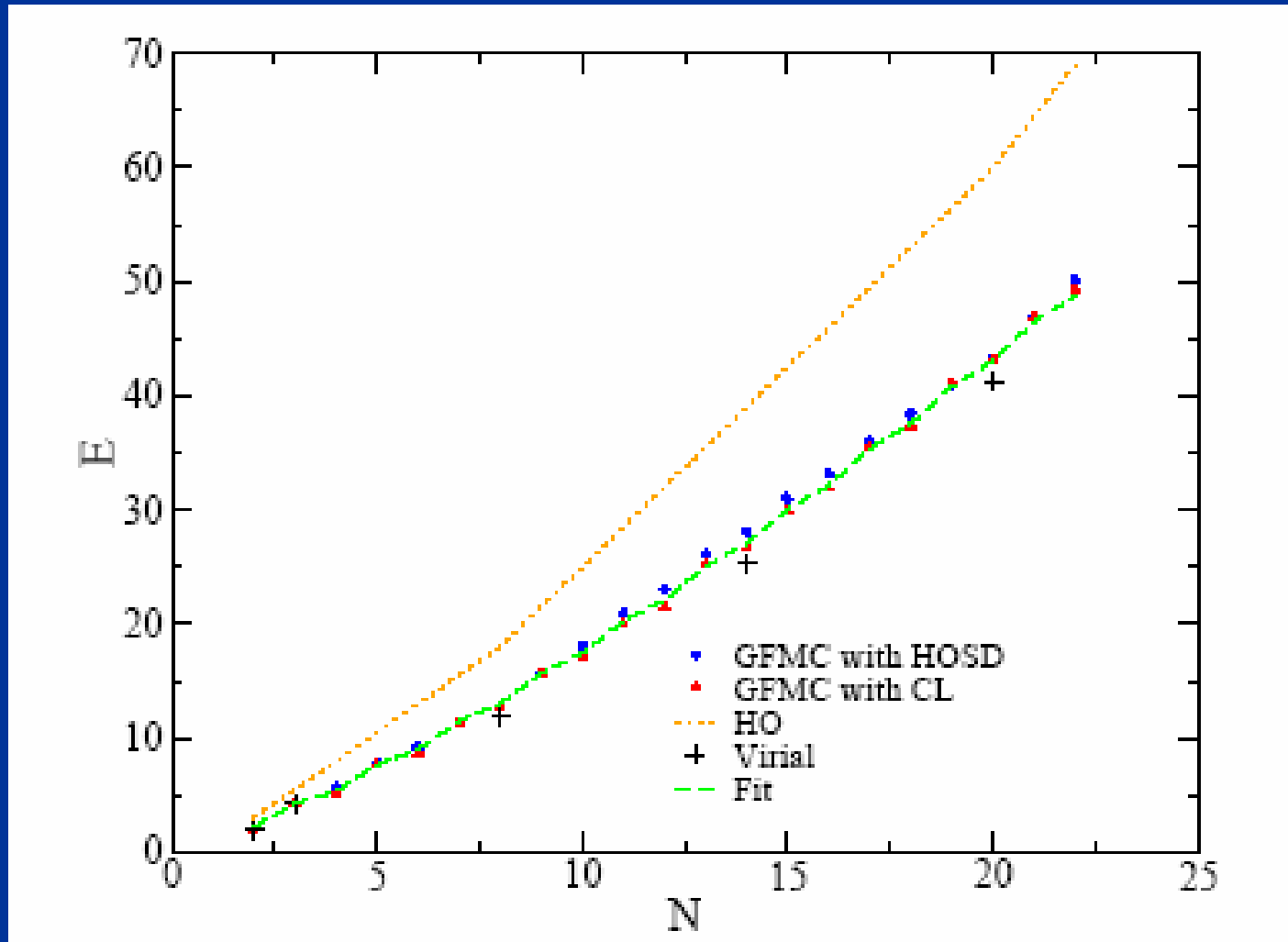


Local Density Functional Theory for Superfluid Fermionic Systems

The Unitary Fermi Gas

Unitary Fermi gas in a harmonic trap

Chang and Bertsch, physics/0703190



Outline:

- **What is a unitary Fermi gas**
- **Very brief/skewed summary of DFT**
- **Bogoliubov-de Gennes equations, renormalization**
- **Superfluid Local Density Approximation (SLDA) for a unitary Fermi gas**
- **Fermions at unitarity in a harmonic trap**

What is a unitary Fermi gas

Bertsch Many-Body X challenge, Seattle, 1999

What are the ground state properties of the many-body system composed of spin $\frac{1}{2}$ fermions interacting via a zero-range, infinite scattering-length contact interaction.

In 1999 it was not yet clear, either theoretically or experimentally, whether such fermion matter is stable or not.

- *systems of bosons are unstable (Efimov effect)*
- *systems of three or more fermion species are unstable (Efimov effect)*
- Baker (winner of the MBX challenge) concluded that the system is stable. See also Heiselberg (entry to the same competition)
- Chang et al (2003) Fixed-Node Green Function Monte Carlo and Astrakharchik et al. (2004) FN-DMC provided best the theoretical estimates for the ground state energy of such systems.
- Thomas' Duke group (2002) demonstrated experimentally that such systems are (meta)stable.

Consider Bertsch's MBX challenge (1999): "Find the ground state of infinite homogeneous neutron matter interacting with an infinite scattering length."

$$r_0 \rightarrow 0 \ll \lambda_F \ll |a| \rightarrow \infty$$

- Carlson, Morales, Pandharipande and Ravenhall, PRC 68, 025802 (2003), with Green Function Monte Carlo (GFMC)

$$\frac{E_N}{N} = \alpha_N \frac{3}{5} \varepsilon_F, \quad \alpha_N = 0.54$$

normal state

- Carlson, Chang, Pandharipande and Schmidt, PRL 91, 050401 (2003), with GFMC

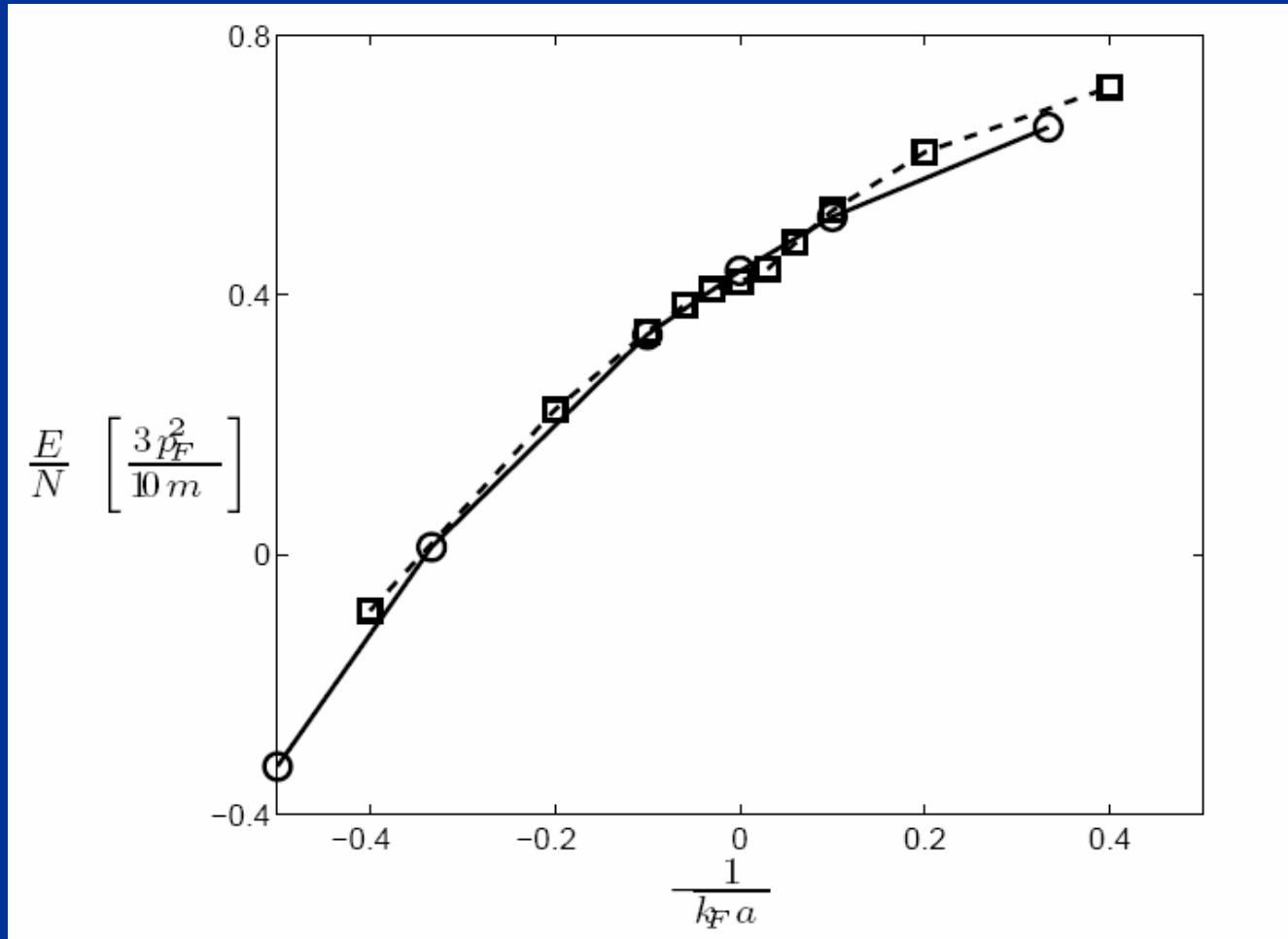
$$\frac{E_S}{N} = \alpha_S \frac{3}{5} \varepsilon_F, \quad \alpha_S = 0.44$$

superfluid state

This state is half the way from BCS→BEC crossover, the pairing correlations are in the strong coupling limit and HFB invalid again.

BEC side

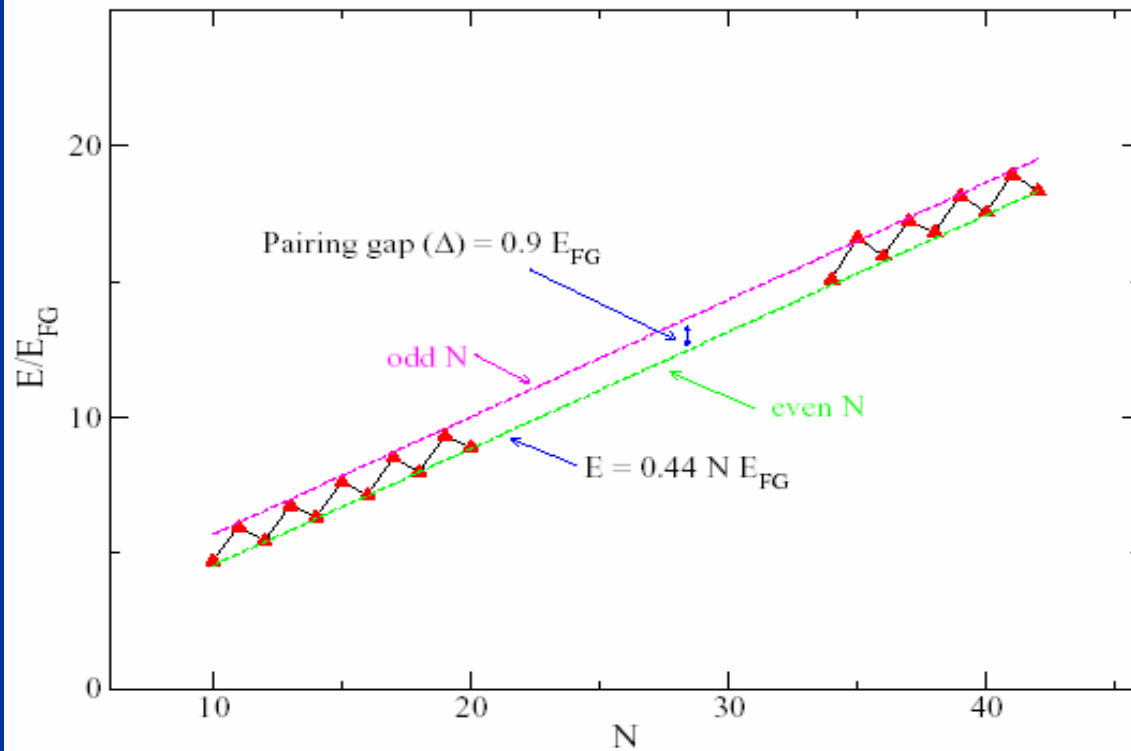
BCS side



Solid line with open circles – Chang *et al.* physics/0404115

Dashed line with squares - Astrakharchik *et al.* cond-mat/0406113

$$\Delta(2n+1) = E(2n+1) - \frac{1}{2}(E(2n) + E(2n+2))$$

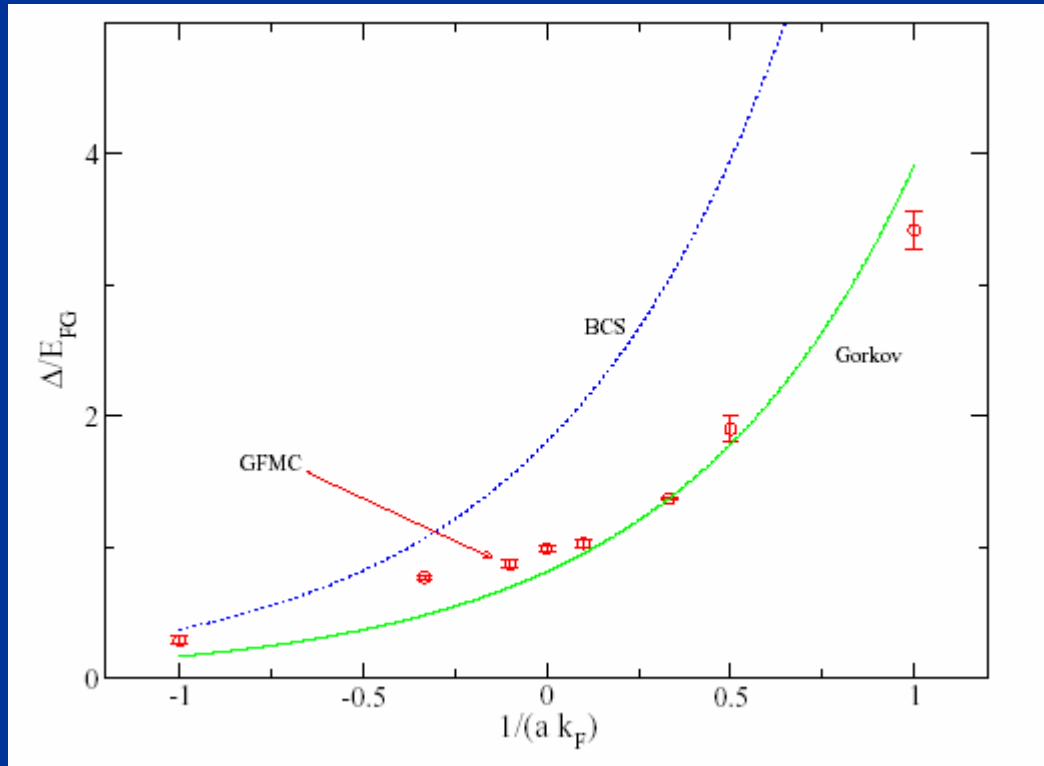


Result for $ak_F = -\infty$

$$E_{FG} = \frac{3 \hbar^2 k_F^2}{5 \cdot 2m}$$

Green Function Monte Carlo with Fixed Nodes

S.-Y. Chang, J. Carlson, V. Pandharipande and K. Schmidt
 physics/0403041



$$\Delta_{Gorkov} = \left(\frac{2}{e}\right)^{7/3} \frac{\hbar^2 k_F^2}{2m} \exp\left(\frac{\pi}{2k_F a}\right)$$

$$\Delta_{BCS} = \frac{8}{e^2} \frac{\hbar^2 k_F^2}{2m} \exp\left(\frac{\pi}{2k_F a}\right)$$

Fixed node GFMC results, S.-Y. Chang *et al.* (2003)

BCS \rightarrow BEC crossover

Leggett (1980), Nozieres and Schmitt-Rink (1985), Randeria *et al.* (1993),...

If $a < 0$ at $T=0$ a Fermi system is a BCS superfluid

$$\Delta \approx \left(\frac{2}{e}\right)^{7/3} \frac{\hbar^2 k_F^2}{2m} \exp\left(\frac{\pi}{2k_F a}\right) \ll \varepsilon_F, \quad \text{iff } k_F |a| \ll 1 \text{ and } \xi = \frac{1}{k_F} \frac{\varepsilon_F}{\Delta} \gg \frac{1}{k_F}$$

If $|a| = \infty$ and $nr_0^3 \ll 1$ a Fermi system is strongly coupled and its properties are universal. Carlson *et al.* PRL 91, 050401 (2003)

$$\frac{E_{\text{normal}}}{N} \approx 0.54 \frac{3}{5} \varepsilon_F, \quad \frac{E_{\text{superfluid}}}{N} \approx 0.44 \frac{3}{5} \varepsilon_F \quad \text{and } \xi = O(\lambda_F), \quad \Delta = O(\varepsilon_F)$$

If $a > 0$ ($a \gg r_0$) and $na^3 \ll 1$ the system is a dilute BEC of tightly bound dimers

$$\varepsilon_2 = -\frac{\hbar^2}{ma^2} \quad \text{and} \quad n_b a^3 \ll 1, \quad \text{where} \quad n_b = \frac{n_f}{2} \quad \text{and} \quad a_{bb} = 0.6a > 0$$

Very brief/skewed summary of DFT

Density Functional Theory (DFT)

Hohenberg and Kohn, 1964

$$E_{gs} = E[\rho(\vec{r})]$$

particle density only!

Local Density Approximation (LDA)

Kohn and Sham, 1965

The energy density is typically determined in *ab initio* calculations of infinite homogeneous matter.

$$E_{gs} = \int d^3r \left\{ \frac{\hbar^2}{2m} \tau(\vec{r}) + \varepsilon[\rho(\vec{r})]\rho(\vec{r}) \right\}$$

$$\rho(\vec{r}) = \sum_{i=1}^N |\psi_i(\vec{r})|^2 \quad \tau(\vec{r}) = \sum_{i=1}^N |\vec{\nabla} \psi_i(\vec{r})|^2$$

$$-\frac{\hbar^2 \Delta}{2m} \psi_i(\vec{r}) + U(\vec{r})\psi_i(\vec{r}) = \varepsilon_i \psi_i(\vec{r})$$

Kohn-Sham equations

One can construct however an EDF which depends both on particle density and kinetic energy density and use it in an extended Kohn-Sham approach (perturbative result)

$$\begin{aligned} E[\rho(\mathbf{x}), \tau(\mathbf{x})] = & \int d^3\mathbf{x} \left\{ \frac{1}{2M} \tau(\mathbf{x}) + v(\mathbf{x}) \rho(\mathbf{x}) + \frac{1}{2} \frac{(\nu - 1)}{\nu} \frac{4\pi a_s}{M} [\rho(\mathbf{x})]^2 \right. \\ & + (B_2 a_s^2 r_s + B_3 a_p^3) \frac{1}{2M} \rho(\mathbf{x}) \tau(\mathbf{x}) + (3B_2 a_s^2 r_s - B_3 a_p^3) \frac{1}{8M} [\nabla \rho(\mathbf{x})]^2 \\ & \left. + b_1 \frac{a_s^2}{2M} [\rho(\mathbf{x})]^{7/3} + b_4 \frac{a_s^3}{2M} [\rho(\mathbf{x})]^{8/3} \right\} . \end{aligned}$$

Notice that dependence on kinetic energy density and on the gradient of the particle density emerges because of finite range effects.

Bhattacharyya and Furnstahl, nucl-phys/0408014

Extended Kohn-Sham equations

Position dependent mass

$$E_{gs} = \int d^3r \left\{ \frac{\hbar^2}{2m^*[\rho(\vec{r})]} \tau(\vec{r}) + \varepsilon[\rho(\vec{r})]\rho(\vec{r}) \right\}$$

$$\rho(\vec{r}) = \sum_{i=1}^N |\psi_i(\vec{r})|^2 \quad \tau(\vec{r}) = \sum_{i=1}^N |\vec{\nabla} \psi_i(\vec{r})|^2$$

$$-\vec{\nabla} \frac{\hbar^2}{2m^*[\rho(\vec{r})]} \vec{\nabla} \psi_i(\vec{r}) + U(\vec{r})\psi_i(\vec{r}) = \varepsilon_i \psi_i(\vec{r})$$

Normal Fermi systems only!

However, not everyone is normal!

Superconductivity and superfluidity in Fermi systems

- Dilute atomic Fermi gases $T_c \approx 10^{-12} - 10^{-9} \text{ eV}$
- Liquid ^3He $T_c \approx 10^{-7} \text{ eV}$
- Metals, composite materials $T_c \approx 10^{-3} - 10^{-2} \text{ eV}$
- Nuclei, neutron stars $T_c \approx 10^5 - 10^6 \text{ eV}$
- QCD color superconductivity $T_c \approx 10^7 - 10^8 \text{ eV}$

units (1 eV \approx 10⁴ K)

Bogoliubov-de Gennes equations and renormalization

SLDA - Extension of Kohn-Sham approach to superfluid Fermi systems

$$E_{gs} = \int d^3r \varepsilon(\rho(\vec{r}), \tau(\vec{r}), \nu(\vec{r}))$$

$$\rho(\vec{r}) = 2 \sum_k |\mathbf{v}_k(\vec{r})|^2, \quad \tau(\vec{r}) = 2 \sum_k |\vec{\nabla} \mathbf{v}_k(\vec{r})|^2$$

$$\nu(\vec{r}) = \sum_k \mathbf{u}_k(\vec{r}) \mathbf{v}_k^*(\vec{r})$$

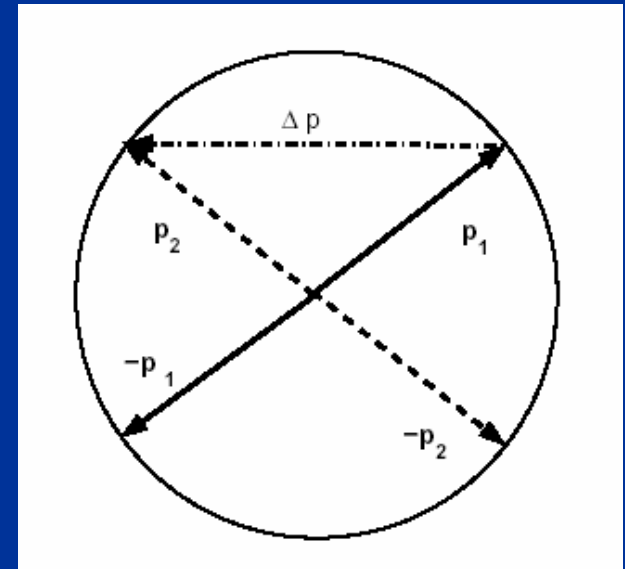
$$\begin{pmatrix} T + U(\vec{r}) - \mu & \Delta(\vec{r}) \\ \Delta^*(\vec{r}) & -(T + U(\vec{r}) - \mu) \end{pmatrix} \begin{pmatrix} \mathbf{u}_k(\vec{r}) \\ \mathbf{v}_k(\vec{r}) \end{pmatrix} = E_k \begin{pmatrix} \mathbf{u}_k(\vec{r}) \\ \mathbf{v}_k(\vec{r}) \end{pmatrix}$$

Mean-field and pairing field are both local fields!
(for sake of simplicity spin degrees of freedom are not shown)

There is a little problem! The pairing field Δ diverges.

Why would one consider a local pairing field?

- ✓ Because it makes sense physically!
- ✓ The treatment is so much simpler!
- ✓ Our intuition is so much better also.



$$r_0 \cong \frac{\hbar}{p_F} = k_F^{-1}$$

radius of interaction inter-particle separation

$$\Delta = \omega_D \text{Exp} \left(-\frac{1}{|V|N} \right) \ll \varepsilon_F$$

$$\xi \approx \frac{1}{k_F} \frac{\varepsilon_F}{\Delta} \gg r_0$$

coherence length
size of the Cooper pair

Nature of the problem

$$\nu(\vec{r}_1, \vec{r}_2) = \sum_{E_k > 0} v_k^*(\vec{r}_1) u_k(\vec{r}_2) \propto \frac{1}{|\vec{r}_1 - \vec{r}_2|} \quad \leftarrow \text{at small separations}$$

$$\Delta(\vec{r}_1, \vec{r}_2) = -V(\vec{r}_1, \vec{r}_2) \nu(\vec{r}_1, \vec{r}_2)$$

It is easier to show how this singularity appears in infinite homogeneous matter.

$$v_k(\vec{r}_1) = v_k \exp(i\vec{k} \cdot \vec{r}_1), \quad u_k(\vec{r}_2) = u_k \exp(i\vec{k} \cdot \vec{r}_2)$$

$$v_k^2 = \frac{1}{2} \left(1 - \frac{\varepsilon_k - \mu}{\sqrt{(\varepsilon_k - \mu)^2 + \Delta^2}} \right), \quad u_k^2 + v_k^2 = 1, \quad \varepsilon_k = \frac{\hbar^2 \vec{k}^2}{2m} + U, \quad \Delta = \frac{\hbar^2 \delta}{2m}$$

$$\nu(r) = \frac{\Delta m}{2\pi^2 \hbar^2} \int_0^\infty dk \frac{\sin(kr)}{kr} \frac{k^2}{\sqrt{(k^2 - k_F^2)^2 + \delta^2}}, \quad r = |\vec{r}_1 - \vec{r}_2|$$

Pseudo-potential approach

(appropriate for very slow particles, very transparent, but somewhat difficult to improve)

Lenz (1927), Fermi (1931), Blatt and Weiskopf (1952)

Lee, Huang and Yang (1957)

$$-\frac{\hbar^2 \Delta_{\vec{r}}}{m} \psi(\vec{r}) + V(\vec{r})\psi(\vec{r}) = E \psi(\vec{r}), \quad V(\vec{r}) \approx 0 \text{ if } r > R$$

$$\psi(\vec{r}) = \exp(i\vec{k} \cdot \vec{r}) + \frac{f}{r} \exp(ikr) \approx 1 + \frac{f}{r} + \dots \approx 1 - \frac{a}{r} + O(kr)$$

$$f^{-1} = -\frac{1}{a} + \frac{1}{2} r_0 k^2 - ik, \quad g = \frac{4\pi \hbar^2 a}{m(1 + ika)} + \dots$$

$$\text{if } kr_0 \ll 1 \text{ then } V(\vec{r})\psi(\vec{r}) \Rightarrow g \delta(\vec{r}) \frac{\partial}{\partial r} [r \psi(\vec{r})]$$

$$\text{Example : } \psi(\vec{r}) = \frac{A}{r} + B + \dots \Rightarrow \delta(\vec{r}) \frac{\partial}{\partial r} [r \psi(\vec{r})] = \delta(\vec{r}) B$$

The SLDA (renormalized) equations

$$E_{gs} = \int d^3r \left\{ \underline{\varepsilon_N [\rho(\vec{r}), \tau(\vec{r})]} + \underline{\varepsilon_S [\rho(\vec{r}), \nu(\vec{r})]} \right\}$$

$$\varepsilon_S [\rho(\vec{r}), \nu(\vec{r})] \stackrel{\text{def}}{=} -\Delta(\vec{r})\nu_c(\vec{r}) = g_{\text{eff}}(\vec{r})|\nu_c(\vec{r})|^2$$

$$\begin{cases} [h(\vec{r}) - \mu]u_i(\vec{r}) + \Delta(\vec{r})v_i(\vec{r}) = E_i u_i(\vec{r}) \\ \Delta^*(\vec{r})u_i(\vec{r}) - [h(\vec{r}) - \mu]v_i(\vec{r}) = E_i v_i(\vec{r}) \end{cases} \quad \begin{cases} h(\vec{r}) = -\vec{\nabla} \frac{\hbar^2}{2m(\vec{r})} \vec{\nabla} + U(\vec{r}) \\ \Delta(\vec{r}) = -g_{\text{eff}}(\vec{r})\nu_c(\vec{r}) \end{cases}$$

$$\frac{1}{g_{\text{eff}}(\vec{r})} = \frac{1}{g[\rho(\vec{r})]} - \frac{m(\vec{r})k_c(\vec{r})}{2\pi^2\hbar^2} \left\{ 1 - \frac{k_F(\vec{r})}{2k_c(\vec{r})} \ln \frac{k_c(\vec{r}) + k_F(\vec{r})}{k_c(\vec{r}) - k_F(\vec{r})} \right\}$$

$$\rho_c(\vec{r}) = 2 \sum_{E_i \geq 0} |v_i(\vec{r})|^2, \quad \nu_c(\vec{r}) = \sum_{E_i \geq 0} v_i^*(\vec{r})u_i(\vec{r})$$

$$E_c + \mu = \frac{\hbar^2 k_c^2(\vec{r})}{2m(\vec{r})} + U(\vec{r}), \quad \mu = \frac{\hbar^2 k_F^2(\vec{r})}{2m(\vec{r})} + U(\vec{r})$$

Position and momentum dependent running coupling constant

Observables are (obviously) independent of cut-off energy (when chosen properly).

**Superfluid Local Density Approximation (SLDA)
for a unitary Fermi gas**

The naïve SLDA energy density functional suggested by dimensional arguments

$$\varepsilon(\vec{r}) = \alpha \frac{\tau(\vec{r})}{2} + \beta \frac{3(3\pi^2)^{2/3} n^{5/3}(\vec{r})}{5} + \gamma \frac{|\mathbf{v}(\vec{r})|^2}{n^{1/3}(\vec{r})}$$

$$n(\vec{r}) = 2 \sum_k |\psi_k(\vec{r})|^2$$

$$\tau(\vec{r}) = 2 \sum_k \left| \vec{\nabla} \psi_k(\vec{r}) \right|^2$$

$$\mathbf{v}(\vec{r}) = \sum_k \mathbf{u}_k(\vec{r}) \psi_k^*(\vec{r})$$

The renormalized SLDA energy density functional

$$\varepsilon(\vec{r}) = \alpha \frac{\tau_c(\vec{r})}{2} + \beta \frac{3(3\pi^2)^{2/3} n^{5/3}(\vec{r})}{5} + g_{eff}(\vec{r}) |\nu_c(\vec{r})|^2$$

$$\tau_c(\vec{r}) = 2 \sum_{E < E_c} \left| \vec{\nabla} v_k(\vec{r}) \right|^2, \quad \nu_c(\vec{r}) = \sum_{E < E_c} u_k(\vec{r}) v_k^*(\vec{r})$$

$$\frac{1}{g_{eff}(\vec{r})} = \frac{n^{1/3}(\vec{r})}{\gamma} - \frac{k_c(\vec{r})}{2\pi^2 \alpha} \left[1 - \frac{k_0(\vec{r})}{2k_c(\vec{r})} \ln \frac{k_c(\vec{r}) + k_0(\vec{r})}{k_c(\vec{r}) - k_0(\vec{r})} \right]$$

$$E_c + \mu = \alpha \frac{k_c^2(\vec{r})}{2} + U(\vec{r}), \quad \mu = \alpha \frac{k_0^2(\vec{r})}{2} + U(\vec{r})$$

$$U(\vec{r}) = \beta \frac{(3\pi^2)^{2/3} n^{2/3}(\vec{r})}{2} - \frac{|\Delta(\vec{r})|^2}{3\gamma n^{2/3}(\vec{r})} + V_{ext}(\vec{r}) + \text{small correction}$$

$$\Delta(\vec{r}) = -g_{eff}(\vec{r}) \nu_c(\vec{r})$$

How to determine the dimensionless parameters α , β and γ ?

$$n = \frac{k_F^3}{3\pi^2} = \int \frac{d^3k}{(2\pi)^3} \left(1 - \frac{\alpha k^2 / 2 + \bar{\beta} k_F^2 / 2 - \mu}{\sqrt{(\alpha k^2 / 2 + \bar{\beta} k_F^2 / 2 - \mu)^2 + \Delta^2}} \right)$$

$$= \int \frac{d^3k}{(2\pi)^3} \left(1 - \frac{\varepsilon_k}{E_k} \right)$$

$$\frac{3}{5} \varepsilon_F n \xi_S = \frac{3}{5} \varepsilon_F n \beta + \int \frac{d^3k}{(2\pi)^3} \left[\alpha \frac{k^2}{2} \left(1 - \frac{\varepsilon_k}{E_k} \right) - \frac{\Delta^2}{2E_k} \right]$$

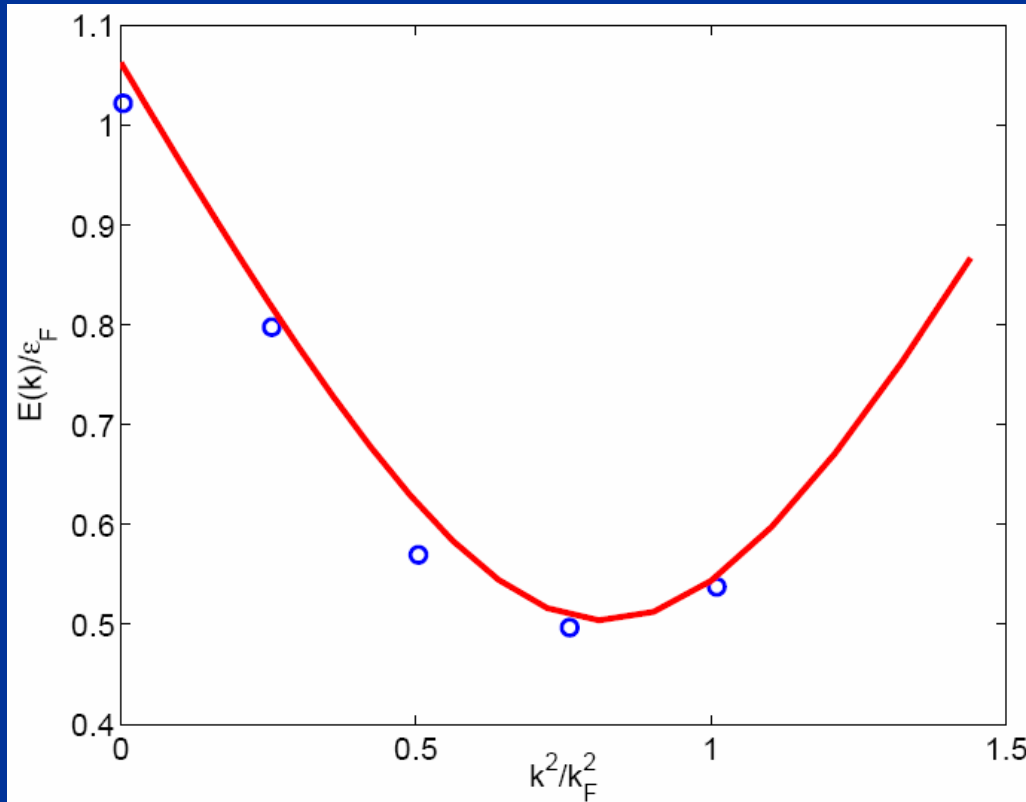
$$\frac{n^{1/3}}{\gamma} = \int \frac{d^3k}{(2\pi)^3} \left(\frac{1}{\alpha k^2} - \frac{1}{2E_k} \right)$$

One thus obtains:

$$\left\{ \begin{array}{l} \xi_s = \frac{5E}{3N \varepsilon_F} = 0.42(2) \\ \eta = \frac{\Delta}{\varepsilon_F} = 0.504(24) \\ \zeta = \frac{\mu}{\varepsilon_F} = 0.42(2) \end{array} \right. \Rightarrow \left\{ \begin{array}{l} \alpha = 1.14 \\ \beta = -0.553 \\ \frac{1}{\gamma} = -0.0906 \end{array} \right.$$

Bonus!

Quasiparticle spectrum in homogeneous matter



solid line - SLDA

circles - GFMC due to Carlson and Reddy

Extra Bonus!

The normal state has been also determined in GFMC

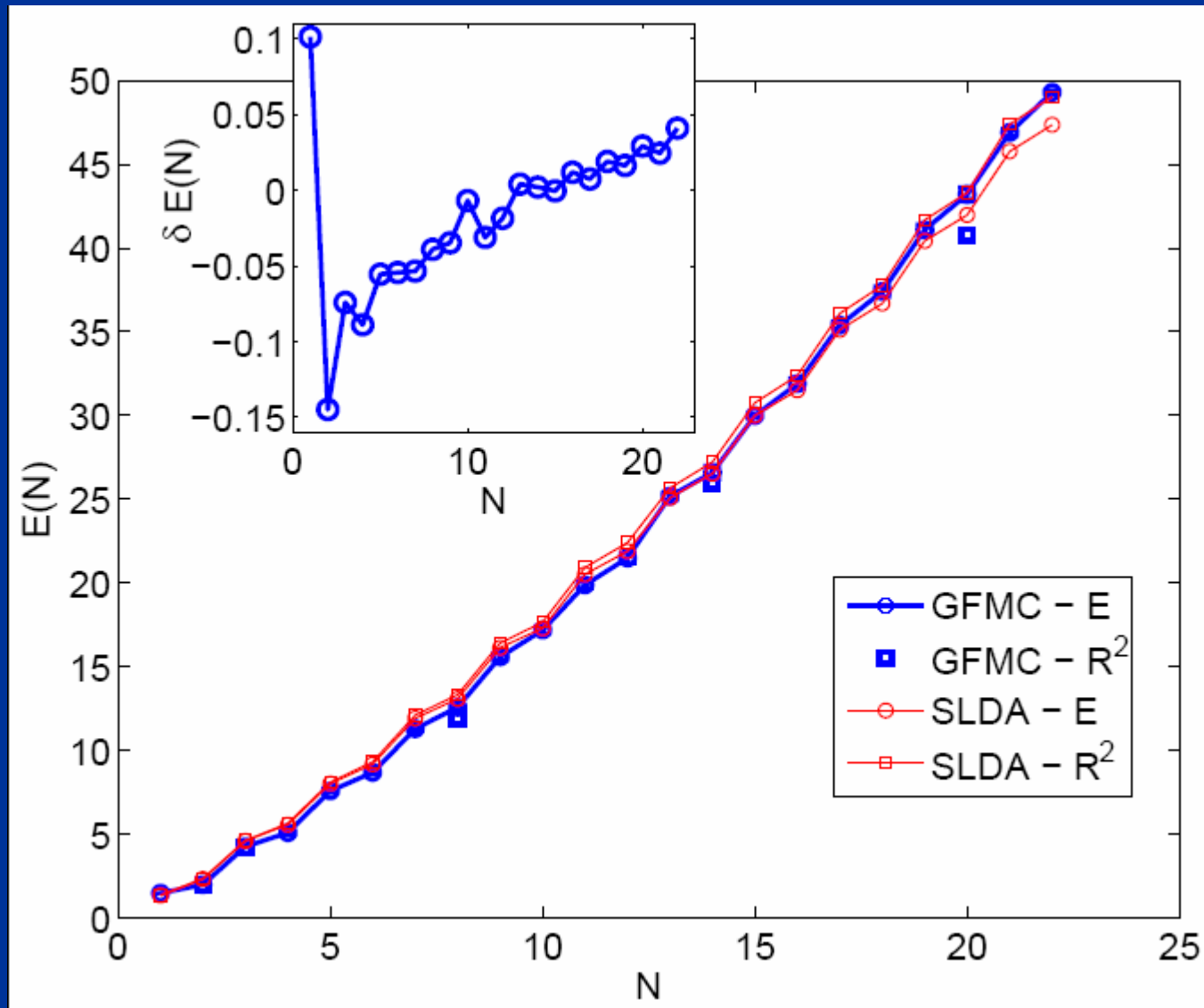
$$\xi_N = \frac{5E}{3N\varepsilon_F} = 0.55(2)$$

SLDA functional predicts

$$\xi_N = \alpha + \beta = 0.59$$

Fermions at unitarity in a harmonic trap

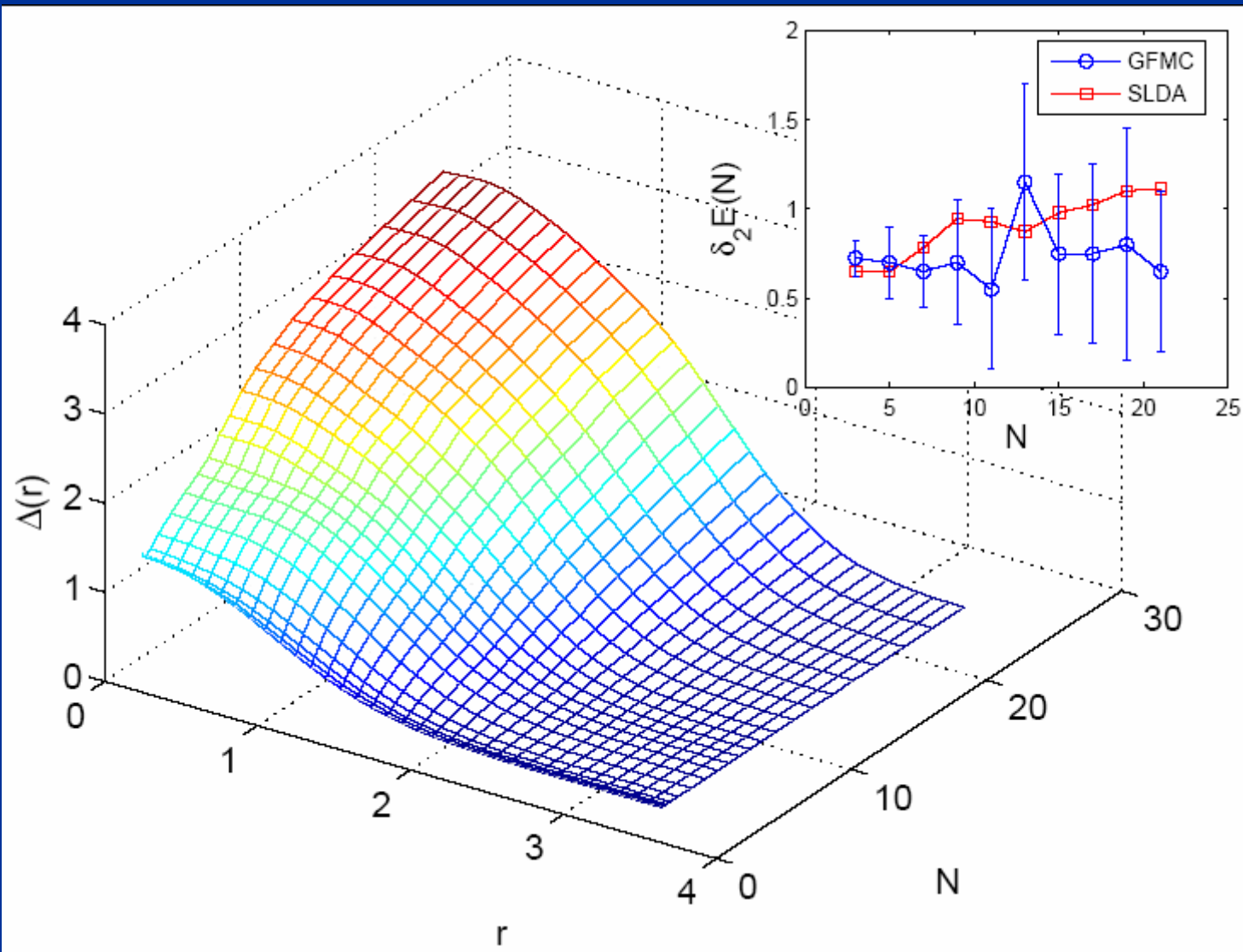
GFMC calculations of Chang and Bertsch



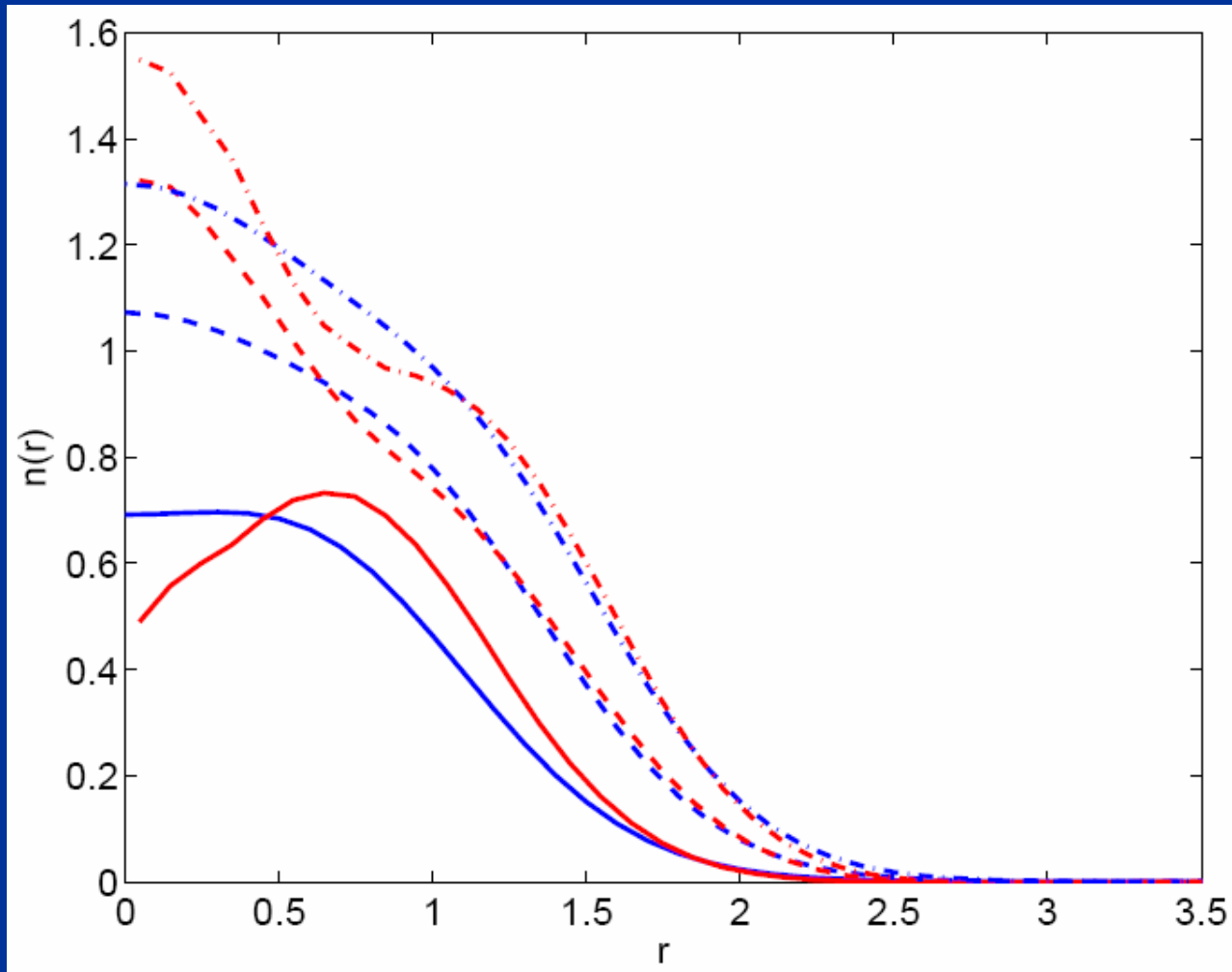
GFMC - Chang and Bertsch

TABLE I: Table 1. The energies $E(N)$ calculated within the GFMC [15] and SLDA. When two numbers are present the first was calculated as the expectation value of the Hamiltonian/functional, while the second is the value obtained using the virial theorem [20].

N	E_{GFMC}	E_{SLDA}	N	E_{GFMC}	E_{SLDA}
1	1.5	1.36	12	21.5	21.9/22.4
2	2.01/1.95	2.35/2.36	13	25.2	25.1/25.7
3	4.28/4.19	4.62/4.63	14	26.6/26.0	26.5/27.2
4	5.10	5.60/5.66	15	30.0	30.0/30.8
5	7.60	8.04/8.12	16	31.9	31.5/32.4
6	8.70	9.20/9.34	17	35.4	35.1/36.1
7	11.3	11.9/12.1	18	37.4	36.7/37.8
8	12.6/11.9	13.1/13.3	19	41.1	40.4/41.7
9	15.6	16.2/16.4	20	43.2/40.8	42.0/43.3
10	17.2	17.3/17.7	21	46.9	45.8/47.4
11	19.9	20.5/20.9	22	49.3	47.4/49.0



$$\delta_2 E(N) = E(N) - \frac{1}{2} [E(N+1) + E(N-1)]$$



**Densities for $N=8$ (solid), $N=14$ (dashed) and $N=20$ (dot-dashed)
GFMC (red), SLDA (blue)**

- **Agreement between GFMC and SLDA very good (a few percent accuracy)**

Why not better?

A better agreement would have really signaled big troubles!

- **Energy density functional is not unique, in spite of the strong restrictions imposed by unitarity**
- **Self-interaction correction neglected
smallest systems affected the most**
- **Absence of polarization effects
spherical symmetry imposed, odd systems mostly affected**
- **Spin number densities not included
extension from SLDA to SLSD(A) needed
ab initio results for asymmetric system needed**
- **Gradient corrections not included**