}) \sqrt{h^{2} + 2W^{2}}} = -\frac{25}{\pi \varepsilon_{0}} \int_{0}^{\frac{4}{2}} d\phi = \frac{25}{\pi \varepsilon_{0}} (\phi_{0} - \phi_{1/2})$$

$$\tan \phi_{\circ} = \frac{h}{h} = 1 \implies \phi_{\circ} = \frac{\pi}{4}, \quad \tan \phi_{1/2} = \frac{h}{\sqrt{h^2 + 2(1/2)^2}}$$

$$\overrightarrow{E}_{square} = \underbrace{\overline{2}}_{2\varepsilon_{\circ}} \left(1 - \frac{4}{\pi} \tan^{-1} \frac{1}{\sqrt{1 + \frac{1}{2}}} \right) \widehat{k}$$

Now taking the limit $h \ll L$ is easy $\tan^2 0 = 0$ so $\vec{E}_{square} \rightarrow \frac{1}{2}\vec{E}_{square} \hat{k}$.