

# Neutrino Mass and the Evolution of the Universe

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# The Story So Far

- Neutrinos were proposed as a solution to problems with  $\beta$ -decay.
- For years, the three varieties neutrinos were presumed to be massless.
- Eventually, it was found that neutrinos can change their flavor, indicating that they must have some mass.
- Neutrino oscillation is described by the “mixing matrix”:

$$\begin{pmatrix} \mathbf{v}_e \\ \mathbf{v}_\mu \\ \mathbf{v}_\tau \end{pmatrix} = \begin{pmatrix} c_{13}c_{12} & c_{13}s_{12} & s_{13}e^{-i\delta} \\ -c_{23}s_{12} - s_{13}s_{23}c_{12}e^{i\delta} & c_{23}c_{12} - s_{13}s_{23}s_{12}e^{i\delta} & c_{13}s_{23} \\ s_{23}s_{12} - s_{13}c_{23}c_{12}e^{i\delta} & -s_{23}c_{12} - s_{13}c_{23}s_{12}e^{i\delta} & c_{13}c_{23} \end{pmatrix} \begin{pmatrix} \mathbf{v}_1 \\ \mathbf{v}_2 \\ \mathbf{v}_3 \end{pmatrix}$$

# The Mixing Matrix

$$\begin{pmatrix} \mathbf{v}_e \\ \mathbf{v}_\mu \\ \mathbf{v}_\tau \end{pmatrix} = \begin{pmatrix} c_{13}c_{12} & c_{13}s_{12} & s_{13}e^{-i\delta} \\ -c_{23}s_{12} - s_{13}s_{23}c_{12}e^{i\delta} & c_{23}c_{12} - s_{13}s_{23}s_{12}e^{i\delta} & c_{13}s_{23} \\ s_{23}s_{12} - s_{13}c_{23}c_{12}e^{i\delta} & -s_{23}c_{12} - s_{13}c_{23}s_{12}e^{i\delta} & c_{13}c_{23} \end{pmatrix} \begin{pmatrix} \mathbf{v}_1 \\ \mathbf{v}_2 \\ \mathbf{v}_3 \end{pmatrix}$$

- The mixing matrix describes the probability of a neutrino with a definite mass (1, 2 or 3) to be in a particular flavor eigenstate (e,  $\mu$  or  $\tau$ ).
- The matrix depends on the phase  $\delta$ , and three mixing angles:  $\theta_{12}$ ,  $\theta_{23}$ , and  $\theta_{13}$ .  $\theta_{12}$  and  $\theta_{23}$  are known (approximately):

$$\theta_{12} \sim 32^\circ \quad \text{and} \quad \theta_{23} \sim 45^\circ$$

- The mixing angle  $\theta_{13}$  is not known, but is believed to be much smaller ( $< 3.2^\circ$ )

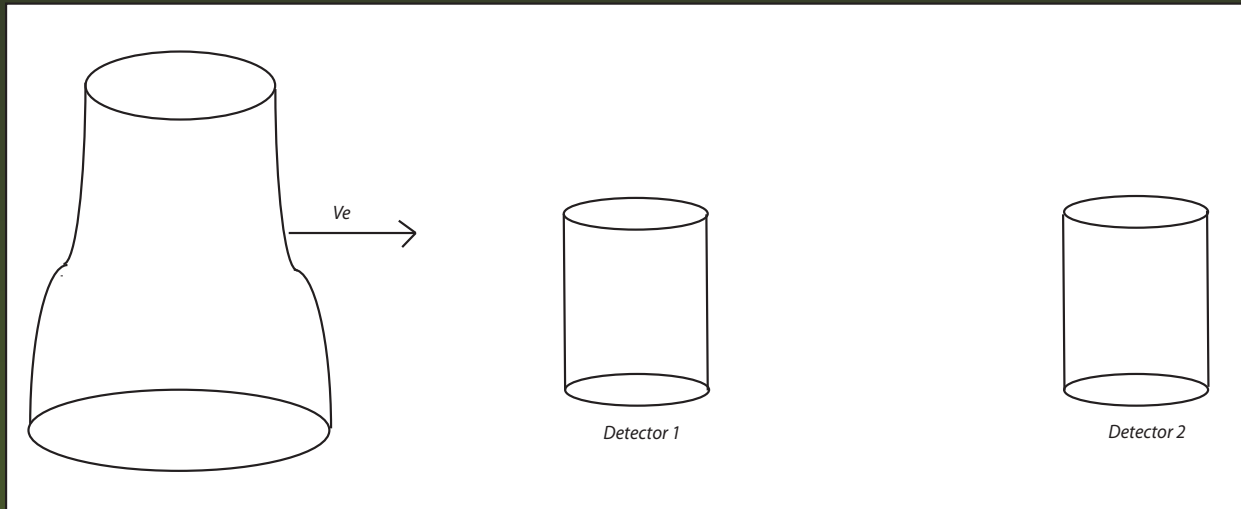
# Finding $\theta_{13}$

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- Determining a value of  $\theta_{13}$  will complete our knowledge of mixing
- The other two mixing angles are measured by solar neutrinos ( $\theta_{12}$ ) and atmospheric neutrinos ( $\theta_{23}$ )
- There are two major types of experiments being done in the attempt to find  $\theta_{13}$ : accelerator experiments and reactor experiments.
- I'll focus on two of the reactor experiments: Daya Bay (China) and Double Chooz (France)

# Reactor Experiments

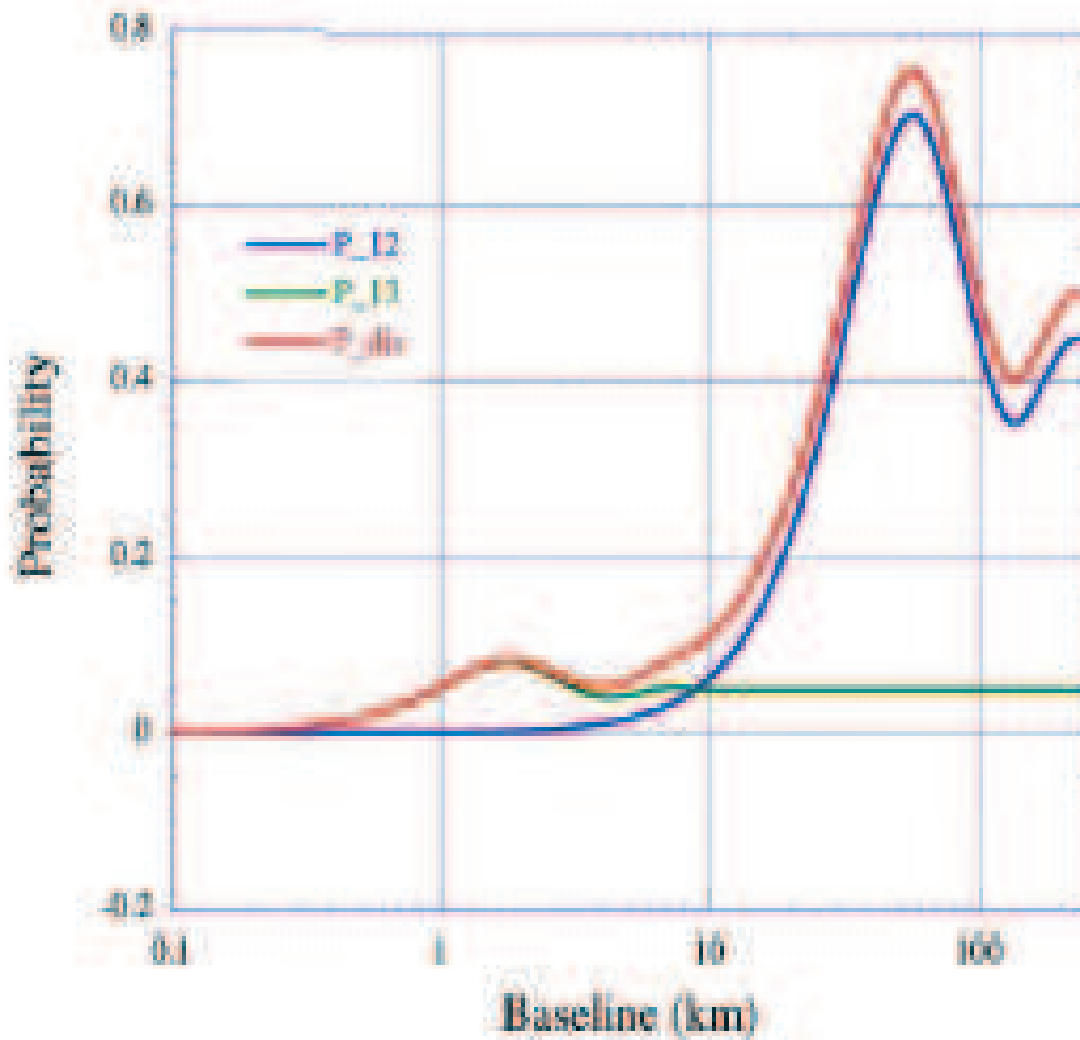
- In reactor experiments, physicists look for a change in flux of electron antineutrinos over some distance.



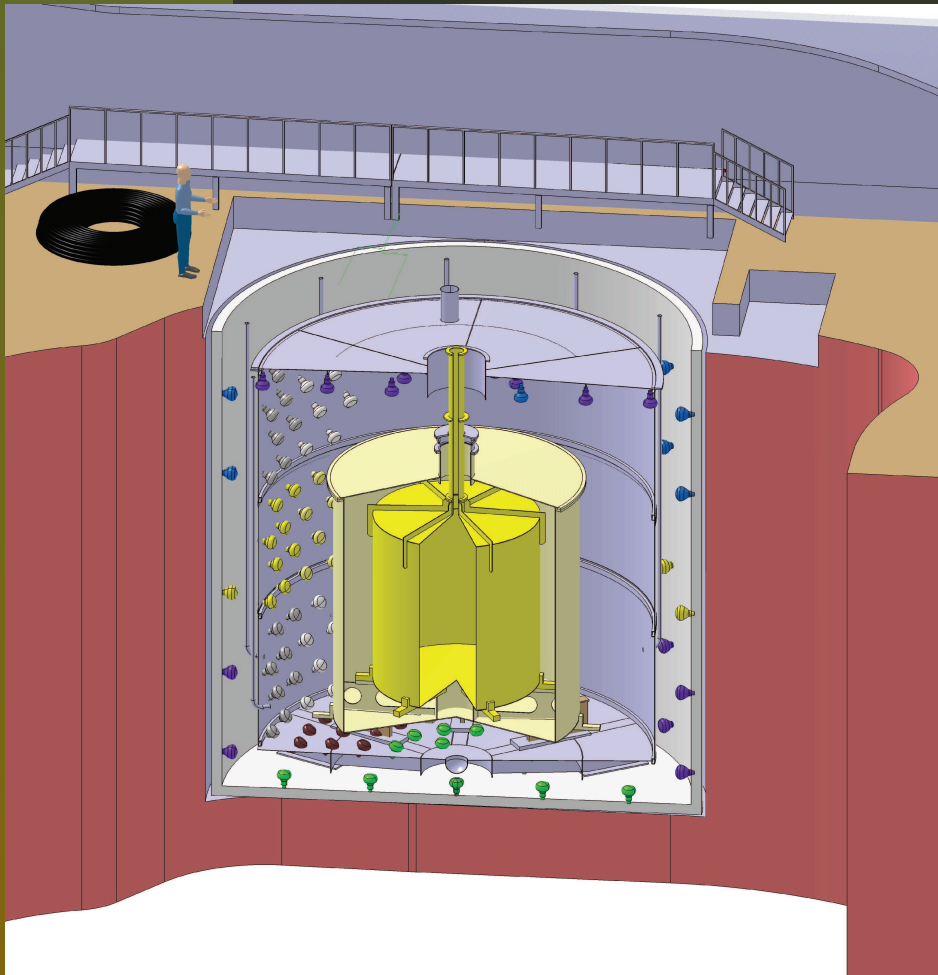
- The survival probability of reactor antineutrinos is described approximately by

$$P_{ee} \simeq 1 - \sin^2 2\theta_{13} \sin^2 \frac{\Delta m_{13}^2 L}{4E_\nu} + \left( \frac{\Delta m_{21}^2 L}{4E_\nu} \right) \cos^4 \theta_{13} \sin^2 2\theta_{12}$$

# Reactor Experiments



# Detecting Neutrinos



- Detectors are filled with Gd-doped fluid that “detects” reverse  $\beta$ -decay:  
$$\bar{\nu}_e + p \rightarrow n + e^+$$
- Researches look for distinctive  $\gamma$ -ray “double-burst” from  $e^- + e^+$  and  $n$  capture

# Daya Bay

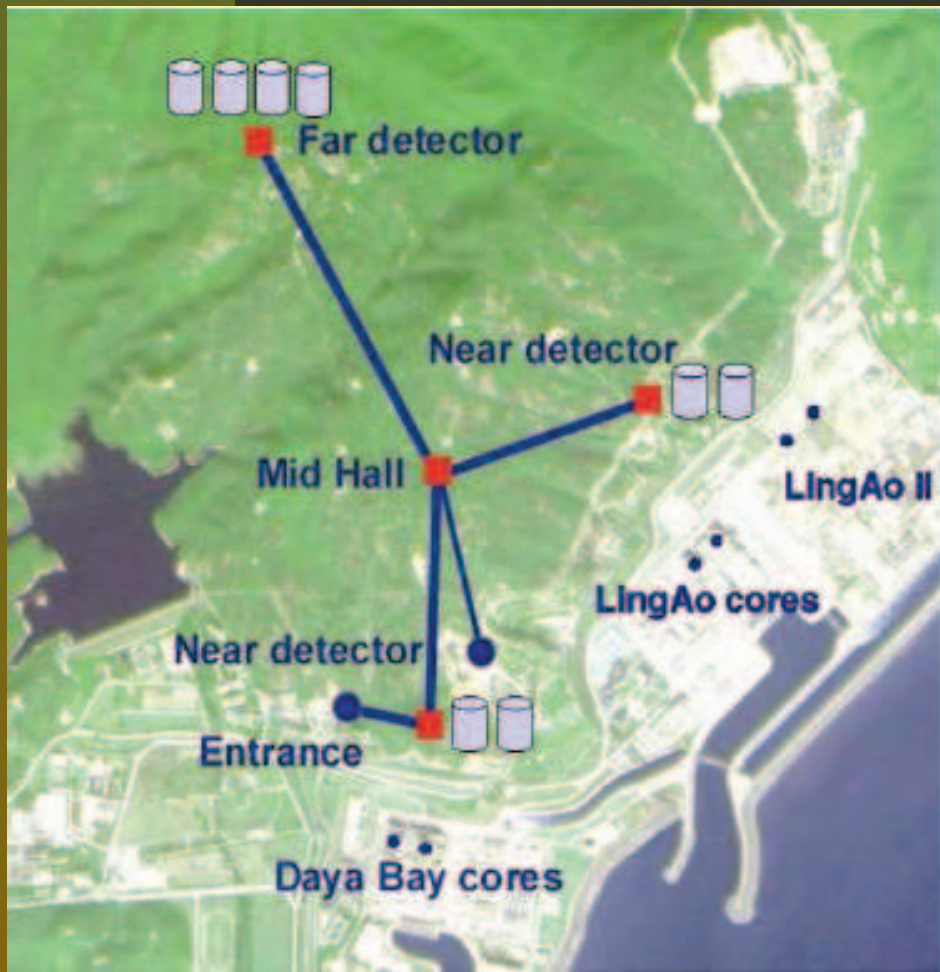


<http://dayabay.bnl.gov/>

- Being constructed in China, outside of Hong Kong; data collection starts in June 2010.
- Will detect antineutrinos from two (eventually three) nearby plants: Daya Bay and LingAo (total power 11.6 GW), with a third plant (with an additional 5.8 GW power) being commissioned by 2010.



# Daya Bay



Daya Bay Proposal, Daya Bay Collaboration

- Three detector sites with a total of 8 identical detectors
- After 3 years of data collection, experimenters hope to have a value of  $\sin^2 2\theta_{13}$  with a sensitivity of 0.008

# Double Chooz



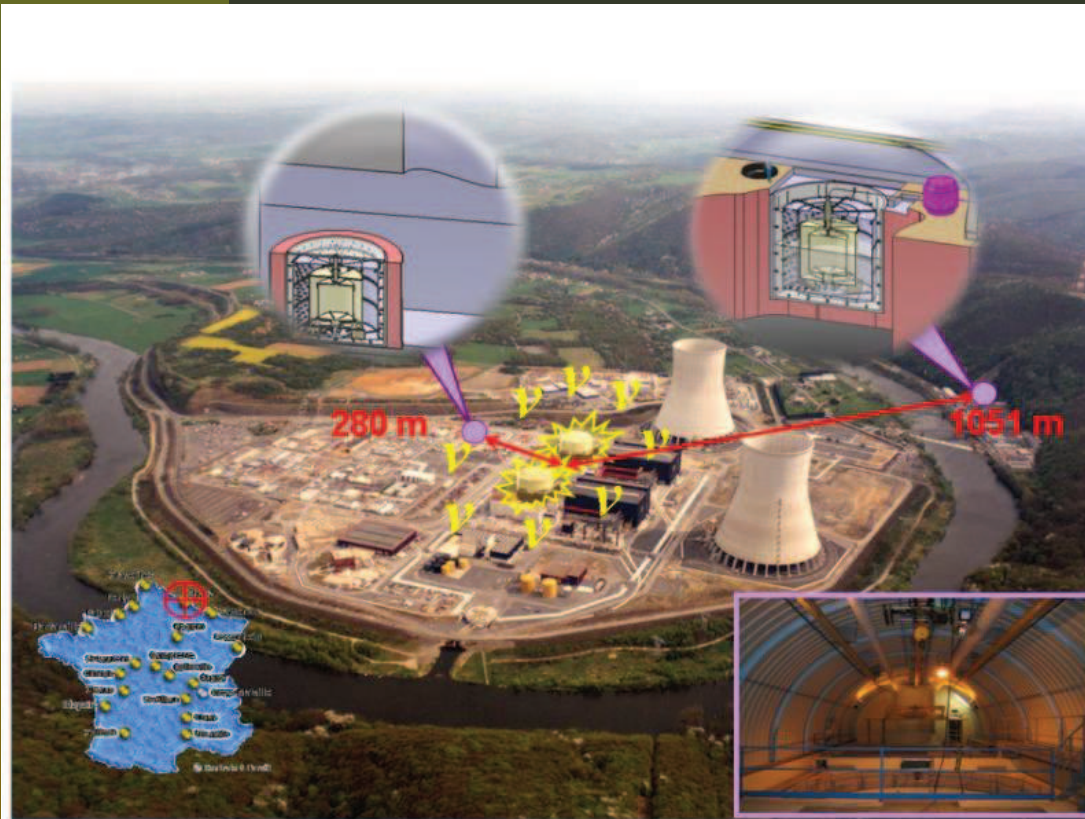
[http://doublechooz.in2p3.fr/Scientific/Photos/vue\\_laterale\\_Centrale\\_Chooz.jpg](http://doublechooz.in2p3.fr/Scientific/Photos/vue_laterale_Centrale_Chooz.jpg)

- Detecting antineutrinos from a pair of reactor cores at Chooz nuclear power plant, in France.

- Total output of power plant is  $\sim 8.6$  MW.

# Double Chooz

- Two sites with near-identical detectors
- After three years of data collection, experimenters expect to have a value of  $\sin^2 2\theta_{13}$  with a sensitivity of 0.025.



[http://irfu.cea.fr/Spp/en/Phocea/Vie\\_des\\_labos/Ast/alltec.php?id\\_ast=2260](http://irfu.cea.fr/Spp/en/Phocea/Vie_des_labos/Ast/alltec.php?id_ast=2260)

# $\theta_{13}$ and CP violation

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} c_{13}c_{12} & c_{13}s_{12} & s_{13}e^{-i\delta} \\ -c_{23}s_{12} - s_{13}s_{23}c_{12}e^{i\delta} & c_{23}c_{12} - s_{13}s_{23}s_{12}e^{i\delta} & c_{13}s_{23} \\ s_{23}s_{12} - s_{13}c_{23}c_{12}e^{i\delta} & -s_{23}c_{12} - s_{13}c_{23}s_{12}e^{i\delta} & c_{13}c_{23} \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

- A non-vanishing value of  $\theta_{13}$  would imply that neutrino oscillation is CP violating, which would be the first evidence of leptonic CP violation.
- If the reactor experiments put an upper bound of  $< 0.01$  on  $\theta_{13}$ , then CP violation in neutrinos would be essentially insignificant.
- Even if neutrino oscillations are found to be CP violating, it's possible that it's not the CP violation for which we are seeking.

# Neutrino mass and the evolution of the universe

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- “What are the masses of the neutrinos and how have they shaped the evolution of the universe?”
- The universe has an apparent “overabundance” of matter, and “deficit” of antimatter.
- It is known that CP violation in quarks contributes to this matter/antimatter imbalance, but it is also known that their contribution is not large enough to explain the imbalance
- If neutrino oscillations are significantly CP violating, we have another candidate to help explain the imbalance.