

Bayesian computation using INLA 5: Spatial Markovian models—The SPDE approach



Outline

Spatial models Gaussian random fields

Low-dimensional methods

Spatial Matérn fields

Example: Continuous vs Discrete

Useful features for manipulating the latent field

Joint modelling of covariates

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Spatial interpolants

(observed value)_i = (true value at location i) + (error)_i

- ▶ We treat the *observed* covariates as being measured with error
- The errors are usually assumed to be independent and identically distributed (i.i.d.)
 - Usually, we take them to be Gaussian
 - If we think there may be outliers, we might use something else (e.g. a Student-T distribution)
 - The only change in R-INLAis in the family argument in the INLA call

So how does that help us fill in the field?

(observed value)_i = (true value at location i) + (error)_i

or

$$y_i = x(s_i) + \epsilon_i$$

We need priors!

- We have chosen the error distribution to be $\epsilon_i \sim N(0, \sigma^2)$
 - A zero mean means that there is no systemic measurement error
 - A common variance means that everything was measured the same way

Now we need a prior on the truth...

If we have a process that is occurring everywhere in space, it is natural to try to model it using some sort of function.

- This is hard!
- We typically make our lives easier by making everything Gaussian.

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What makes a function Gaussian?

- We don't ever observe a function *everywhere*.
- If x is a vector of observations of x(s) at different locations, we want this to be normally distributed:

$$\mathbf{x} = (x(s_1), \ldots, x(s_p))^T \sim N(\mathbf{0}, \mathbf{\Sigma}_{x(s_1), \ldots, x(s_p)})$$

- This is actually quite tricky: the covariance matrix Σ will need to depend on the set of observation sites and always has to be positive definite.
- ► It turns out you can actually do this by setting $\Sigma_{ij} = c(s_i, s_j)$ for some *covariance function* $c(\cdot, \cdot)$.
- Not every function will ensure that Σ is positive definite!

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Stationary random fields

A GRF is stationary if:

- has mean zero.
- the covariance between two points depends only on the distance and direction between those points.

It is **isotropic** if the covariance only depends on the *distance between the points*.

- Zero mean —> remove the mean
- Stationarity is a mathematical assumption and may have no bearing on reality

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The three typical parameters for a GRF

• The variance (or precision) parameter:

- This controls how wildly the function can deviate from its mean
- The range parameter
 - This controls the range over which the correlation between x(s) and x(s + h) is essentially zero
 - Often the "range" parameter is some transformation of this distance

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The smoothness parameter

- Controls how differentiable the field is.
- This essentially controls how similar nearby points are
- Often not jointly identifiable with the range

For isotropic random fields, these parameters are constant.

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Defn: Gaussian random fields

A Gaussian random field x(s) is defined by a mean function $\mu(s)$ and a covariance function $c(s_1, s_2)$. It has the property that, for every finite collection of points $\{s_1, \ldots, s_p\}$,

$$\mathbf{x} \equiv (x(s_1),\ldots,x(s_p))^T \sim N(\mathbf{0},\mathbf{\Sigma}),$$

where $\Sigma_{ij} = c(s_i, s_j)$.

- **Σ** will almost never be sparse.
- It is typically very hard to find families of parameterised covariance functions.
- It isn't straightforward to make this work for multivariate, spatiotemporal, or processes on non-flat spaces.

How big is the problem?

Let's do a quick operation count!

- Parameter estimation: Requires the field at N data points: Must factor an N × N matrix
- Spatial prediction (Kriging): Requires the field at *m* points, probably densely through the domain: Must factor an *m* × *m* matrix.

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▶ Joint parameter and spatial estimation: Needs it at both. Must factor a (N + m) × (N + m).

Working with dense matrices

Storage:

- ▶ $\mathcal{O}(N^2)$
- $\blacktriangleright\,$ 2500 points for 20 years requires \sim 20 Gbytes
- Computation:
 - Each sample, solve, or determinant costs $\mathcal{O}(N^3)$.
 - ▶ We always need quite a few of these (likelihood, VB, INLA)
 - If we use MCMC we need a gargantuan number!
 - \blacktriangleright Remember 1,000,000 samples give \sim 3 decimal places of accuracy.

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Clearly this won't work if N is large.

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What if our matrices magically became sparse?

Sparse covariance matrices can be formed using covariance tapering or compactly supported covariance functions.

- Storage:
 - ▶ $\mathcal{O}(N)$
 - $\blacktriangleright\,$ 2500 points for 20 years requires \sim 400 Kilobytes
- Computation:
 - Each sample, solve, or determinant costs $\sim O(N^{3/2})$.

Question: Can we find random field that 'magically' give us sparse matrices?

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Low-dimensional methods

Kernel methods Approximation Why do kernel methods fail? Better spatial approximation

Spatial Matérn fields

Example: Continuous vs Discrete

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The Goldilocks principle

So modelling directly with GRFs is too hard, while GMRFs can be difficult when they're not on a lattice.

Is there something "just right"?

It is instructive to consider another "awkward" method that aims to fix the computational problems of GRFs while keeping a continuous specification (hence avoiding the lattice problem).

Combining this method with GMRFs will lead to (finally!) a stable, flexible, computationally feasible method.

Reducing the dimension

Most of the methods aimed at reducing the "big N problem" in spatial statistics is based on some sort of low-dimensional approximation:

$$\mathbf{x}(s) pprox \sum_{i=1}^{n} w_i \phi_i(s),$$

where \boldsymbol{w} is jointly Gaussian and $\phi_i(s)$ are a set of known deterministic functions.

If $\boldsymbol{w} \sim N(\boldsymbol{0}, \boldsymbol{\Sigma})$, then the covariance function of $x(\boldsymbol{s})$ is

$$c(\boldsymbol{s}_1, \boldsymbol{s}_2) = \boldsymbol{\Phi}(\boldsymbol{s}_1)^T \boldsymbol{\Sigma} \boldsymbol{\Phi}(\boldsymbol{s}_2),$$

where $\Phi(s)$ is a vector with the ϕ_i functions evaluated at point s.

We will take a close look at Kernel methods and, in particular, when they fail.

- They are popular!
- They are easy to analyse!
- They are prototypical of this low-dimensional (please don't say "low-rank"!!) approach.

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Kernel representations

Most GRFs can be represented as

$$x(s) = \int_{\mathbb{R}^2} k(s,t) \, dW(t),$$

where W(t) is white noise, and k(s, t) is a deterministic "kernel" function.

- It is often suggested that we model $k(\cdot, \cdot)$ directly.
- ▶ We can approximate the integral by a sum (Higdon, '98)

$$x(s) \approx \sum_{i=1}^n k(x,t_i)\xi_i,$$

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where ξ_i are i.i.d. normals.

 This does not work well. (S, Lindgren, Rue, '10, Bolin and Lindgren '10)

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So what happens?



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Approximation properties

Realisations of Gaussian Random Fields are functions.

Appropriate question

How well can realisations of $x(\cdot)$ be approximated by functions of the form $\sum_{i=1}^{n} w_i \phi_i(s)$.

- This is not an asymptotic question! n never goes to infinity.
- Without considering these questions, you *cannot* know how a method will work!

Best Kernel approximation to a constant



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Why did kernel methods perform badly?

 Kernel methods performed badly because there weren't enough points.

- Kernel methods performed badly because the range was smaller than the grid spacing.
- Kernel methods performed badly because the basis functions depend on the parameter being inferred!

This is a common problem and leads to "spotty" spatial predictions and bad uncertainty estimates.

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Piecewise linear approximation of surfaces



NB: The basis functions are only non-zero on a small part of the domain.

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Known approximation properties



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How can we use these functions?

There is no obvious way to use piecewise linear functions...

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Spatial Matérn fields The spatial Matérn property The finite element method

Notes

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In this section we're going to look at a class of flexible, computationally efficient spatial models.

- These models will give you a lot of the flexibility of GRFs without the pain
- They have the good computational properties of GMRFs, but with more flexibility
- They are defined continuously like Kernel methods, but are stable
- Don't focus too much on the theory (unless you want to!)

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The secret is in the Markov property



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How does this translate to maths?

General Result

The power spectrum of a stationary Markovian Gaussian random field has the form $R(\mathbf{k}) = 1/p(\mathbf{k})$, where $p(\mathbf{k})$ is a positive, symmetric polynomial.

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Oh dear!

Can we salvage something from this?

Sometimes it's useful to be an engineer!

An engineering calculation

Let L be a differential operator. Then the solution to

 $Lx(s) = W(\cdot)$

is a Gaussian random field and it has the Markov property.

"Prove" it using Fourier transforms.

- ► The derivatives (local) produce the Markov property (local)
- ▶ Now we're solving (partial) differential equations: *standard!*

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What does this remind us of?

Recall our SAR(1) model (slightly re-written)

$$4x_i - (x_n + x_s + x_e + x_w) \sim N(0, \sigma^2).$$

Also remember that

$$-\frac{d^2x}{ds^2}\approx\frac{-x(s+h)+2x(s)-x(s-h)}{h^2}$$

So if we scale our SAR(1) model and let the lattice spacing h → 0, we get

$$-\Delta x(s) \equiv -\left(rac{d^2x}{ds_1^2}+rac{d^2x}{ds_2^2}
ight) \stackrel{d}{=} W(s)$$

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In the context of GMRFs

In this context, this was first noted by Whittle in the 50s (!!) who noted that Matérn fields, which have covariance function of the form

$$c(x,y) \propto (\kappa ||x-y||)^{\nu} K_{\nu} (\kappa ||x-y||),$$

are the stationary solutions to the SPDE

$$(\kappa^2 - \Delta)^{\frac{\nu + d/2}{2}} x(s) = W(s),$$

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where

•
$$\Delta = \sum_{i=1}^{d} \frac{\partial^2}{\partial s_i^2}$$
 is the Laplacian

- W(s) is spatial white noise.
- The parameter ν controls the smoothness.
- The parameter κ controls the range.

Practical interpretation of the parameters

We have

$$(\kappa^2 - \Delta)^{\frac{lpha}{2}}(\tau x(s)) = W(s),$$

where $\alpha = \nu + d/2$ is an integer.

• κ^2 is a range parameter. The approximate range is

range
$$\approx rac{\sqrt{8
u}}{\kappa}.$$

The variance of the model is

$$\sigma^2 = \frac{\Gamma(\nu)}{\gamma(\nu+d/2)(4\pi)^{d/2}\kappa^{2\nu}\tau^2}.$$

So, according to the Whittle characterisation of the Matérn covariance functions, we get a Markovian random field when $\alpha = \nu + d/2$ is an integer. When d is odd, Matérn models with $\nu \in 1/2\mathbb{N}$.

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- This include the Thin Plate Spline model ($\nu = 1$)
- And the exponential covariance ($\nu = 1/2$).

When d is even, we get Matérn models with $u \in \mathbb{N}$

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Let's simplify things: set $\nu + d/2 = 2$

The SPDE becomes

$$(\kappa^2 - \Delta)x(s) = W(s),$$

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which only involves second derivatives, which are nice.

Connection to thin plate splines!

Approximating the SPDE

We are looking for the piecewise linear random field

$$\mathsf{x}_n(s) = \sum_{i=1}^n w_i \phi_i(s)$$

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for piecewise linear functions $\phi_i(s)$ that *best* approximates the solution to $(\kappa^2 - \Delta)x(s) = W(s)$.

Step 1: The 'weak' solution

White noise is weird, but if we integrate it, it becomes nicer. So we require that for every suitable function $\phi(s)$,

$$\int_{\Omega} \psi(s)(\kappa^2 - \Delta) x(s) \, ds \stackrel{D}{=} \int_{\Omega} \psi(s) \, dW(s).$$

White noise integrals aren't scary!

$$\int_{\Omega} \psi(s) \, dW(s) \sim N(0, \int_{\Omega} \psi(s)^2 \, ds)$$

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Step 2: Plug in the basis functions

Replace x(s) with the basis function expansions and chose $\phi(s)$ to be the set of basis functions

We get the system of linear equations

$$\int_{\Omega} \phi_j(s)(\kappa^2 - \Delta) \left(\sum_i w_i \phi_i(s)\right) ds \stackrel{D}{=} \int_{\Omega} \phi_j(s) dW(s)$$

This is good— LHS has things we can compute, RHS has integrals of white noise.

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What comes out?

We get two matrices:

The (scary) SPDE becomes the (normal) equation $(w^2C + K)w = \Lambda((0, C))$

$$(\kappa \mathbf{C} + \mathbf{K}) \mathbf{w} \sim \mathcal{N}(\mathbf{0}, \mathbf{C})$$

and therefore \boldsymbol{w} is a GMRF with precision matrix

$$\mathbf{Q} = \left(\kappa^{2}\mathbf{C} + \mathbf{K}\right)^{T}\mathbf{C}^{-1}\left(\kappa^{2}\mathbf{C} + \mathbf{K}\right).$$

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What comes out?

We get two matrices:

•
$$C_{ii} = \int_{\Omega} \phi_i(s) ds$$
 (the constant terms)

• $\mathsf{K}_{ij} = \int_{\Omega} \nabla \phi_i(s) \cdot \nabla \phi_j(s) \, ds$ (the Laplacian term)

The (scary) SPDE becomes the (normal) equation

$$(\kappa^{2}\mathbf{C}+\mathbf{K})\mathbf{w}\sim\mathcal{N}\left(0,\mathbf{C}
ight)$$

and therefore \mathbf{w} is a GMRF with precision matrix

$$\mathbf{Q} = \left(\kappa^2 \mathbf{C} + \mathbf{K}\right)^T \mathbf{C}^{-1} \left(\kappa^2 \mathbf{C} + \mathbf{K}\right).$$

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- ► This works for any Matérn field where α = ν − d/2 is an integer.
- More importantly, this works for any SPDE Lx = W.
- ► More importantly, if we Q = L*L, this method can be applied directly to the precision operator Q.
- ► We can approximate non-Markov fields by Markovian fields by approximating 1/R(k) by a polynomial.

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Spatial models

Low-dimensional methods

Spatial Matérn fields

Example: Continuous vs Discrete

Useful features for manipulating the latent field

Joint modelling of covariates

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Leukaemia survival

Leukaemia survival data (Henderson et al, 2002, JASA), 1043 cases.



Fig. 1. Leukaemia survival data: districts of Northwest England and locations of the observations.

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Survival models are different for many models in statistics. There are two types of observations: an event (death) or we stop measuring (censoring).

Rather than directly modelling the *hazard* (instantaneous risk)

$$\begin{array}{lcl} h(y) \ dy & = & \operatorname{Prob}(y \leq Y < y + \ dy | Y > y) \\ h(y) & = & \frac{f(t)}{S(t)} \end{array}$$

Cox proportional hazards model

Write the hazard function for each patient as:

$$h(y_i|w_i, \boldsymbol{x}_i) = h_0(y_i) w_i \exp(\boldsymbol{c}_i^T \boldsymbol{\beta}) \exp(x(s_i)); i = 1, \dots, 1043$$

where

- $h_0(\cdot)$ is the baseline hazard function
 - w_i is the log-Normal frailty effect associated with patient *i*

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- \boldsymbol{c}_i is the vector of observed covariates for patient i
- β is a vector of unknown parameters
- $x(s_i)$ is the value of the spatial effect x(s) for patient *i*.

Spatial survival: example

log(hazard) = log(baseline) + f(age) + f(white blood cell count) + f(deprivation index) + f(spatial) + sex



Fig. 1. Leukaemia survival data: districts of Northwest England and locations of the obse
R-code (regions)

```
data(Leuk)
g = system.file("demodata/Leuk.graph", package="INLA")
```

```
formula = inla.surv(Leuk$time, Leuk$cens) ~ sex + age +
  f(inla.group(wbc), model="rw1")+
  f(inla.group(tpi), model="rw2")+
  f(district, model="besag", graph = g)
```

result = inla(formula, family="coxph", data=Leuk)

source(system.file("demodata/Leuk-map.R", package="INLA"))
Leuk.map(result\$summary.random\$district\$mean)
plot(result)

baseline.hazard



PostMean 0.025% 0.5% 0.975%

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● = ◆□



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inla.group(wbc)



PostMean 0.025% 0.5% 0.975%

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inla.group(tpi)



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Continuous spatial effect



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SPDE models

We call spatial Markov models defined on a mesh SPDE models.

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SPDE* models have 3 parts

- A mesh
- A range parameter κ
- A precision parameter τ

SPDE=Stochastic Partial Differential Equation

Meshes can be created using two different functions:

inla.mesh.create: The workhorse function. An interface to the meshing code written by Finn Lindgren.

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inla.mesh.2d: A slightly more user friendly interface for creating practical meshes (we will focus on this one).

Typical use



inla.mesh.create.helper

```
inla.mesh.create.helper(loc = NULL,
                 loc = NULL.
              loc.domain = NULL.
              offset = NULL,
              n = NULL,
                boundary = NULL,
              interior = NULL,
              max.edge,
              min.angle = NULL,
              cutoff = 0,
              plot.delay = NULL)
```

This function contains a mesh with two regions: the interior mesh, which is where the action happens; and the exterior mesh, which is designed to alleviate the boundary effects.

- Ioc: Points to be included as vertices in the triangulation.
- loc.domain: Points not in the mesh, but that are used to define the internal mesh section (taken as the convex hull of these points).
- offset=c(a,b): Distance from the points to the inner (outer) boundary. Negative numbers = relative distance.
- boundary: Prescribed boundary. (inla.mesh.segment type)
- max.edge = c(a,b): Maximum triangle edge length in the inner (outer) segment.
- min.angle = c(a,b): Minimum angle for the inner and outer segments (bigger angles are better, but harder to make)
- cutoff: Minimum distance between two distinct points.

Good and bad meshes



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Between the mesh and the data

- So a good mesh probably doesn't have vertices at the data locations
- This means we need to have a way to get between values of the field at the vertices and the value of the field at the data points
- The trick is that the SPDE model is *linear* on the triangles, so the value of the field at any point is a weighted sum of the vertices of the triangle the point is in.

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- ► In maths speak, we are observing **Ax** rather than **x**
- We call A the "A-matrix" or the "observation matrix"

Making observation matrices in INLA

When the observations don't occur at mesh points, we need some way to map between the latent field and the observation process.

- ▶ inla.spde.make.A constructs the matrix A_{ij} = φ_j(s_i) that maps a field defined on the mesh to the observation locations s_i.
- The function will also automatically deal with space-time models and replicates.
- A related function (inla.mesh.projector) builds an A-matrix for projecting onto a lattice. This is useful for plotting.

The inla.spde.make.A call

```
inla.spde.make.A(mesh = NULL,
                 loc = NULL.
                 index = NULL.
                 group = NULL,
                 repl = 1L,
                 n.mesh = NULL,
                 n.group = max(group),
                 n.repl = max(repl),
                 group.mesh = NULL,
                 group.method = c("nearest", "S0", "S1"),
                 weights = NULL)
```

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- The first two arguments are needed.
- group is needed to build space-time models
- The other arguments are fairly advanced!

Other mesh commands

- inla.mesh.segment: Constructs an object that can be given to inla.mesh.create as a boundary or interior segment
- inla.mesh.boundary: Extracts a boundary segment from a mesh.
- inla.mesh.project and inla.mesh.projector: Projects results from a mesh to a lattice. Useful for plotting.
- inla.mesh.basis: Constructs a B-spline basis of a given degree on a mesh.
- inla.mesh.query: Extracts information about the topology of the mesh (advanced!)

Constructing SPDE models

For historical reasons there are two different SPDE classes (spde1 and spde2)

- spde1 is the "classic" SPDE model!
- The spde2 class is more flexible and defines non-stationarity in a more natural way.
- The primary difference between the two models is in the prior specification.
- At some point there will probably be an spde3 class: We are interested in backwards-compatability!
- For "stationary" models, these are fairly much the same (up to prior specification)

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The spde1 call

inla.spde.create(mesh,

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- "imatern" is the intrinsic model ($\kappa^2 = 0$).
- "matern.osc" is an oscillating Matérn model.
- param is a list that contains alpha (1 or 2) and stuff about non-stationarity.

The problem with the spde1 comes when specifying non-stationarity.

- Suppose we want to model τ(s) = Σ^k_{i=1} θ^τ_i b_i(s) for some basis functions {b_i(s)}. (Similar for κ²(s))
- The spde1 model put i.i.d. log-normal priors on the θ_i .
- This is not a good idea: what if we want a smooth effect—should have a spline prior...
- ▶ We also penalise the (log) variance directly:

$$\log(\sigma^2) = \text{const.} - 2\log(\kappa) - 2\log(\tau)$$

spde2 fixes this by putting a multivariate normal prior on

$$\log(\tau) = \boldsymbol{B}^{ au} \boldsymbol{ heta}, \qquad \log(\kappa^2) = \boldsymbol{B}^{\kappa} \boldsymbol{ heta}$$

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$$\log(\sigma^2) = ext{const.} - 2\log(\kappa) - 2\log(au)$$

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$$\log(oldsymbol{ au}) = oldsymbol{B}^ au oldsymbol{ heta}, \qquad \log(\kappa^2) = oldsymbol{B}^\kappa oldsymbol{ heta}$$

The spde2 call

inla.spde2.matern(mesh,

```
alpha = 2,
B.tau = matrix(c(0,1,0),1,3),
B.kappa = matrix(c(0,0,1),1,3),
prior.variance.nominal = 1,
prior.range.nominal = NULL,
prior.tau = NULL,
prior.tau = NULL,
theta.prior.mean = NULL,
theta.prior.mean = NULL,
fractional.method = c("parsimonious", "null"))
```

mesh: An inla.mesh object. (Necessary)

- ▶ alpha =2: The smoothness. Exact fields if it's an integer, approximate fields for non-integer α
- **b** B.tau: The matrix \boldsymbol{B}^{τ} use to define non-stationary $\tau(s)$
- B.kappa: As above, but for $\kappa^2(s)$
- > prior.variance.nominal, prior.range.nominal: Helps the automatic prior know the scale of the variance and the range
- ▶ prior.tau, prior.kappa: Prior specification for τ and κ^2 . (not often used)
- theta.prior.mean, theta.prior.prec: Mean vector and precision matrix for θ prior.
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A few more useful commands

- inla.spde.precision(spde,tau=...,kappa2=...)—
 computes precision matrix. Less straightforward for spde2
 models
- inla.qsample(n,Q,...)—Computes a sample and various
 other quantities needed for MCMC for precision matrix Q
- inla.qreordering—Computes a fill-in reducing reordering.
- inla.qsovle—Solve a linear system
- inla.qinv(Q)—Calculates the elements of the inverse corresponding to the non-zero elements of Q. Needed for computing derivatives of Gaussian likelihoods.

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Useful features

- replicate and group
- more than one "family"
- copy
- linear combinations
- A matrix in the linear predictor

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- values
- remote computing

Feature: replicate

"replicate" generates iid replicates from the same f()-model with the same hyperparameters.

If $\boldsymbol{x} \mid \boldsymbol{\theta} \sim \mathsf{AR}(1)$, then nrep=3, makes

 $\boldsymbol{x} = (\boldsymbol{x}_1, \boldsymbol{x}_2, \boldsymbol{x}_3)$

with mutually independent x_i 's from AR(1) with the same θ Arguments

f(..., replicate = r [, nrep = nr])

where replicate are integers $1, 2, \ldots, etc$

Example: replicate

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What do NAs do?

- In the covariates, an NA is treated as a zero.
- In the random effect, NAs indicate that the effect does not contribute to the likelihood

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► In the data, an NA indicates a location for prediction.

This feature fixes a limitation in the formula-formulation of the model $% \left[{{\left[{{{\rm{T}}_{\rm{T}}} \right]}_{\rm{T}}}} \right]$

The model

formula = $y \sim f(i, ...) + ...$

Only allow ONE element from each sub-model, to contribute to the linear predictor for each observation.

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Sometimes/Often this is not sufficient.

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Sometimes/Often this is not sufficient.

Suppose

$$\eta_i = u_i + u_{i+1} + \dots$$

Then we can code this as

► The copy-feature, creates internally an additional sub-model which is *e*-close to the target

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Many copies allowed, and copies of copies

Suppose

$$\eta_i = u_i + \beta u_{i+1} + \dots$$

Then we can code this as

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Suppose that

$$\eta_i = a_i + b_i z_i + \dots$$

where

$$(a_i,b_i) \stackrel{\mathsf{iid}}{\sim} \mathcal{N}_2(\mathbf{0},\mathbf{\Sigma})$$

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Multiple likelihoods

In many situations, you need to combine data from different sources and need to be able to handle multiple likelihoods.

Examples:

 Joint modelling of longitudinal and event time data (Guo and Carlin, 2004)

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- Preferential sampling (Diggle et al, 2010)
- "Marked" point processes
- Animal breeding modelling with multiple traits
- Combining data from multiple experiments

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How to do this in INLA

Make response y a matrix rather than a vector.

- > Y[1:n, 1] = y[1:n]
- > Y[1:n + n, 2] = y[(n + 1):(2 * n)]
- NAs are used to select components in the formula

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> cov1 = c(cov, rep(NA,n))

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Back to covariates

Consider a model with a linear predictor that looks like

$$\eta_i = \ldots + \beta c(s_i) + \ldots$$

where c is an unknown spatial covariate.

- c_i is unknown, but we have some measurements $\{c'_j\}$ at points $\{s'_j\}$
- We can model the true covariate field as above

$$egin{aligned} c_j' | c(\cdot) &= c(s_j') + \epsilon_j \ c(\cdot) &\sim \mathsf{SPDE} egin{aligned} \mathsf{model} \end{aligned}$$

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Joint modelling of the covariate

We can then fit these models at the same time!

Likelihood:

 $y_i \mid \eta_i \sim \text{Any INLA}$ likelihood with latent field η $c_j' | c(\cdot) \sim \mathcal{N}(\xi_j, \tau_c^{-1})$

Latent field:

$$\eta_i = \ldots + \beta c(s_i) + \ldots$$

$$\xi_j = c(s'_j)$$

- We have two likelihoods (data and covariate)
- We use the covariate field c(s) twice $\longrightarrow copy$

We begin by putting the observations and the observed covariates *together as data*

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Setting up the formula

We need to set up the formula carefully to separate out the two things. The trick is NAs in indices

The formula is then

> formula = Y ~ ...+ f(covariate_first_lik, model=spde)
+ f(covariate_second_like, copy=covariate_first_lik) -

Finally, we need to make an inla call for this model.

```
> result = inla(formula, family = c("____", "gaussian"),
    data = list(Y=Y,
    covariate_first_lik=covariate_first_lik,
    covariate_second_lik=covariate_second_lik),
    verbose=TRUE)
```

where _____ is the data likelihood.

This model in practice

 Joint modelling the covariate adds 3 hyperparameters (range, precision, noise precision)

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- This can be done any type of data (eg point patterns)
- ► If there is *misalignment*, it can get tricky
- ► In this case, you need A-matrices

Organising data, latent fields and A matrices

Real life is hard!

- In complicated models, we will have multiple sources of data occurring in different places with different likelihood.
- The latent field may also be composed of sections defined at different resolutions (grid for a spatial covariate, mesh for random field, etc).
- So we need a function that takes these components and chains them together in a way that makes sense.
- (You can "roll your own" here, but I *really* don't recommend it!)

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We are rescued by inla.stack!

The inla.stack call

- The trick here is lists!
- The first element of the effects list is mapped to the first element of the data list by the first element for the A list.
- Slightly more tricky when there are replicates and grouping (time!)
- The functions inla.stack.data(stack) and inla.stack.A(stack) are used to extract the data.frame and the A-matrix for use in the inla call.