

SER-SAG

in-kind proposal

Serbian AGN (SER-SAG) Team
LSST AGN Science Collaboration

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SER-SAG Team

- Proposal Lead: Luka Popović (AOB)
- Core Team: Anđelka Kovačević (MatF), Maša Lakičević (AOB), Dragana Ilić (MatF)
- Other members:
Saša Simić (FSUK), Marko Stalevski (AOB), Oliver Vince (AOB), Edi Bon (AOB), Djordje Savić (AOB), Jelena Kovačević-Dojčinović (AOB), Nataša Bon (AOB), Nemanja Rakić (PhD Student, MatF), Isidora Jankov (PhD student, MatF), Sladjana Mandić-Marčeta (PhD student, AOB), Iva Čvorović Hajdinjak (PhD Student, MatF)
- Currently Serbian science facing huge challenges:
 - complete change in financing: institutional + limited projects
 - advices from the region most welcomed

Research: AGN & microlensing

- Long experience of AGN research and microlensing
- Time-domain: special focus on AGN variability
 - Periodicity (oscillation in light curves; time-delays in light curves)
- Monitoring-campaigns (spectroscopy & photometry)

Vidojevica 1.4m



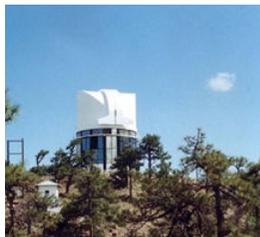
BTA 6m



Rozhen 2m



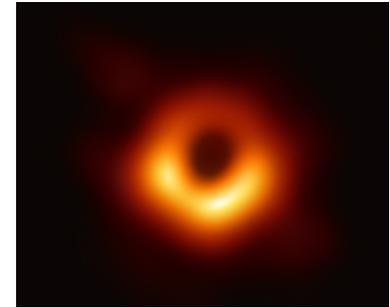
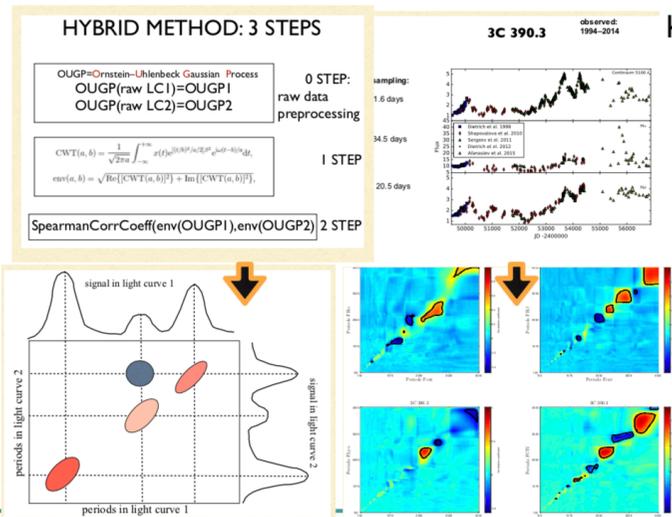
OAN SPM 2m



GHO 2m



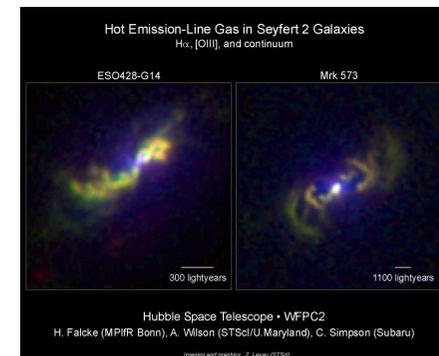
Calar Alto 2m



SMBH, EHT Collaboration, 2019



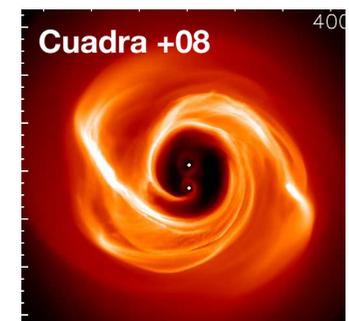
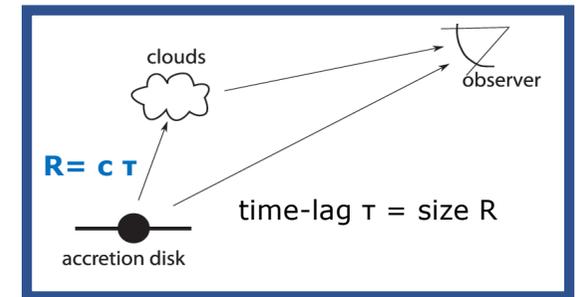
Broad line region, dusty torus



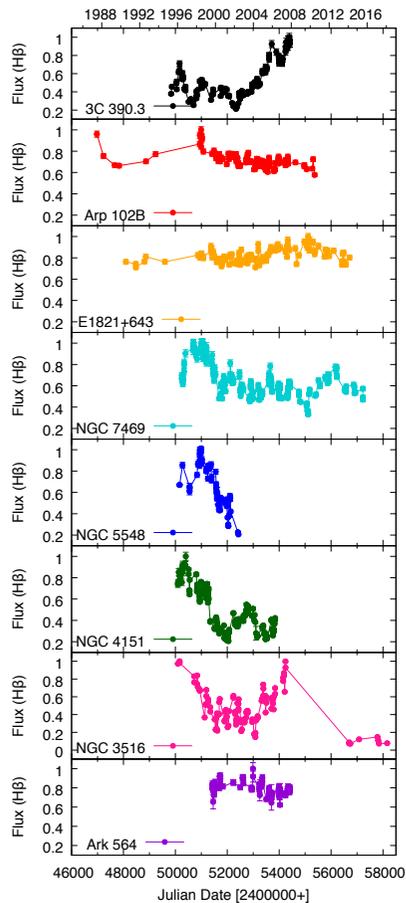
Narrow line region

AGN variability hot topics

- AGN core difficult to resolve with current optical telescopes (except w/interferometry, e.g. GRAVITY Sturm+2018, EHT Collaboration, 2019)
→ we can resolve it in time-domain
- study AGN accretion disk and BLR through **reverberation mapping**:
→ one of the priorities of LSST AGN SC (e.g. Brandt et al. 2018)
- detect oscillation in AGN light curves, **searching for periodicities**
→ important for detection of close binary SMBHs, and possible GW sources (for a recent review see De Rosa et al. 2020)



AGN long-term (decades) monitoring campaign



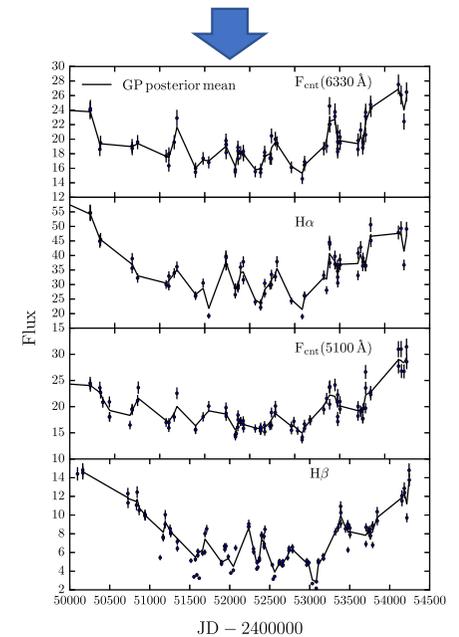
our RM campaign features:

- long (10+ years) uniform dataset
- different sub-types of type 1 AGN of different variability and optical spectra behavior
- spectral campaign → some problems in light curves such as e.g. **gaps**, **nonuniform cadence**, etc.
- applied machine learning, i.e. Gaussian processes to model light curves → extract time lags

Gaussian PROcesses for TIme-Delays Estimates in AGN **GPro-TIDE**

- a complementary tool for time-delays measurements
- utilizes generalized Gaussian processes to model the observed light curves used for extraction of time-delays (e.g. Kovačević et al. 2015, 2018)

GP modeled light curves of a changing look AGN: NGC 3516



e.g. Shapovalova et al. 2016, 2017, 2019
Ilic et al. 2020

In-kind contributions of SER-SAG

- Software for analysis of variability of celestial sources
 - Feedback: medium
 - Non-directable software contribution
- Optical follow-up of bright LSST transients
 - Feedback: medium
 - Join via AEON
- Access to data on long-term variability of AGN
 - Feedback: low
 - Discouraged from including

In-kind telescope time

- Astronomical Station Vidojevica, Southern Serbia
 - Average seeing $\sim 1.5''$
- 1.4m telescope
- possibility of fast response (telescope moving speed is 4-6 degree/sec)
- Photometer:
 - Andor iKon-L, pixel scale 0.244 arcsec/pixel, Field of view 8.3x8.3 arcmin
 - Filters: BVRI broad bands (+L very broad filter) + Halpha, SII, red continuum narrow bands
- 6-year long experience in the Gaia-FUN-TO
- In process of joining AEON



Andjelka Kovačević

METHODOLOGICAL STRATEGY TO CATCH VARIABLE SIGNAL IN LSST LIGHT CURVES

the observing strategy will affect LSST efficiency to detect CB-SMBBH candidates:
 -the “rolling cadence” will prohibit LSST to build the long base-lines that are necessary for the search of CB-SMBBH
 -a nominal cadence of 3 days will produce, due to the filters successively alternation, the light curves in each band with just a dozen points every year

➔ **MUST INCLUDE NOVEL MACHINE LEARNING BASED PERIODICITY DETECTION METHODS INTO LSST PERIODICITY SEARCH FLOW:**

2DHYBRID METHOD BASED ON GAUSSIAN PROCESS MODELING

Successfull on decadal AGN LC with cadences of > 80days!

HYBRID METHOD: 3 STEPS

0 STEP: raw data preprocessing
 OUGP=Ornstein-Uhlenbeck Gaussian Process
 OUGP(raw LC1)=OUGP1
 OUGP(raw LC2)=OUGP2

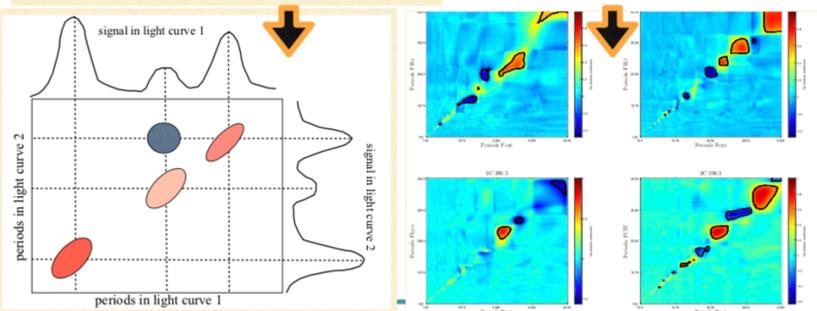
1 STEP

$$CWT(a, b) = \frac{1}{\sqrt{2\pi a}} \int_{-\infty}^{\infty} x(t) e^{i2\pi f t} e^{-\frac{(t-a)^2}{2a}} dt$$

$$env(a, b) = \sqrt{\text{Re}[\{CWT(a, b)\}^2] + \text{Im}[\{CWT(a, b)\}^2]}$$

2 STEP
 SpearmanCorrCoeff(env(OUGP1), env(OUGP2))

3C 390.3 observe: 1994-2014 Kovačević et al. 18,19,20 Gaussian processes: Kovačević et al. 17



2DHYBRID METHOD IN POSSIBLE LSST PERIODICITY SEARCH FLOW

