



# Revisiting the surface-energy-flux perspective on the sensitivity of global precipitation to climate change

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## Abstract

Climate models simulate an increase in global precipitation at a rate of approximately 1–3% per Kelvin of global surface warming. This change is often interpreted through the lens of the atmospheric energy budget, in which the increase in global precipitation is mostly offset by an increase in net radiative cooling. Other studies have provided different interpretations from the perspective of the surface, where evaporation represents the turbulent transfer of latent heat to the atmosphere. Expanding on this surface perspective, here we derive a version of the Penman–Monteith equation that allows the change in ocean evaporation to be partitioned into a thermodynamic response to surface warming, and additional diagnostic contributions from changes in surface radiation, ocean heat uptake, and boundary-layer dynamics/relative humidity. In this framework, temperature is found to be the primary control on the rate of increase in global precipitation within model simulations of greenhouse gas warming, while the contributions from changes in surface radiation and ocean heat uptake are found to be secondary. The temperature contribution also dominates the spatial pattern of global evaporation change, leading to the largest fractional increases at high latitudes. In the surface energy budget, the thermodynamic increase in evaporation comes at the expense of the sensible heat flux, while radiative changes cause the sensible heat flux to increase. These tendencies on the sensible heat flux partly offset each other, resulting in a relatively small change in the global mean, and contributing to an impression that global precipitation is radiatively constrained.

**Keywords** Hydrologic cycle · Global warming

## 1 Introduction

In Earth’s hydrologic cycle, water evaporates from the surface, condenses in the atmosphere, and returns to the surface as precipitation. It takes an average water molecule 8–10 days to complete this cycle (e.g., Van Der Ent and Tuinenburg 2017). On timescales much longer than this, the rates

of global-mean evaporation and precipitation are essentially equal. Because of the latent heat absorbed and released during evaporation and condensation, the hydrologic cycle plays an important role in the heat engine of the climate system, transferring energy from the warm surface, where most sunlight is absorbed, to the cooler atmosphere, where infrared radiation is emitted back to space.

In response to CO<sub>2</sub>-induced warming, climate models predict that the intensity of the global hydrologic cycle (i.e., global-mean evaporation/precipitation) will increase by around 1–3%/K, which is significantly less than the approximately 7%/K increase in atmospheric water vapor resulting from the Clausius–Clapeyron equation and near-constant relative humidity (e.g., Boer 1993; Allen and Ingram 2002). This disparity has been invoked to explain important aspects of the climate response to CO<sub>2</sub>-induced warming, including a slowing down of the atmospheric circulation (Held and Soden 2006), and an increase in the frequency and intensity of floods and droughts (Allen and Ingram 2002; Trenberth 1999, 2011). Understandably, therefore, the relationship

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between surface temperature and the global hydrologic cycle has been a central topic in climate science for decades (e.g., Manabe and Wetherald 1975)

Many previous studies have investigated the change in global precipitation with warming from the perspective of the global-mean atmospheric energy budget,

$$\overline{LP} = \overline{R_a} - \overline{H}, \quad (1)$$

where  $L$  is the latent heat of vaporization,  $P$  is the rate of precipitation,  $R_a$  is the net heat lost through radiation,  $H$  is the sensible heat flux from the surface to the atmosphere, and  $(\overline{\phantom{x}})$  indicates the global mean of a quantity (e.g., Allen and Ingram 2002; Stephens and Ellis 2008; Previdi 2010; Pendergrass and Hartmann 2014; Fläschner et al 2016). Equation (1) implies that any increase in  $\overline{LP}$  must be offset by a decrease in  $\overline{H}$  and/or an increase in  $\overline{R_a}$ . In most simulations of CO<sub>2</sub>-induced warming by GCMs, the change in  $\overline{H}$  tends to be small compared with changes in  $\overline{LP}$  and  $\overline{R_a}$ , with the latter dominated by an increase in longwave emissions (e.g., Lambert and Webb 2008; Andrews et al 2009; Pendergrass and Hartmann 2014). This result has sometimes been interpreted as evidence that the change in global precipitation is primarily determined by the ability of the atmosphere to radiate more energy as it warms (e.g., Allen and Ingram 2002). Yet in most cases, there is not a one-to-one trade-off between changes in  $\overline{R_a}$  and  $\overline{LP}$ , because changes in  $\overline{H}$  are often not negligible. For example, among GCMs participating in the most recent Coupled Model Intercomparison Project (CMIP5), Pendergrass and Hartmann (2014) found that the change in  $\overline{H}$  was small but significant, with a magnitude about a third as large as the change in  $\overline{LP}$ . Simulations spanning a wider range of climate states have also been found to exhibit large variability in the change in  $\overline{H}$  with global temperature (O’Gorman and Schneider 2008; O’Gorman et al 2012). On its own, however, the atmospheric energy budget provides limited insight into how the changes in  $\overline{H}$ ,  $\overline{LP}$ , and  $\overline{R_a}$  are partitioned.

An alternative to the atmospheric energy budget perspective is to treat the hydrologic cycle as a turbulence-driven process, in which evaporation,  $E$ , represents the turbulent flux of water vapor from the surface to the near-surface atmosphere (e.g., Penman 1948; Priestly and Taylor 1972; Monteith 1981; Boer 1993; Pierrehumbert 2002; Richter and Xie 2008; Lorenz et al 2010; Pierrehumbert 2010). Over the oceans, where 85% of global evaporation occurs (e.g., Trenberth et al 2007), the rate of evaporation is given approximately by the bulk transfer equation,

$$E = [q^*(T_s) - rq^*(T_a)]\rho C_H u \quad (2)$$

where  $q^*$  is the specific humidity at saturation,  $T_s$  and  $T_a$  are the temperatures of the ocean surface and near-surface atmosphere,  $r$  is the near-surface relative humidity,  $\rho$  is the near-surface air density,  $C_H$  is the bulk transfer coefficient,

and  $u$  is the near-surface wind speed. Similarly, the sensible heat flux is approximately given by

$$H = c_p(T_s - T_a)\rho C_H u, \quad (3)$$

where  $c_p$  is the specific heat capacity of air, and  $C_H$  is typically assumed to have the same value as in Eq. (2). Equations (2) and (3) are derived from Monin–Obukhov similarity theory (e.g., Pierrehumbert 2010), and they imply that the fluxes of latent and sensible heat are determined by the speed, temperature, and relative humidity of the near-surface winds, and by the difference in temperature between the ocean surface and the near-surface atmosphere.<sup>1</sup> While these variables are all strongly dependent on atmospheric physics, collectively they must satisfy the surface energy budget,

$$LE + H = R_s - G, \quad (4)$$

where  $L$ ,  $E$ , and  $H$  are defined as in Eq. (1),  $R_s$  is the net downward radiation flux at the surface, and  $G$  is the rate of heat storage by the ocean plus—at high latitudes—the heat required to melt frozen hydrometeors that reach the surface. Unlike Eq. (1), Eqs. (2)–(4) are valid at a specific location, and not just in the global mean.

With three equations instead of one, the turbulent-flux perspective provides insight into the partitioning between  $H$  and  $LE$  that is not possible with the atmospheric energy budget alone. For example, Pierrehumbert (2002) has shown that one implication of Eqs. (2)–(4) is that  $LE \ll H$  at very cold temperatures due to low values of  $q^*$ , while  $LE \gg H$  at very warm temperatures due to high values of  $q^*$ , and correspondingly weak (or even negative) values of  $T_s - T_a$ , which are necessary to satisfy the surface energy budget (e.g., Le Hir et al 2009). Interpolating between these two extremes, it must be true that, in more moderate climates like that of the present day, warming will tend to cause  $LE$  to increase (via an increase in  $q^*$ ), and  $H$  to decrease (via a decrease in  $T_s - T_a$ ). However, while such changes are easily diagnosed in GCM simulations (e.g., Richter and Xie 2008; Lorenz et al 2010), they are difficult to quantify from first principles, since they also depend on changes in net surface radiation, ocean heat uptake, and boundary-layer dynamics, which are only weakly constrained by Eqs. (2)–(4).

As we argue in this paper, however, the turbulent-flux perspective may have more explanatory power than previously assumed. By combining Eqs. (2)–(4) into a variant of the Penman–Monteith equation (Penman 1948; Monteith 1981), we show that the change in evaporation can be partitioned into four distinct terms, each with a straightforward physical interpretation. The first term depends only on surface temperature, and represents the thermodynamic response to warming,

<sup>1</sup> On long timescales,  $T_s - T_a$  is generally positive over the ocean, implying a transfer of both sensible and latent heat from the surface to the atmosphere.

which is independent of other changes in the climate system. Meanwhile, the other terms represent the change in evaporation due to changes in surface radiation, ocean heat uptake, and boundary-layer dynamics/relative humidity. These terms cannot be derived from first principles, and must therefore be diagnosed from GCM simulations. Nevertheless, if the spatial patterns of changes in surface radiation and ocean heat uptake are known, the Penman–Monteith equation provides a way to quantify the resulting spatial pattern of the change in evaporation. When extended to the global mean, the Penman–Monteith framework also sheds light on the partitioning between changes in  $\bar{H}$  and  $\bar{LE}$  in the surface energy budget (or  $\bar{LP}$  in the atmospheric energy budget).

The paper is organized as follows. We begin in Sect. 2 by deriving a variant of the Penman–Monteith equation, which governs the rate of evaporation over the oceans subject to energetic constraints. In Sect. 3, we use this equation to diagnose the contributions to evaporation change in GCM simulations of global warming. In this diagnostic framework, we find that most of the increase in global evaporation is a direct consequence of warmer temperatures, while changes in surface radiation and ocean heat uptake play a secondary role. Based on these results, we then derive an approximation for evaporation change as a function of changes in temperature, surface radiation, and ocean heat uptake, and show that it accurately represents the fast and slow responses of evaporation to  $\text{CO}_2$  forcing, while also shedding light on the partitioning between changes in the latent and sensible heat flux. In Sect. 4, we apply the Penman–Monteith framework to a series of idealized simulations run by O’Gorman and Schneider (2008), and find that thermodynamics alone can account for much of the changes in global precipitation among the simulations. We conclude with a brief summary and discussion in Sect. 5.

## 2 Derivation and interpretation of the Penman–Monteith equation

We build upon a long history of research on the physics of ocean evaporation, beginning with the fundamental equations (2)–(4) that govern the exchange of energy between the ocean surface and the atmosphere (Penman 1948; Monteith 1981; Pierrehumbert 2002). In particular, we seek to solve this system of equations for  $E$  while eliminating  $T_s - T_a$  and  $H$ . To do so, we first approximate  $q^*(T)$  as the first-order Taylor expansion about the point  $T = T_a$ :

$$q^*(T) \approx q^*(T_a) + \frac{dq^*}{dT}[T - T_a]. \quad (5)$$

Given the Clausius–Clapeyron relation,

$$\frac{dq^*}{dT} = \alpha q^*, \quad (6)$$

Equation (5) implies that

$$q^*(T_s) \approx q^*(T_a)(1 + \alpha[T_s - T_a]), \quad (7)$$

where

$$\alpha = \frac{L}{R_v T_a^2} \quad (8)$$

is the Clausius–Clapeyron scaling factor, with  $R_v$  representing the specific gas constant for water vapor. Given Eq. (7), we can express the air–sea moisture difference in Eq. (2) as a function of  $r$ ,  $T_a$ , and  $T_s - T_a$ :

$$E \approx q^*(T_a)(1 - r + \alpha[T_s - T_a])\rho C_H u. \quad (9)$$

This allows us to eliminate  $T_s - T_a$  and  $H$  from Eqs. (2)–(4) to arrive at the following expression for the rate of evaporation from the ocean surface:

$$LE \approx \eta(R_s - G + \kappa), \quad (10)$$

where  $\eta$  and  $\kappa$  are defined as

$$\eta \equiv \frac{1}{1 + \beta_0}, \quad (11)$$

$$\kappa \equiv (1 - r)c_p \rho C_H u \alpha^{-1}, \quad (12)$$

with

$$\beta_0 = \frac{c_p}{\alpha L q^*(T_a)} \quad (13)$$

representing the Bowen ratio ( $H / LE$ ) in the limit of 100% near-surface relative humidity ( $r = 1$ ).

Equation (10) is equivalent to the Penman–Monteith equation for terrestrial evapotranspiration from a saturated surface (i.e., where stomatal resistance is zero) (e.g., Scheff and Frierson 2014). As written above, however, Eq. (10) is much easier to interpret than its more conventional form: besides  $R_s$  and  $G$ , it comprises just two terms ( $\eta$  and  $\kappa$ ), each of which has a straightforward physical meaning.

The first term,  $\eta$ , is a function of  $T_a$  alone. In the context of climate change, therefore,  $\eta$  captures the thermodynamic response of evaporation to warming, which is independent of changes in other variables like wind speed, relative humidity, and net surface radiation. On the other hand,  $\eta$  does implicitly involve  $T_s - T_a$  and  $H$ , as the following thought experiment illustrates.

Let us assume that  $T_s$  and  $T_a$  were to increase by the same amount, while all other variables in Eqs. (2)–(4) were held constant. In this scenario,  $H$  would not change, while  $LE$  would increase with  $T_a$  at the same rate as  $q^*(T_a)$ —i.e., at the Clausius–Clapeyron rate of  $\alpha \approx 7\%/K$  (Eq. 9; Richter and Xie 2008). But without compensating changes in  $R_s - G$ , such an increase in  $LE$  would clearly violate the surface energy budget. In reality, therefore,  $T_s - T_a$  must decrease

with warming, such that  $H$  decreases and  $LE$  increases at a rate less than 7%/K. This constraint on  $T_s - T_a$  has been described in similar terms by Pierrehumbert (2002) and Lorenz et al (2010), but in Eq. (10), it is implicitly captured by a single variable,  $\eta$ .

Meanwhile,  $\kappa$  represents the dependence of evaporation on relative humidity ( $r$ ) and boundary-layer dynamics ( $u$  and  $C_H$ ), while its temperature-dependence (via  $\alpha$  in its denominator) is negligibly small. An increase in  $r$  will reduce the air-sea moisture difference, thereby causing  $LE$  to decrease. Yet in the absence of changes in  $R_s - G$ , this decrease must be accompanied by an increase in  $T_s - T_a$  to ensure no change in  $LE + H$ . Likewise, an increase in  $u$  or  $C_H$  will cause both  $LE$  and  $H$  to increase, and thus requires a simultaneous decrease in  $T_s - T_a$  to satisfy the surface energy budget.

These thought experiments illustrate the crucial role played by  $T_s - T_a$  within the Penman–Monteith framework (Eqs. 2–4). In response to any environmental change that, in isolation, would violate the surface energy budget,  $T_s - T_a$  must adjust to conserve energy and, in the process, alter the partitioning between  $LE$  and  $H$ . Physically, this framework is consistent with previous studies that have emphasized the role of  $T_s - T_a$  in limiting the rate of increase in global precipitation with surface warming (Pierrehumbert 2002; Richter and Xie 2008; Takahashi 2009; Lorenz et al 2010). Changes in the longwave opacity profile could in turn provide a constraint for  $T_s - T_a$  (Takahashi 2009). However, the advantage of the Penman–Monteith surface-energy perspective developed here is that it allows the change in evaporation with surface warming to be partitioned into a predictable component due to changes in  $\eta$ , and a diagnostic component due to changes in surface radiation, ocean heat uptake, and relative humidity/dynamics. In the next section, we show that this partitioning provides insight into the factors controlling the change in global and regional evaporation in response to CO<sub>2</sub>-induced warming.

### 3 Results from comprehensive GCMs

In this section, we use the Penman–Monteith framework (Eq. 10) to quantify the factors contributing to the change in ocean evaporation in GCM simulations of CO<sub>2</sub>-induced warming. We examine both the equilibrium response to an abrupt doubling of atmospheric CO<sub>2</sub>, as simulated by an ensemble of atmosphere-only GCMs with slab oceans, and the transient response a century after an abrupt quadrupling of atmospheric CO<sub>2</sub>, as simulated by an ensemble of coupled atmosphere-ocean GCMs. The transient simulations were performed as part of the most recent Coupled Model Intercomparison Project (CMIP5; Taylor et al 2012), while the equilibrium simulations were performed with a previous

generation of climate models included in CMIP3<sup>2</sup> (Meehl et al 2007). In the equilibrium (slab ocean) case, we focus on the change in ocean evaporation between the last five years of the pre-industrial control simulations and years 21–25 after CO<sub>2</sub> doubling. The transient results reflect the change in evaporation between the pre-industrial control simulations and years 96–100 after CO<sub>2</sub> quadrupling.

Tables 1 and 2 list the names of the 10 CMIP3 models and 12 CMIP5 models included in our analysis. Some models were excluded because the required variables were not readily available, or in the CMIP3 case, because the simulations were not in radiative equilibrium. For consistency with our later analysis in Sect. 3.1, we further restrict the CMIP5 models to those that also performed CO<sub>2</sub> quadrupling simulations with prescribed climatological sea-surface temperatures (SSTs). The second column of each table gives the percent rate of change in global evaporation per Kelvin of global warming, while the third column gives the equivalent rate over the oceans (i.e., using ocean-mean instead of global-mean values of  $\Delta E$ ,  $E$ , and  $\Delta T$ ). Even though the rate of change in ocean evaporation exceeds the global change in all models, the two values are highly correlated across models ( $r = 0.93$  and  $0.81$  for CMIP3 and CMIP5, respectively). This high correlation is not surprising, considering that the oceans account for 85% of global evaporation in the current climate (Trenberth et al 2007), and an even larger share of the increase in evaporation under global warming (e.g., Fu and Feng 2014). To first order, therefore, we can understand the change in global evaporation (and thus precipitation) by focusing on the ocean-mean changes.

Figure 1 shows the fractional change in ocean evaporation per Kelvin of local warming in the ensemble mean of the equilibrium (left column) and transient (right column) simulations, along with the individual contributions to evaporation change due to changes in  $\eta$ ,  $R_s$ ,  $G$ , and  $\kappa$  (rows 2–5). We have calculated these contributions using the discrete form of the fractional derivative of Eq. (10) with respect to  $T_a$ :

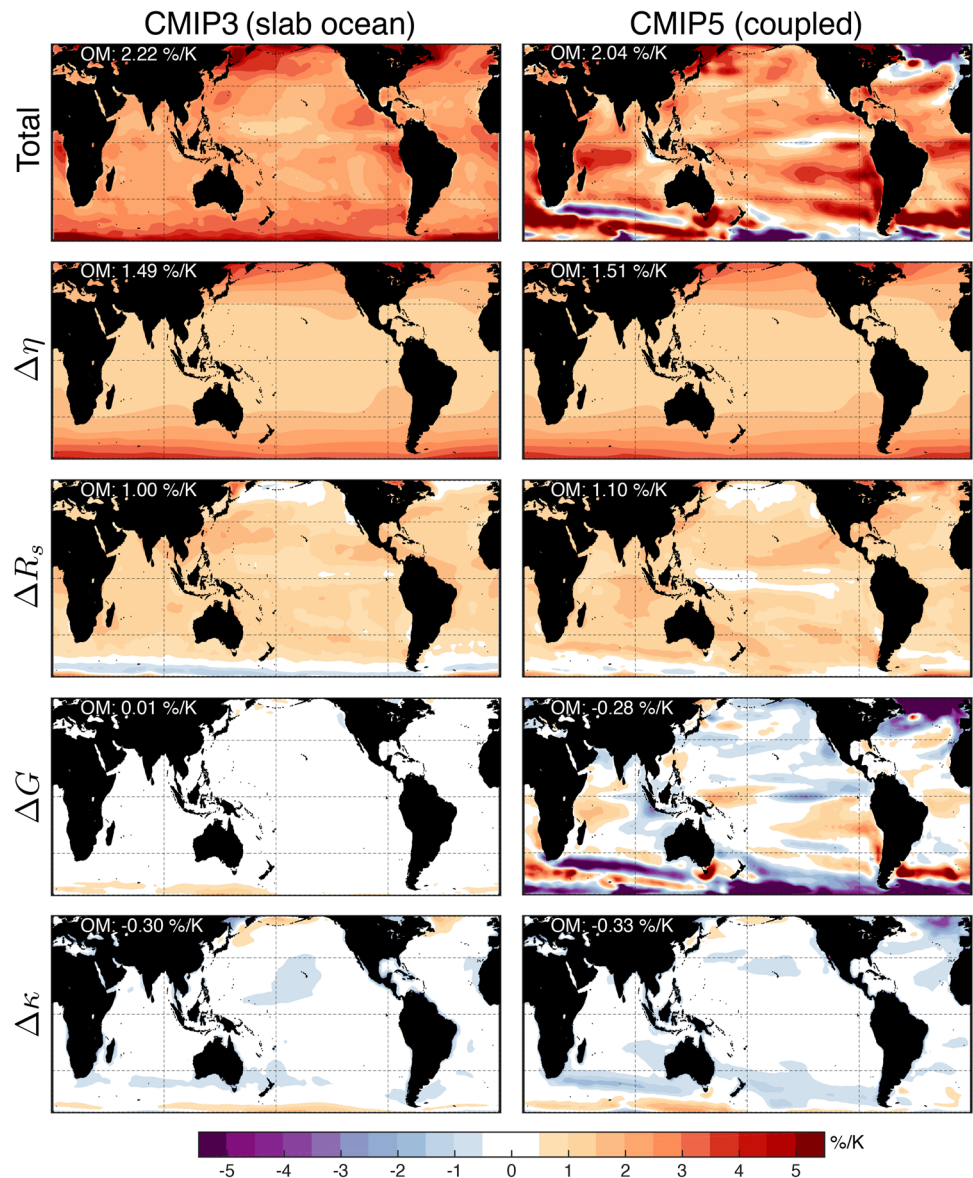
$$\frac{1}{\Delta T_a} \left[ \underbrace{\frac{\Delta E}{E}}_1 \approx \underbrace{\frac{\Delta \eta}{\eta}}_2 + \frac{\eta}{LE} \left( \underbrace{\Delta R_s}_3 - \underbrace{\Delta G}_4 + \underbrace{\Delta \kappa}_5 \right) \right], \quad (14)$$

where  $\Delta$  indicates the change in a variable between the control and warmed climate. The number above each term in Eq. (14) indicates its order, from top to bottom, in Fig. 1. Each contribution was calculated using monthly-mean model output (see “Appendix” for details). The ocean-mean of each contribution is given in the top left corner of each panel.

<sup>2</sup> Equilibrium simulations were not part of CMIP5.



**Fig. 1** The percent change in evaporation over the oceans in response to CO<sub>2</sub>-induced warming (top row), and the individual contributions from changes in  $\eta$  (second row),  $R_s$  (third row),  $G$  (fourth row), and  $\kappa$  (fifth row), in the equilibrium (left column) and transient (right column) simulations. Each contribution was calculated from ensemble-mean output according to Eq. (14) (see “Appendix”), and represents the change per Kelvin of global warming over the oceans, which is equal to 3.15 K in the 2 × CO<sub>2</sub> equilibrium simulations, and 4.13 K in the 4 × CO<sub>2</sub> transient simulations. The top left corner of each panel gives the ocean-mean (OM) values of each contribution. The results are broadly similar between the equilibrium and transient ensembles, with the exception of the contribution from ocean-heat uptake ( $\Delta G$ ), which is negligible in the equilibrium simulations due to the absence of a dynamical ocean



Columns 4–7 of Tables 1 and 2 give the equivalent values for each ensemble member, along with the standard deviation of each contribution across the ensemble. In this Section we discuss only the ensemble-mean results (Fig. 1), but will address the inter-model variability in Sect. 5.

In both the equilibrium and transient cases, we find that  $\Delta\eta$  accounts for the largest contribution to evaporation change, increasing from about 1.2%/K at the equator to more than 4%/K at higher latitudes. To understand this meridional structure—which is also reflected in the pattern of total evaporation change—consider the analytic expression for the  $\Delta\eta$  contribution, which can be derived from Eqs. (6), (8), (11), and (13):

$$\frac{d \ln \eta}{dT_a} = \frac{\beta_0}{1 + \beta_0} \left( \alpha - \frac{2}{T_a} \right). \quad (15)$$

Because  $\beta_0$  decreases almost exponentially with increasing temperature,  $d \ln \eta / dT_a$  is largest at cold temperatures where it approaches  $(\alpha - 2/T_a)$ —not much less than the Clausius–Clapeyron scaling of atmospheric water vapor under fixed relative humidity. As  $T_a$  increases, however, the gap between  $\alpha$  and  $d \ln \eta / dT_a$  grows ever larger. The increase in  $d \ln \eta / dT_a$  with latitude therefore reflects the meridional structure of  $T_a$  in the control climate. In the global mean,  $\Delta\eta$  causes an increase in evaporation of about 1.5% per Kelvin of global warming in both the equilibrium and transient cases, accounting for the largest fraction of the total increase.<sup>3</sup>

<sup>3</sup> The global-mean contributions represent the average of the fractional changes (Fig. 1), weighted by the product of mean-state evaporation and local temperature change. Further details are provided in the “Appendix”.

**Table 1** First column: the names of the CMIP3 (slab-ocean) models included in our analysis

CMIP3 model	$\frac{\Delta E}{E\Delta T}$ (globe)	$\frac{\Delta E}{E\Delta T}$ (oceans)	$\Delta\eta$	$\Delta R_s$	$\Delta G$	$\Delta\kappa$
Can-CGCM3.1 (T47)	2.02	2.14	1.37	0.96	− 0.01	− 0.22
Can-CGCM3.1 (T63)	2.18	2.28	1.37	0.95	0.07	− 0.16
CSIRO-Mk3.0	2.26	2.67	1.56	1.20	− 0.01	− 0.13
GFDL-CM2	1.35	1.76	1.53	0.59	0.06	− 0.43
HadGEM1	1.81	2.18	1.48	0.89	0.09	− 0.34
INM-CM3	1.58	1.85	1.56	0.71	− 0.06	− 0.37
MIROC3.2 (hires)	1.92	2.13	1.47	1.05	0.02	− 0.46
MIROC3.2 (medres)	2.18	2.37	1.51	1.23	− 0.01	− 0.39
MPI-OM	1.95	2.23	1.46	1.07	− 0.10	− 0.23
MRI-CGCM2	2.28	2.48	1.52	1.16	0.01	− 0.24
Ensemble mean	1.97	2.22	1.49	1.00	0.01	− 0.30
Standard deviation	0.30	0.27	0.07	0.21	0.06	0.12

Second column: the rate of increase (in %/K) in global-mean evaporation in response to a doubling of atmospheric CO<sub>2</sub>, after reaching radiative equilibrium. Third column: the rate of increase in ocean-mean evaporation. Columns 4–7: the individual contributions to changes in ocean-mean evaporation from  $\Delta\eta$ ,  $\Delta R_s$ ,  $\Delta G$ , and  $\Delta\kappa$ , according to Eq. (14). The second row from the bottom gives the ensemble-mean rates (Fig. 1). These were calculated from ensemble-mean variables, and therefore differ slightly from the average rates of the individual models. The bottom row gives the standard deviation across models. In the second (third) column, global-mean (ocean-mean) rates were calculated using the global-mean (ocean-mean) values of  $\Delta E$ ,  $E$ , and  $\Delta T$ . Global-mean and ocean-mean rates are highly correlated at  $r = 0.93$ , indicating the dominant influence of the ocean on global evaporation

**Table 2** As in Table 1, but for the CMIP5 (coupled) simulations

CMIP5 model	$\frac{\Delta E}{E\Delta T}$ (globe)	$\frac{\Delta E}{E\Delta T}$ (oceans)	$\Delta\eta$	$\Delta R_s$	$\Delta G$	$\Delta\kappa$
BCC-CSM1.1	1.65	1.98	1.41	1.00	− 0.21	− 0.26
CanESM2	1.41	1.85	1.44	0.74	− 0.19	− 0.17
CCSM4	1.46	1.99	1.42	1.07	− 0.22	− 0.34
CSIRO-Mk3.6.0	1.85	2.29	1.53	1.29	− 0.27	− 0.35
HadGEM2-ES	1.23	1.78	1.43	0.92	− 0.22	− 0.42
INM-CM4	1.35	1.58	1.54	0.79	− 0.46	− 0.31
IPSL-CM5A-LR	2.02	2.64	1.71	1.53	− 0.27	− 0.38
MIROC5	1.46	1.64	1.46	1.02	− 0.34	− 0.59
MPI-ESM-LR	1.63	2.19	1.52	1.20	− 0.34	− 0.24
MPI-ESM-MR	1.77	2.26	1.49	1.21	− 0.27	− 0.23
MRI-CGCM3	2.27	2.27	1.43	1.28	− 0.29	− 0.23
NorESM1-M	1.43	1.96	1.52	1.12	− 0.32	− 0.40
Ensemble mean	1.62	2.04	1.51	1.10	− 0.28	− 0.33
Standard deviation	0.30	0.31	0.08	0.22	0.08	0.11

The rates of change in evaporation are based on years 96–100 after CO<sub>2</sub> quadrupling. The global-mean and ocean-mean rates are correlated at  $r = 0.81$

Compared with  $\Delta\eta$ , the contributions from the other terms in Eq. (14) are generally smaller in magnitude.  $\Delta R_s$  accounts for the second largest contribution to evaporation change (Fig. 1, row 3), increasing ocean-mean evaporation by about 1%/K in both the equilibrium and transient cases due to an increase in net surface radiation in the tropics and

midlatitudes.<sup>4</sup> In contrast, the contribution from  $\Delta G$  is negligible in the equilibrium (slab-ocean) simulations, but significantly negative in the transient simulations due to ocean heat uptake in the subpolar North Atlantic Ocean and Southern

<sup>4</sup> The contribution from  $\Delta R_s$  is negative at high latitudes in the Southern Hemisphere, reflecting a decrease in shortwave absorption as a result of increased cloud cover.

Ocean (row 4). Finally, the contribution from  $\Delta\kappa$  tends to be negative but quite weak (row 5), implying that changes in boundary-layer dynamics and relative humidity play a minor role, particularly in determining the spatial pattern of evaporation change.

Figure 1 reveals two key points about how ocean evaporation responds to CO<sub>2</sub>-induced warming in GCMs. First, the largest contribution to evaporation change comes from  $\Delta\eta$ , which is a direct consequence of warmer temperatures (Eq. 15). This means that much of the increase in ocean evaporation with warming—both globally and regionally—is independent of the nature of both the forcing that drove the temperature change (e.g., aerosols vs. greenhouse gases) and how certain physical processes are parameterized within a model (e.g., boundary-layer dynamics, convection, clouds, etc.).

Second, the small contribution from  $\Delta\kappa$  suggests that changes in boundary layer dynamics and relative humidity play a limited role in determining the change in ocean evaporation in response to CO<sub>2</sub>-induced warming, particularly at regional scales. We know of no a priori reason why this should be the case, since in principle the boundary layer could adjust in any number of ways that would alter the surface energy balance (e.g., Pierrehumbert 2002). However, if we take as given that the contribution from  $\Delta\kappa$  is small, the total change in evaporation can then be approximated as

$$L\Delta E \approx LE \frac{d \ln \eta}{dT_a} \Delta T_a + \eta \Delta(R_s - G), \quad (16)$$

where the first term represents the thermodynamic response to warming (Eq. 15), and the second term represents the diagnostic component of evaporation change due to the combined changes in surface radiation and ocean heat uptake.

### 3.1 Testing Eq. (16) on the fast and slow responses of evaporation to CO<sub>2</sub> forcing

Although  $\Delta\kappa$  has little impact on evaporation in Fig. 1, it is reasonable to question the generality of this result. For example, it could be that  $\Delta\kappa$  is not physically independent of the other terms in Eq. (14), in which case its small contribution in Fig. 1 might reflect a cancellation between larger competing effects driven by changes in  $\eta$  and  $R_s - G$ . If true, this would imply that Eq. (16) does not generally hold.

To test the robustness of the approximation in Eq. (16), therefore, we revisit the coupled CMIP5 simulations analyzed previously, but now with the change in evaporation partitioned into “fast” and “slow” components that represent, respectively, the direct response of evaporation to CO<sub>2</sub> quadrupling with fixed SSTs, and the more gradual

changes that occur as the climate warms. Following previous studies (e.g., Lambert and Faull 2007; Lambert and Webb 2008; Lambert et al 2009; Andrews et al 2009, 2010; Andrews and Forster 2010; Frieler et al 2011; Samset et al 2016; Fläschner et al 2016), we define the fast response as the change that occurs when CO<sub>2</sub> is quadrupled and SSTs are fixed at pre-industrial values (i.e., the difference between sstClim4xCO2 and sstClim experiments in CMIP5 parlance), and the slow response as the difference between the coupled 4xCO2 simulations and the fixed-SST 4xCO2 simulations. Defined in this way, the sum of the slow and fast responses gives the total change in the top right panel of Fig. 1.

In the fast response to CO<sub>2</sub> forcing, an increase in the longwave optical depth of the atmosphere will cause an increase in net surface radiation (e.g., Allen and Ingram 2002; McInerney and Moyer 2012) and a decrease in outgoing longwave radiation. The latter effect adds heat to the climate system, which is mostly absorbed by the ocean (thus increasing  $G$ ). Meanwhile, little warming occurs over the oceans because  $T_s$  is held constant, implying that the first term on the RHS of Eq. (16) will be small. As a result, the fast change in evaporation is approximately given by

$$L\Delta E_{\text{fast}} \approx \eta \Delta(R_s - G)_{\text{fast}}, \quad (17)$$

where  $\Delta(R_s - G)_{\text{fast}}$  represents the change in  $R_s - G$  that is directly caused by CO<sub>2</sub> forcing, independent of surface warming.

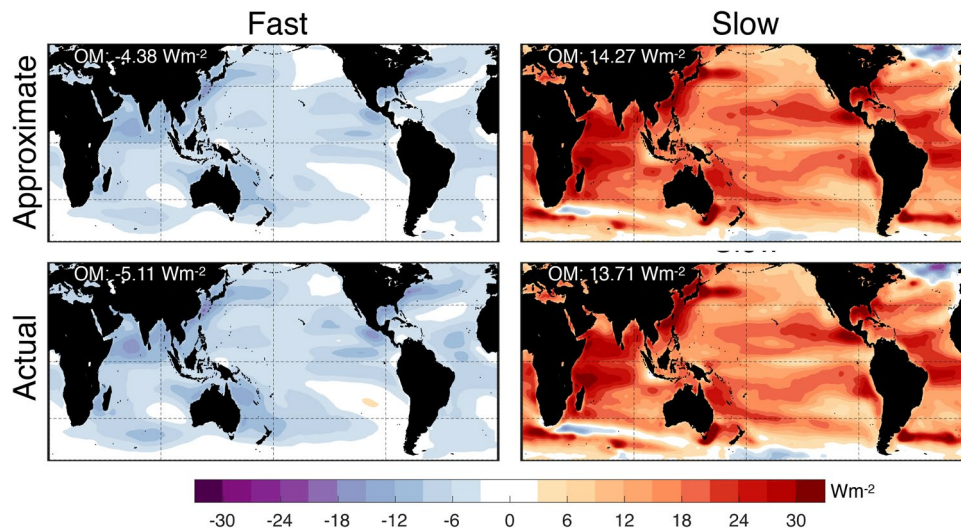
On longer timescales,  $T_a$  will gradually rise, impacting ocean evaporation in two ways. First, there will be a thermodynamic response represented by the first term on the RHS of Eq. (16). Second,  $R_s$  and  $G$  will also change as the atmosphere warms and the earth approaches top-of-atmosphere radiative balance (e.g., Pierrehumbert 1999). Applying the regression method of Gregory et al (2004) at each grid point, we find that the slow change in  $R_s - G$  over most of the global oceans is well approximated in all models as a linear surface-temperature feedback,

$$\Delta(R_s - G)_{\text{slow}} \approx \lambda \Delta T_a, \quad (18)$$

where  $\lambda$  represents the slope of the regression. The slow response of evaporation can then be expressed as

$$L\Delta E_{\text{slow}} \approx \left( LE \frac{d \ln \eta}{dT_a} + \eta \lambda \right) \Delta T_a. \quad (19)$$

Figure 2 shows the approximations in Eqs. (17) and (19) (top row) along with the actual fast and slow changes in evaporation from the ensemble mean of the CMIP5 simulations (bottom row). Comparing the two rows, we find that the approximations capture the spatial pattern of evaporation change remarkably well, with spatial correlations of  $r > 0.99$



**Fig. 2** The “fast” (left column) and “slow” (right column) changes in ocean evaporation in GCM simulations of an abrupt quadrupling of atmospheric  $\text{CO}_2$  (in  $\text{Wm}^{-2}$ ). The fast component represents the direct response of evaporation to  $\text{CO}_2$  quadrupling with fixed SSTs, while the slow component represents the gradual changes

that occur as the climate warms. Top row: the approximate changes calculated from Eqs. (17) and (19), using ensemble-mean values of  $\Delta(R_s - G)_{\text{fast}}$ ,  $\lambda$ ,  $E$ ,  $T_a$ , and  $\Delta T_a$ . Bottom row: the actual changes in the ensemble mean of CMIP5 simulations. Ocean-mean (OM) values are given in the top left corner of each panel

in both cases. The approximations are also quite accurate at the global scale, deviating from the actual change in ocean-mean evaporation by 16 and 4% in the fast and slow cases, respectively.<sup>5</sup> These results show that the contribution from  $\Delta\kappa$  is small even when  $\Delta(R_s - G)$  is independent of surface warming, suggesting that Eq. (16) is robust within GCM simulations of  $\text{CO}_2$ -induced warming.

### 3.2 Implications of the Penman–Monteith perspective for changes in the sensible heat flux

Having demonstrated the approximate validity of Eq. (16), let us now consider its implications for the change in sensible heat flux ( $\Delta H$ ). When combined with the surface energy budget (Eq. 4), Eq. (16) implies that

$$\Delta H \approx -LE \frac{d \ln \eta}{dT_a} \Delta T_a + (1 - \eta) \Delta(R_s - G), \quad (20)$$

which can be further decomposed into fast and slow components, following the same line of reasoning used to derive Eqs. (17) and (19):

$$\Delta H_{\text{fast}} \approx (1 - \eta) \Delta(R_s - G)_{\text{fast}}, \quad (21)$$

$$\Delta H_{\text{slow}} \approx \left( -LE \frac{d \ln \eta}{dT_a} + (1 - \eta) \lambda \right) \Delta T_a. \quad (22)$$

These equations provide insight into how  $L\Delta E$  and  $\Delta H$  are partitioned in the fast and slow responses to  $\text{CO}_2$  warming. In the fast case, Eqs. (17) and (21) imply that

$$\frac{\Delta H_{\text{fast}}}{L\Delta E_{\text{fast}}} \approx \frac{1 - \eta}{\eta} = \beta_0, \quad (23)$$

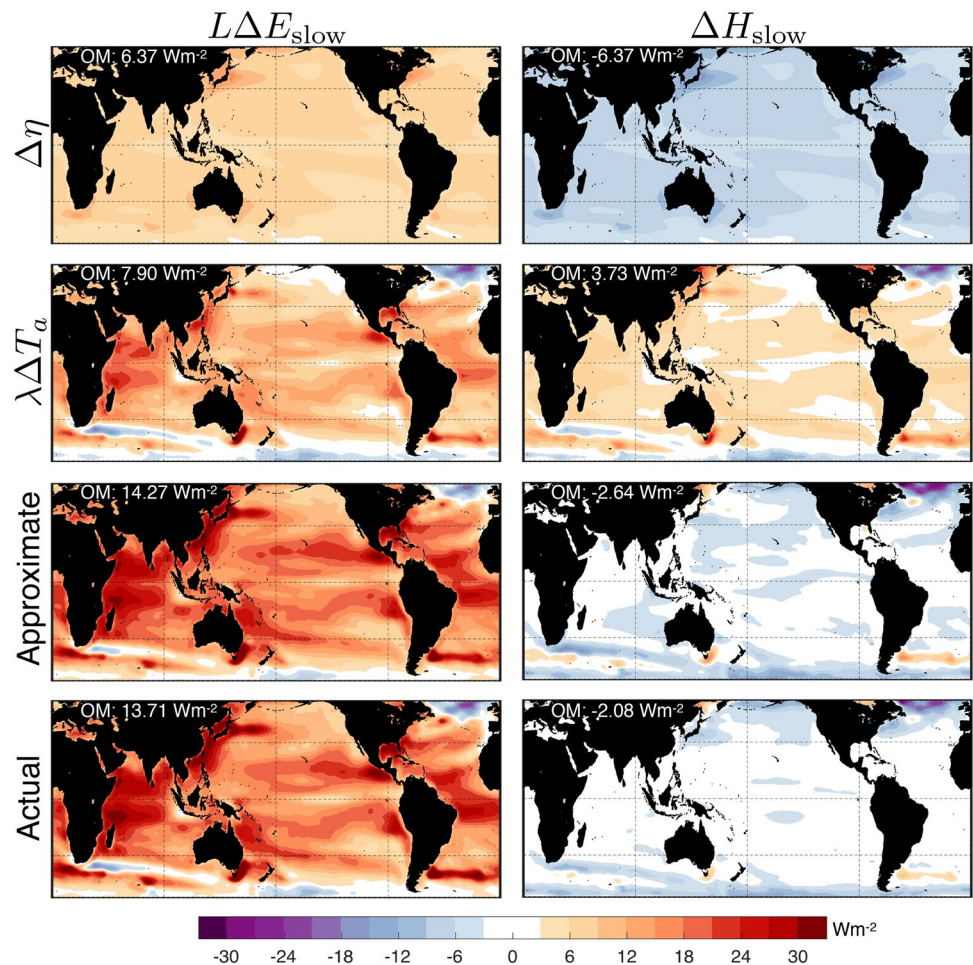
which is a function of  $T_a$  alone (Eq. 13). In the current climate,  $\beta_0$  ranges from around 0.25 in the tropical warm pool to more than 1.5 at high latitudes, with an ocean-mean value of around 0.5. Even though this result is specific to the oceans, it helps explain why in the fast response to  $\text{CO}_2$  forcing,  $L\Delta E$  significantly exceeds  $\Delta H$  in the global mean (e.g., Bala et al 2008).

In the slow case,  $L\Delta E$  (Eq. 19) and  $\Delta H$  (Eq. 22) both consist of two terms: one stemming from a change in  $\eta$ , and the other from the slow change in  $R_s - G$  (i.e.,  $\lambda \Delta T_a$ ). Figure 3 shows the contributions from these terms in the ensemble mean of the CMIP5 simulations. The  $\Delta\eta$  terms (top row) are equal and opposite to each other, representing an increase in  $LE$  at the expense of  $H$ . In contrast, both  $\lambda \Delta T_a$  terms (second row) are generally positive, with magnitudes that differ by a factor of  $\beta_0$ , mirroring the response to  $\Delta(R_s - G)_{\text{fast}}$  (Eq. 23). When these contributions are combined (third row), the positive tendencies on  $L\Delta E$  reinforce each other, while the opposing tendencies on  $\Delta H$  mostly cancel. This explains why ocean-mean  $\Delta H$  is close to zero in the slow response to  $\text{CO}_2$  forcing (bottom row; Andrews et al 2009), and why it is significantly smaller in magnitude than ocean-mean  $L\Delta E$

<sup>5</sup> These percentages are based on a comparison of the ocean-mean values that appear in the top left of each panel in Fig. 2.



**Fig. 3** The slow changes in  $LE$  (left column) and  $H$  (right column) in GCM simulations of an abrupt quadrupling of atmospheric  $CO_2$  (in  $Wm^{-2}$ ). Top row: the contribution from the  $\Delta\eta$  term in Eqs. (19) and (22). Second row: the contribution from the  $\lambda\Delta T_a$  term in Eqs. (19) and (22). Third row: the full approximation given by Eqs. (19) and (22). Fourth row: the actual change in the ensemble mean of CMIP5 simulations. Note that the two panels on the bottom left match those in the right column of Fig. 2



in the total (fast + slow) response, where  $\Delta(R_s - G)$  is also generally positive.

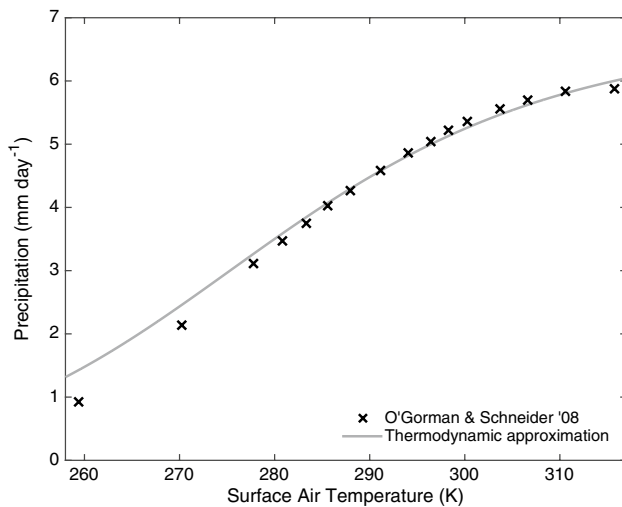
This result points to an important conceptual difference between the Penman–Monteith perspective and other interpretations of evaporation change based on diagnostic assessments of the surface or atmospheric energy budgets. In the surface energy budget, the relatively small decrease in  $H$  over the global oceans implies that the increase in  $LE$  is mostly offset by an increase in  $R_s - G$ . Similarly, in the global-mean atmospheric energy budget, the small magnitude of  $\Delta\bar{H}$  means that  $\Delta\bar{L}\bar{P}$  is mostly offset by  $\Delta\bar{R}_a$  (Eq. 1).

Within the Penman–Monteith framework, however, Fig. 1 shows that  $\Delta(R_s - G)$  accounts for just 45 and 41% of the change in ocean-mean evaporation in the equilibrium and transient simulations, respectively—a very different diagnosis than suggested by the surface energy budget alone. This difference stems from the fact that, in addition to the surface energy budget, the Penman–Monteith equation also incorporates other physical constraints related to the turbulent transfer of latent and sensible heat from the surface (Eqs. 2–3). In particular,  $\eta$  represents a thermodynamic constraint on the partitioning between  $LE$  and  $H$ , as noted in Sect. 2. When

this constraint is combined with energy conservation,  $\Delta H$  and  $L\Delta E$  are found to be closely related, each responding to  $\Delta T_a$  and  $\Delta(R_s - G)$  according to Eqs. (16) and (20). In GCM simulations,  $\Delta T_a$  and  $\Delta(R_s - G)$  conspire to make  $\Delta H$  relatively small over most of the oceans, but this is by no means a general result. In a much cooler climate, for example, Eq. (20) indicates that  $\Delta H$  would be significantly larger in response to a similar amount of warming.

#### 4 Thermodynamic constraints on global precipitation over a wide range of climates

In the previous section, we showed that the change in ocean evaporation in simulations of  $CO_2$ -induced warming can be separated into a thermodynamic component that is a direct consequence of surface warming, and a diagnostic component that represents the effects of changes in net surface radiation, ocean heat uptake, and boundary-layer dynamics/relative humidity. Of these, the thermodynamic component was found to account for 2/3 of the total increase in ocean



**Fig. 4** Global mean precipitation ( $\bar{P}$ ) as a function of global-mean surface-air temperature ( $\bar{T}_a$ ), according to the idealized simulations of O’Gorman and Schneider (2008) (black x’s), and the thermodynamic approximation assuming that  $\eta$  varies with global-mean surface temperature and  $R_s - G + \kappa$  is held fixed at  $197 \text{ W m}^{-2} \text{ K}^{-1}$  (Eq. 24; gray line)

evaporation in equilibrium slab-ocean simulations, and for an even larger share in transient coupled simulations. The thermodynamic component is also strongly dependent on mean-state temperature, explaining why the largest fractional increase in evaporation tends to occur at high latitudes in Fig. 1. In light of these results, it is worth considering how thermodynamics might influence global precipitation over a wide range of climates much warmer and cooler than our own.

For this portion of our analysis, we revisit a series of idealized gray-radiation simulations run by O’Gorman and Schneider (2008), which were designed to exhibit a wide range of global-mean surface temperatures in response to imposed changes in atmospheric longwave optical depth, albeit with no representation of the radiative effects of clouds and atmospheric water vapor. Since the output from these simulations is not publicly available, we are unable to diagnose the factors contributing to their differences in global-mean precipitation, as we did for the CMIP ensembles. However, we can estimate the thermodynamic contribution to these differences by assuming that  $\eta$  varies with global-mean surface temperature (Eq. 11), and that all other variables are approximately constant ( $R_s - G + \kappa = 197 \text{ W m}^{-2}$ ; see “Appendix”). This gives the following approximation for global precipitation as a function of global-mean surface temperature alone (Eq. 10):

$$L\bar{P} \approx \eta(\bar{T}_a) \times 197 \text{ W m}^{-2}. \quad (24)$$

Figure 4 shows the estimated global precipitation based on Eq. (24) (gray line) compared with the actual results of

O’Gorman and Schneider (2008) (black x’s). Despite some discrepancies at low temperatures, the agreement is good overall, suggesting that most of the change in global precipitation among their idealized simulations can be explained as a direct consequence of the change in global-mean surface temperature.

Of course, the real world is more complicated than the idealized, gray-radiation GCM used by O’Gorman and Schneider (2008). This might explain why changes in surface radiation appear to be smaller in these simulations than in the more realistic CMIP simulations discussed previously.

Yet even if surface radiation is not constant, it is still instructive to consider how global precipitation varies with surface temperature as a consequence of thermodynamics alone. Taking the derivative of  $\eta$  with respect to  $T_a$ , we find that the slope of the  $\bar{P}$ -vs.- $\bar{T}_a$  curve in Fig. 4 is proportional to

$$\frac{d\bar{P}}{dT_a} \propto \frac{\beta_0}{(1 + \beta_0)^2} \left( \alpha - \frac{2}{\bar{T}_a} \right). \quad (25)$$

Importantly, this equation encapsulates what Pierrehumbert (2002) has identified as distinct constraints on global precipitation operating in different temperature regimes. At the cold extreme,  $\beta_0 \gg 1$  as  $q^*$  approaches 0. While this results in a large *fractional* increase in global precipitation with warming (Eq. 15), the actual increase is small (Eq. 25), reflecting the atmosphere’s limited capacity at cool temperatures to maintain water vapor against condensation. Conversely, at very warm temperatures,  $\beta_0 \ll 1$  due to high values of  $q^*$ . In this regime, atmospheric water vapor is plentiful, but the change in global precipitation with warming is limited by net radiation at the surface (Eq. 4). It is therefore between these limits, where  $\beta_0$  is  $O(1)$  (i.e.,  $\bar{T}_a \approx 280 \text{ K}$ ), that global precipitation is thermodynamically most sensitive to changes in global mean surface temperature. While changes in surface radiation will affect the upper bound on global precipitation in the warm limit, the broad shape of the  $\bar{P}$ -vs.- $\bar{T}_a$  curve is guaranteed by thermodynamic constraints inherent in  $\eta$ .

Finally, it is interesting to consider how the primacy of  $\eta$  in GCM simulations of  $\text{CO}_2$ -induced warming relates to two other ideas for how global precipitation changes with warming. The first, put forth by Kleidon and Renner (2013a, b), is that the hydrologic-cycle heat engine operates near the thermodynamic limit of maximum power. Using an idealized energy balance model, Kleidon and Renner (2013b) explore the implications of this assumption for how global evaporation/precipitation scales with warming, assuming no change in net surface radiation. Consistent with Eq. (24), they find that global evaporation is proportional to a term that can be shown to be

identical to  $\eta$ , except that  $\overline{T_a}$  in their model represents the average temperature at which water vapor condenses in the atmosphere. This suggests that Kleidon and Renner's foundational assumption—that atmospheric convection approaches the thermodynamic limit of maximum power—may indeed have some relevance to GCM simulations, and perhaps also to the real atmosphere.

A second, similarly idealized conceptualization of the hydrologic cycle was proposed by Takahashi (2009), who argued that the change in precipitation with global warming is controlled by radiative cooling from the free troposphere rather than the full atmosphere. This idea—which represents a variant of the atmospheric energy budget perspective—is based on the principle that little of the surface sensible heat flux makes it out of the marine boundary layer. Using a one-dimensional radiative–convective–equilibrium model constrained by this and a few other a priori assumptions, Takahashi found changes in precipitation with warming that were similarly consistent with the idealized results of O'Gorman and Schneider in Fig. 4 (O'Gorman et al 2012). While the idealized models of Takahashi (2009) and Kleidon and Renner (2013a, b) do not appear to be incompatible, it remains unclear how they might relate to each other, or to the Penman–Monteith framework presented in this paper. We hope that future research will shed light on this important question.

## 5 Discussion

In this paper, we have shown that the Penman–Monteith equation, as expressed in Eq. (10), allows any change in ocean evaporation to be partitioned into distinct contributions from changes in surface temperature, net surface radiation, ocean heat uptake, and boundary layer dynamics/relative humidity. In GCM simulations of CO<sub>2</sub>-induced warming, we find that the majority of the change in ocean evaporation is a direct consequence of warming, represented by  $\Delta\eta$  in Eq. (14). This component of evaporation change derives from fundamental thermodynamics, and therefore does not depend on the specific nature of the radiative forcing or on the model physics. Physically, this term represents a change in the partitioning between latent and sensible heat fluxes due to an increase in the surface-air moisture gradient (required by the Clausius–Clapeyron equation), and a corresponding decrease in the surface-air temperature gradient (required by energy conservation). In fractional terms, the change in evaporation due to this effect diminishes with warming, explaining why the largest fractional changes in ocean evaporation tend to occur at high latitudes in GCM simulations (Fig. 1). Compared with this thermodynamic effect, the contribution to evaporation change from changes

in net surface radiation ( $\Delta R_s$ ) and ocean heat uptake ( $\Delta G$ ) were found to be secondary but still significant, while changes in boundary-layer dynamics and relative humidity ( $\Delta\kappa$ ) were found to be less important, particularly at regional scales.

Because  $\Delta\kappa$  is small, the Penman–Monteith framework allows the change in evaporation to be estimated from the spatial pattern of  $\Delta T_a$ ,  $\Delta R_s$ , and  $\Delta G$  alone (Eq. 16). For example, in the fast response to CO<sub>2</sub> forcing,  $R_s - G$  decreases due to significant ocean heat uptake, causing a decrease in global evaporation (and thus precipitation). Because SSTs are fixed, this decrease in evaporation is well approximated as  $\eta\Delta(R_s - G)$ . On longer timescales, surface temperatures rise and evaporation increases, in part due to thermodynamics, and in part because ocean heat uptake declines as the climate system returns to radiative equilibrium. Combined with the surface energy budget, Eq. (16) also leads to an equally accurate approximation of the change in sensible heat flux,  $H$  (Eq. 20). Thus, from the Penman–Monteith perspective,  $LE$  and  $H$  represent two sides of the same coin, each responding to  $\Delta T_a$ ,  $\Delta R_s$ , and  $\Delta G$  according to Eqs. (16) and (20).

This interpretation of global hydrologic change is somewhat different from those based on the atmospheric or surface energy budgets, in which the change in global precipitation ( $L\Delta P$ ) is offset by  $\Delta\overline{H}$  and a change in net atmospheric radiative cooling ( $\Delta\overline{R_a}$ ). The energy-budget perspective provides little insight into  $\Delta\overline{H}$ , but has much to say about the physics behind  $\Delta\overline{R_a}$  (e.g., Lambert and Webb 2008; Stephens and Ellis 2008; Previdi 2010; Pendergrass and Hartmann 2014; DeAngelis et al 2015; Fläschner et al 2016). In contrast, the Penman–Monteith equation provides new insight into the partitioning between  $L\Delta E$  and  $\Delta H$  over the oceans (Eqs. 16, 20), but only if  $\Delta T_a$  and  $\Delta(R_s - G)$  (which is equal to  $\Delta R_a$  in the global mean) are already known. This shows that the energy-budget and Penman–Monteith perspectives are fully complementary, and together provide a more complete understanding of evaporation change than either can provide by itself.

The Penman–Monteith perspective may also shed light on the response of the global hydrologic cycle to a change in the solar constant (e.g., Wetherald et al 1975; Andrews et al 2009), to changes in radiation due to solar geoengineering (Bala et al 2008), or to non-greenhouse forcings like a volcanic eruption (Trenberth and Dai 2007). Of course, the accuracy of Eq. (16) is contingent on the forcing having little impact on the dynamics or relative humidity of the atmospheric boundary layer. In certain scenarios—e.g., a change in the concentration of absorbing aerosols in the boundary layer (Ming et al 2010; Samset et al 2016)—this condition might not be met. Yet even in these cases, the Penman–Monteith framework could prove to be a powerful tool for diagnosing the various contributions to changes in ocean evaporation at both global and regional scales.



Finally, it is important to note that while changes in net surface radiation are of secondary importance to the overall change in ocean evaporation, they account for most of the inter-model spread, as evidenced by the standard deviations in the bottom row of Tables 1 and 2. Relative to the other terms, the standard deviation of the  $\Delta R_s$  contribution is roughly 2–3 times larger across both the equilibrium and transient ensembles. This is not surprising given that  $\Delta R_s$  is closely tied to  $\Delta R_a$  (Eqs. 1, 4), which depends on several model variables that are not well constrained, including clouds, tropospheric humidity, and the radiative transfer parameterization for calculating shortwave absorption by water vapor (DeAngelis et al 2015; Fläschner et al 2016). In contrast, the  $\Delta \eta$  contribution is more consistent due to broad model agreement in the spatial patterns of warming, mean-state temperature, and mean-state evaporation. Altogether, these results suggest that thermodynamics alone will contribute to an increase in global precipitation with surface warming at a rate of about 1.5%/K; whether global precipitation increases at a rate closer to 1 or 3%/K will largely depend on radiative changes.

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## Appendix 1: Calculating the contributions to evaporation change in Eq. (14)

The terms in Eq. (14) were calculated as follows:  $LE$  and  $R_s$  were taken directly from model output;  $G$  was determined from  $R_s$ ,  $LE$ , and  $H$  based on the surface energy budget (Eq. 4);  $\eta$  was calculated from the two-meter air temperature using Eqs. (11) and (13). Finally, given  $LE$ ,  $\eta$ ,  $R_s$ , and  $G$ , we then solved for  $\kappa$  in Eq. (10). The contributions were calculated from ensemble-mean output over the last 5 years of the simulation period. In the equilibrium warming simulations, this was typically 21–25 years after  $\text{CO}_2$  doubling. In the transient warming simulations, we used years 96–100 after  $\text{CO}_2$  quadrupling. The contributions were first calculated for each month, and then the monthly contributions were averaged to arrive at an annual-mean value. However, the results were essentially unchanged when the contributions were calculated from annual-mean output.

To understand the *global* impact of the fractional contributions in Fig. 1, we must account for spatial variability in the magnitude of the mean-state evaporation and surface-air warming. To do so, we multiply each term in Eq. (14) by the following (dimensionless) weighting function,

$$w = \frac{E\overline{\Delta T_a}}{\overline{E\Delta T_a}}, \quad (26)$$

where the overbars in the denominator indicate the ocean-mean values of each variable. These results are then averaged in space, yielding the ocean-mean contributions given in the top left of each panel in Fig. 1.

## Appendix 2: Estimating $R_s - G + \kappa$ in the idealized simulations of O’Gorman and Schneider (2008)

To estimate the value of  $R_s - G + \kappa$  in O’Gorman and Schneider (2008) simulations, we use the fact that their control climate exhibits a global-mean surface-air temperature of  $\overline{T_a} = 288$  K, and a global-mean precipitation of 4.3 mm/day, which equates to  $\overline{LE} = 124 \text{ Wm}^{-2}$ . Given  $\eta \approx 0.63$  at  $\overline{T_a} = 288$  K, this implies a combined value of  $R_s - G + \kappa = 197 \text{ Wm}^{-2}$ . If we assume that this sum is constant, global precipitation is directly proportional to  $\eta$ , resulting in the gray curve in Fig. 4.

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